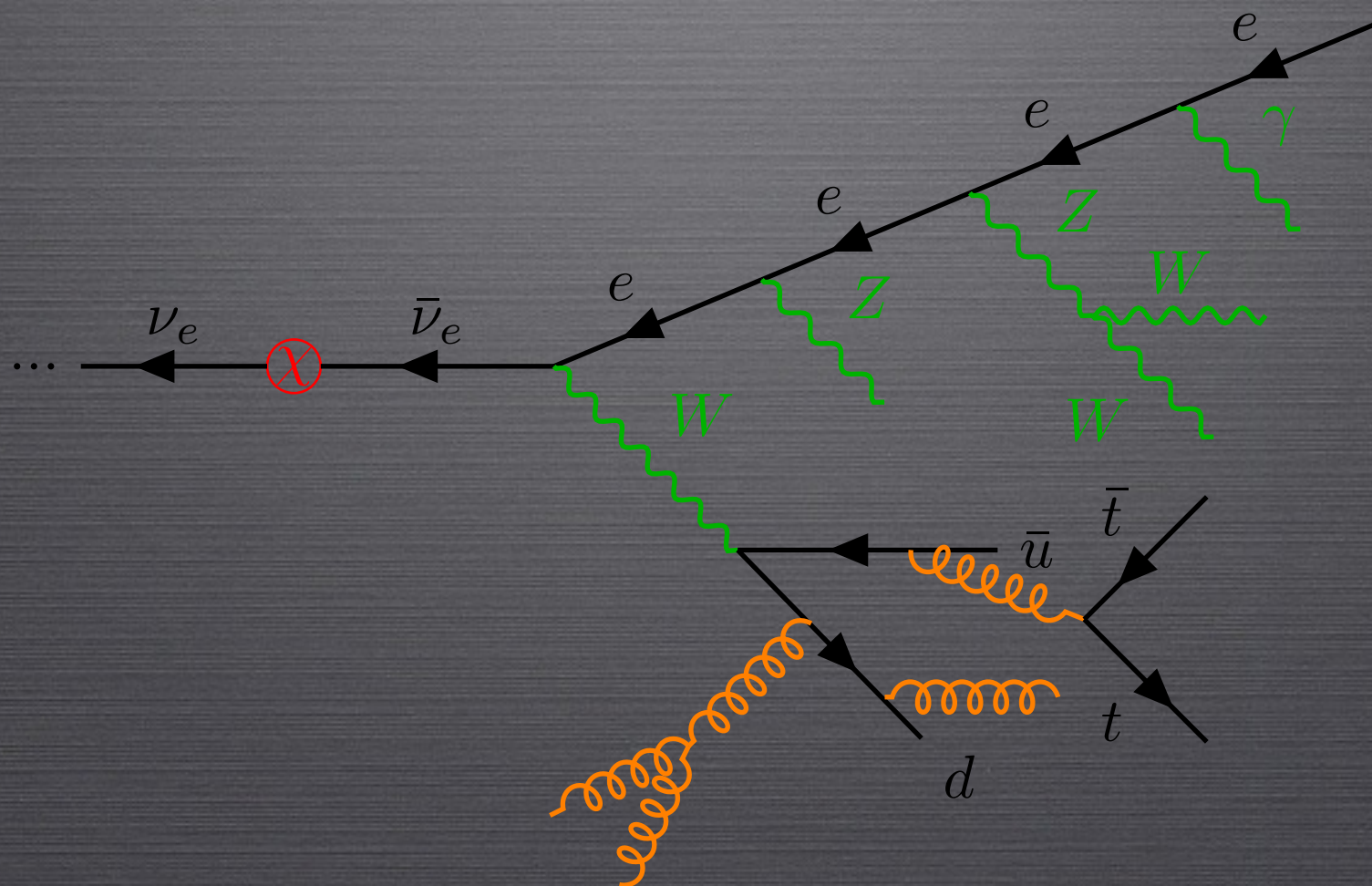




SPECTRA FOR HEAVY DARK MATTER



NICK RODD

W/ CHRISTIAN BAUER AND BRYAN WEBBER

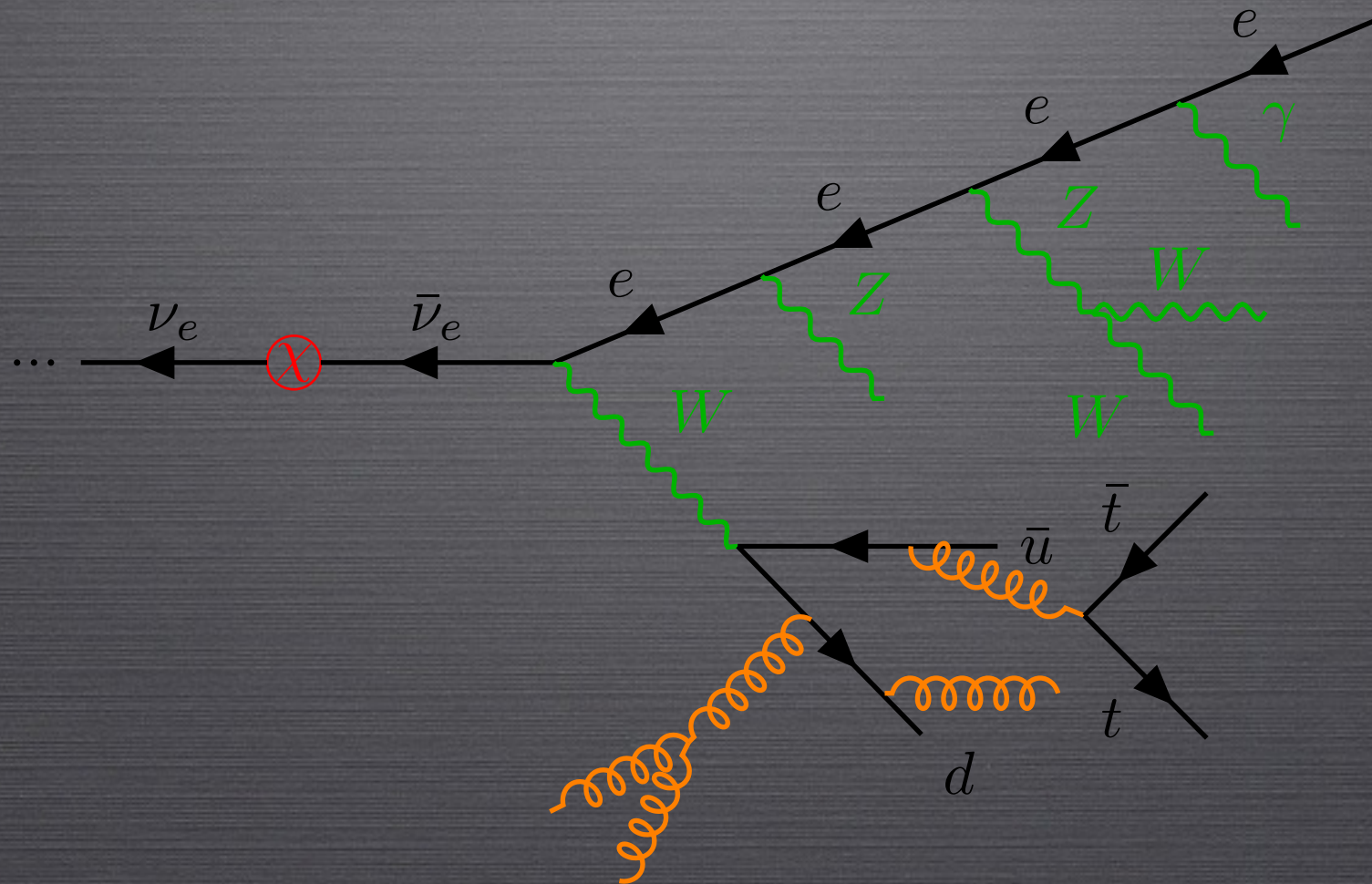
TRIUMF, 12 MARCH 2020



* $m_\chi > 10 \text{ TeV}$



SPECTRA FOR HEAVY* DARK MATTER



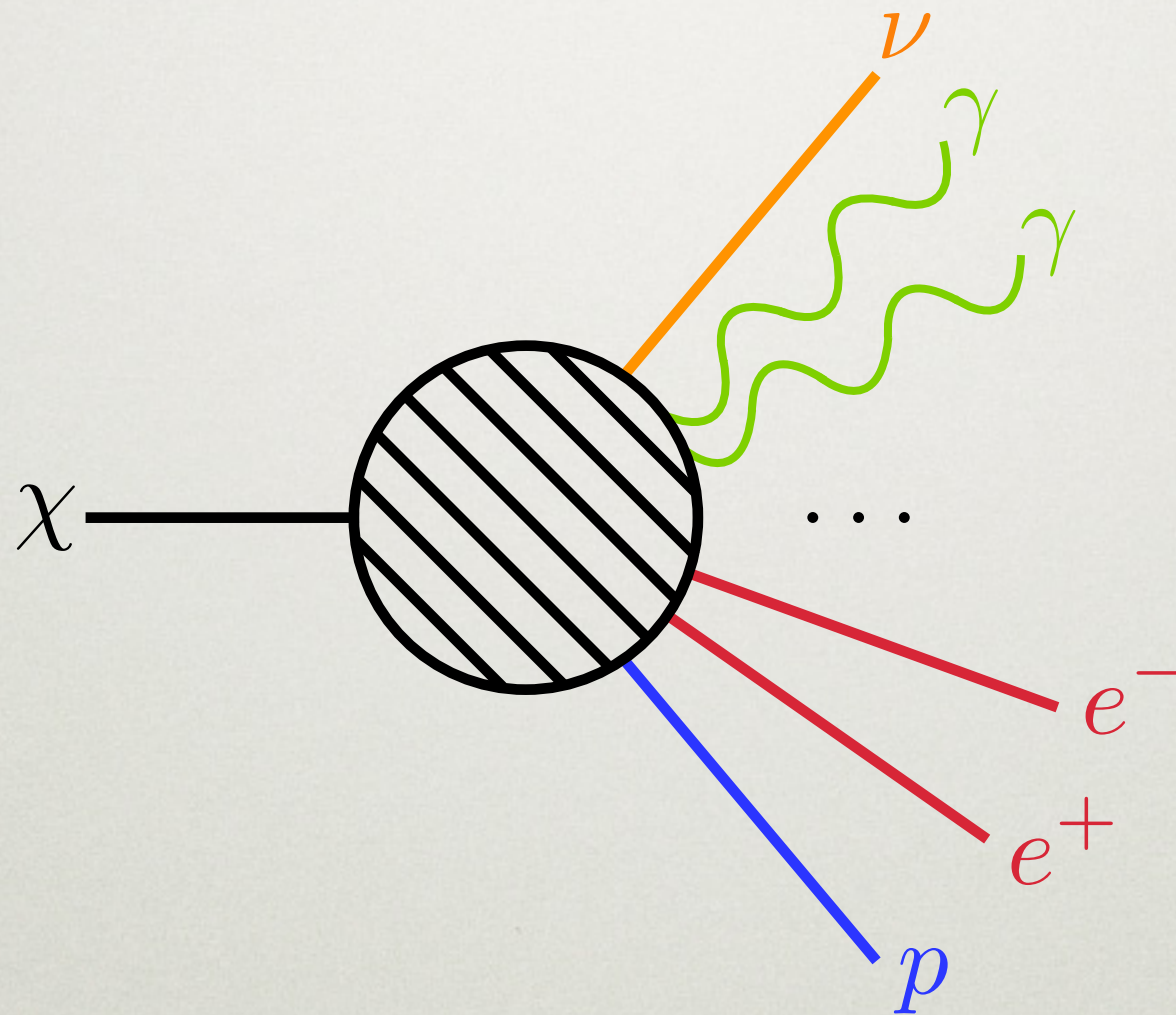
NICK RODD

W/ CHRISTIAN BAUER AND BRYAN WEBBER

TRIUMF, 13 MARCH 2020

ID ABOVE 10 TEV: OVERVIEW

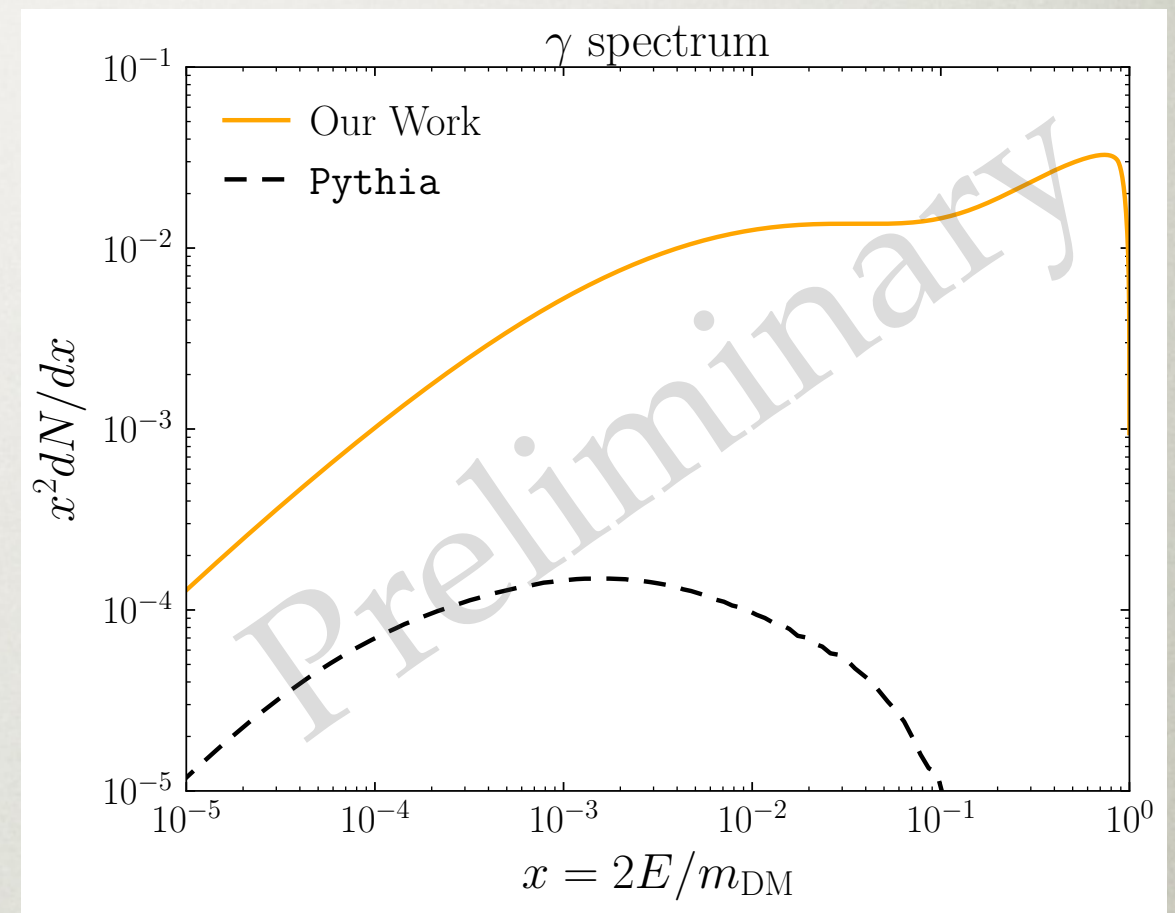
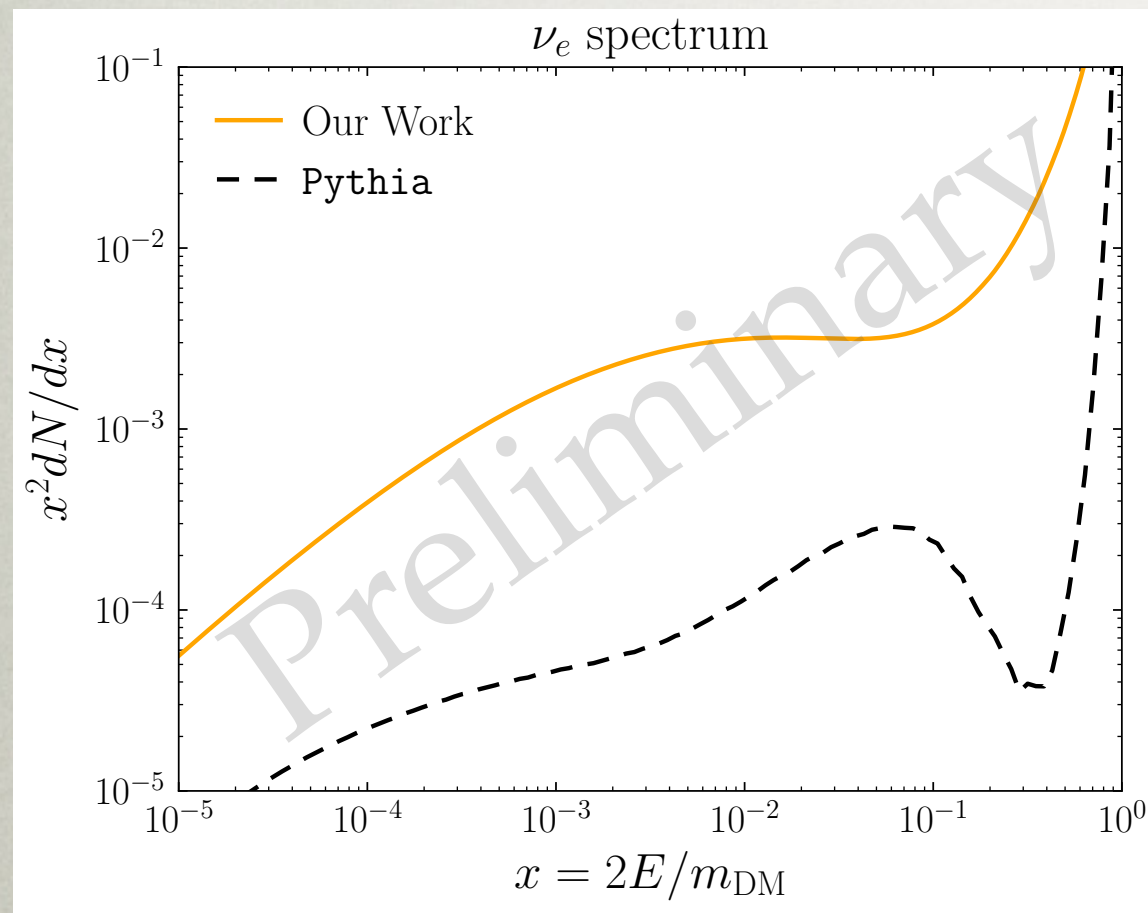
- Imagine EeV DM that can decay to neutrinos
- What is the **prompt*** spectrum of stable particles? $\{\gamma, \nu, e, p\}$



*Note the spectrum at Earth can be quite different; [Cohen, Murase, NLR, Safdi, Soreq 1612.05638], [Murase, Beacom 1206.2595], [Esmaili, Serpico 1505.06486]

ID ABOVE 10 TEV: OVERVIEW

- Imagine EeV DM that can decay to neutrinos
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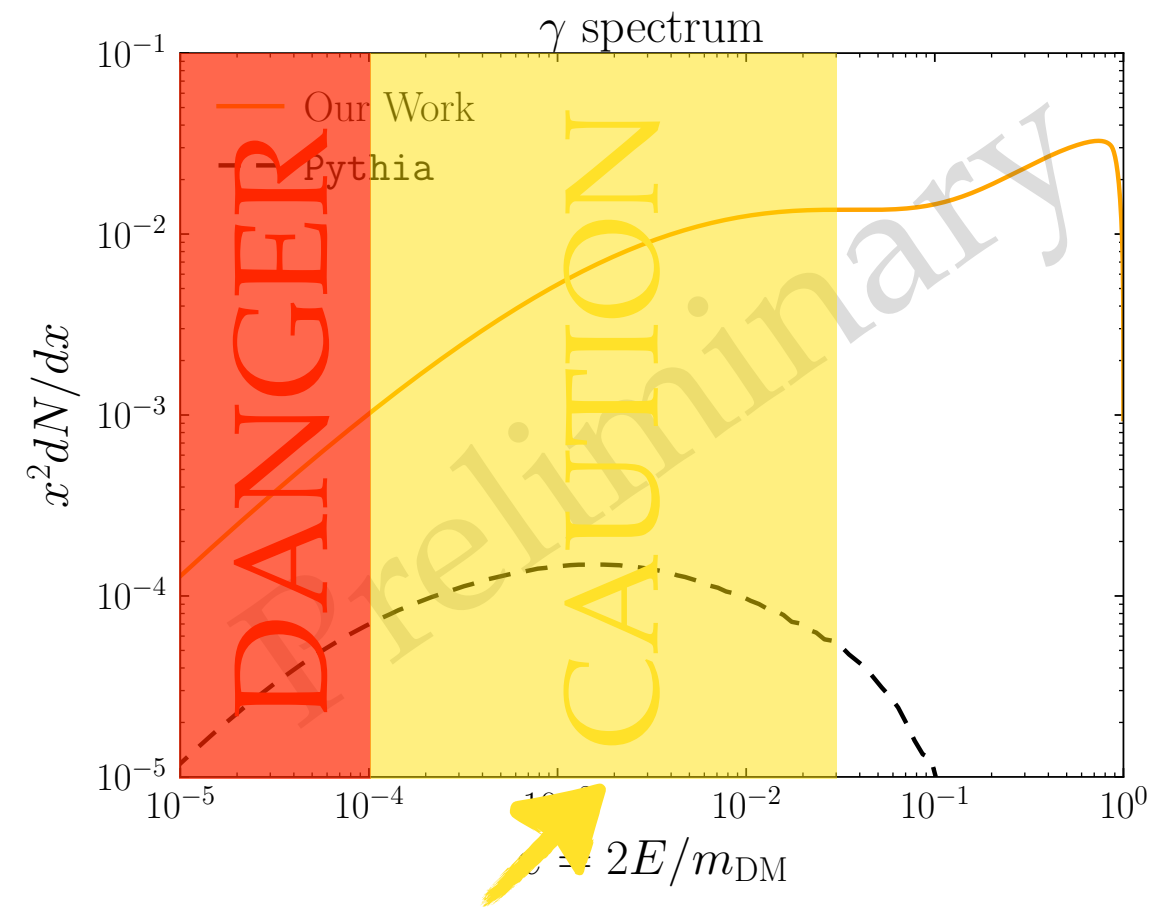
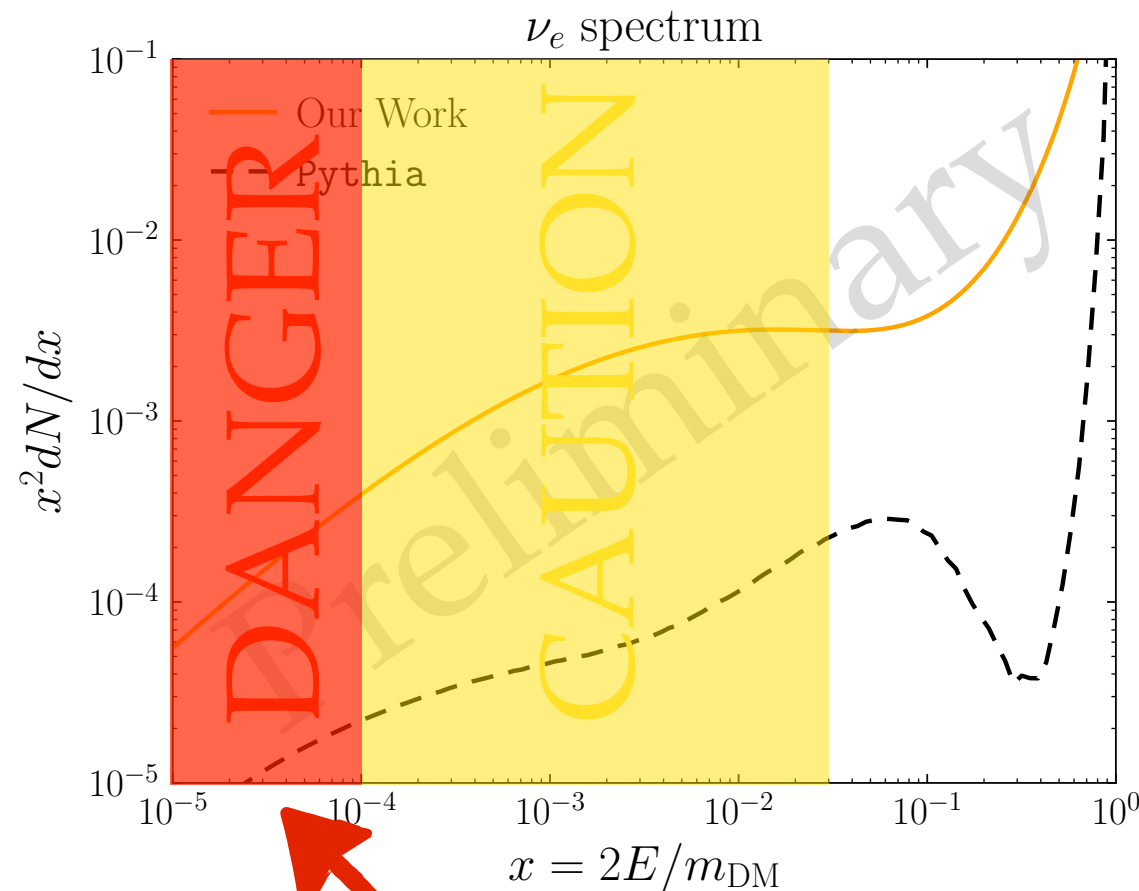
(This "SCHEDULE I" is posted on the internet at www.columbiaconventioncenter.com and is made a part of the License Agreement as described in Section "10" thereof.)



SCHEDULE I
TO ~~COLUMBIA METROPOLITAN CONVENTION CENTER~~
LICENSE AGREEMENT (the "Agreement")

our spectra

GENERAL TERMS AND CONDITIONS

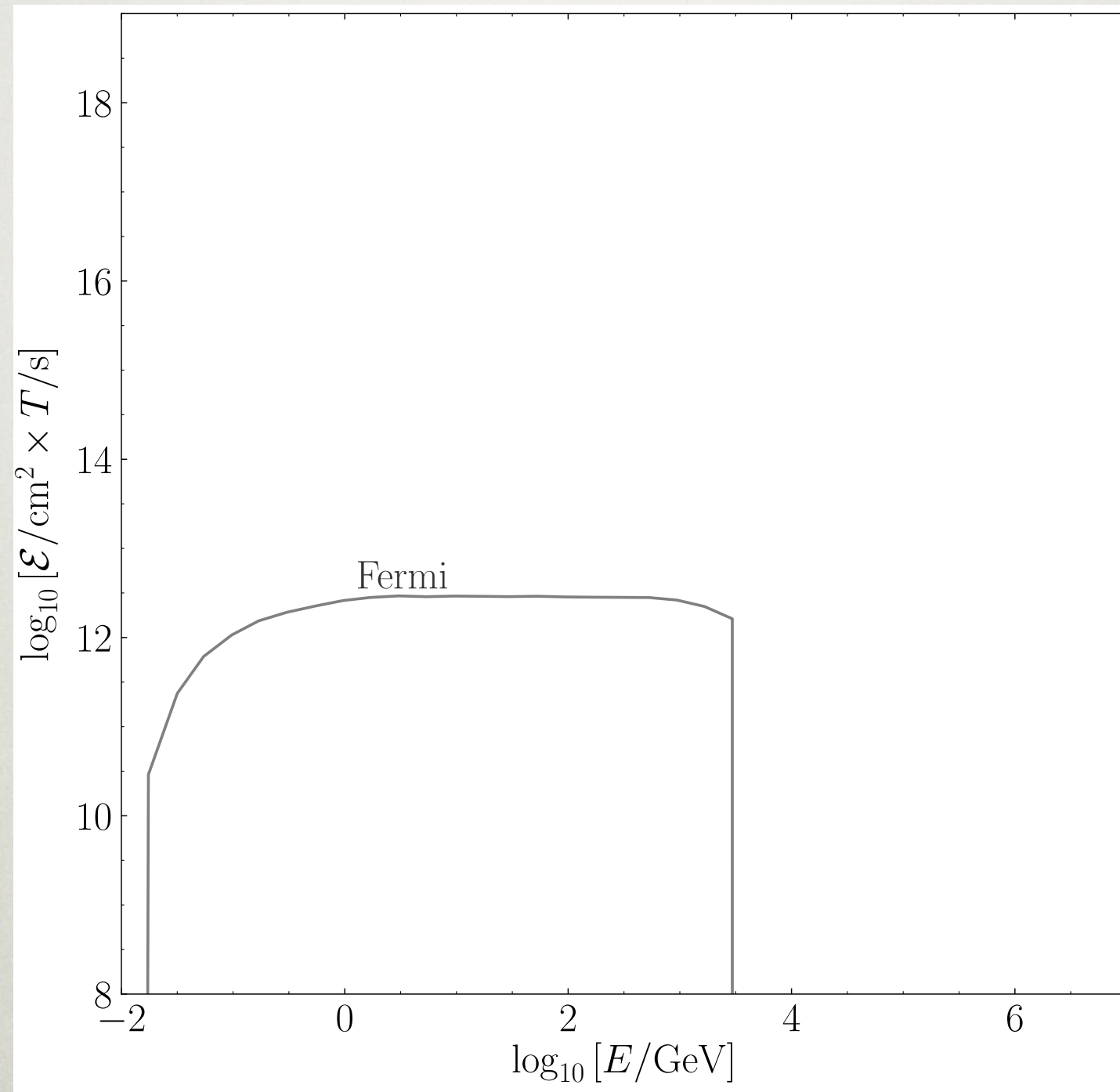


Missing QCD single logs

Incomplete treatment of color coherence

Customer responsible for the increased cost of such services to the extent the Licensed Area contains excessive trash, difficult to remove items (e.g., confetti, glitter, or flower petals) or other conditions for which the Customer is responsible.

ID ABOVE 10 TEV: EXPERIMENTS

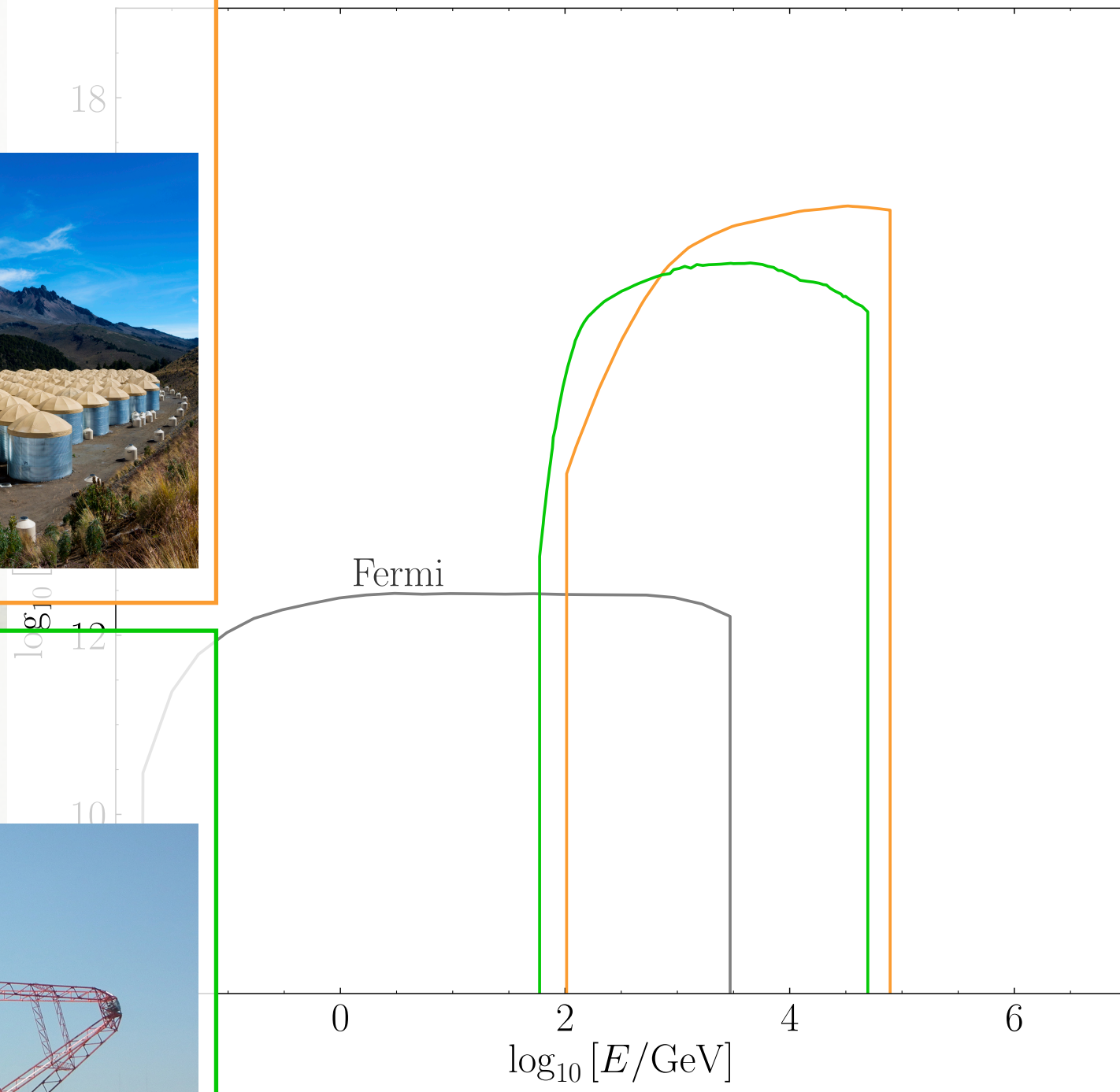


ID ABOVE 10 TEV: EXPERIMENTS

HAWC
2015-PRESENT



H.E.S.S.
2002-PRESENT
(H.E.S.S. II 2012-)



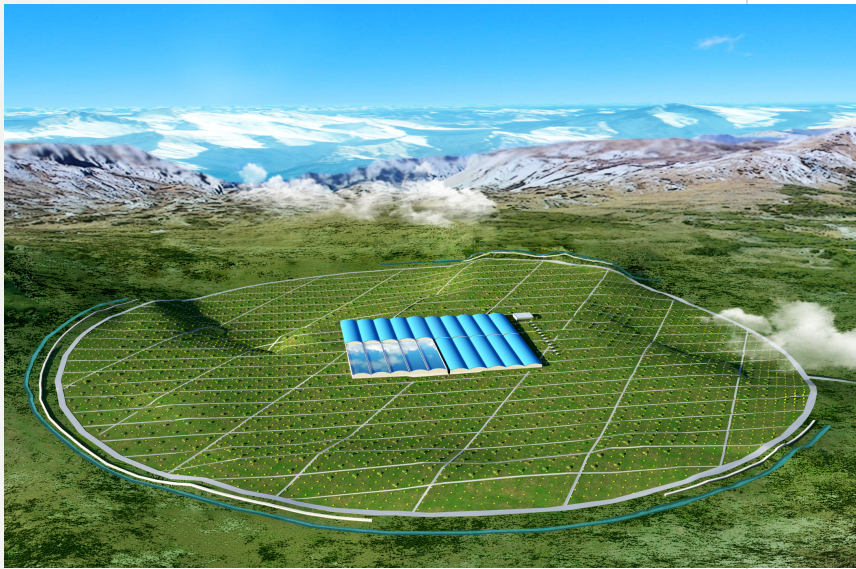
ID ABOVE 10 TEV: EXPERIMENTS

LHAASO

2020

[1905.02773]

18

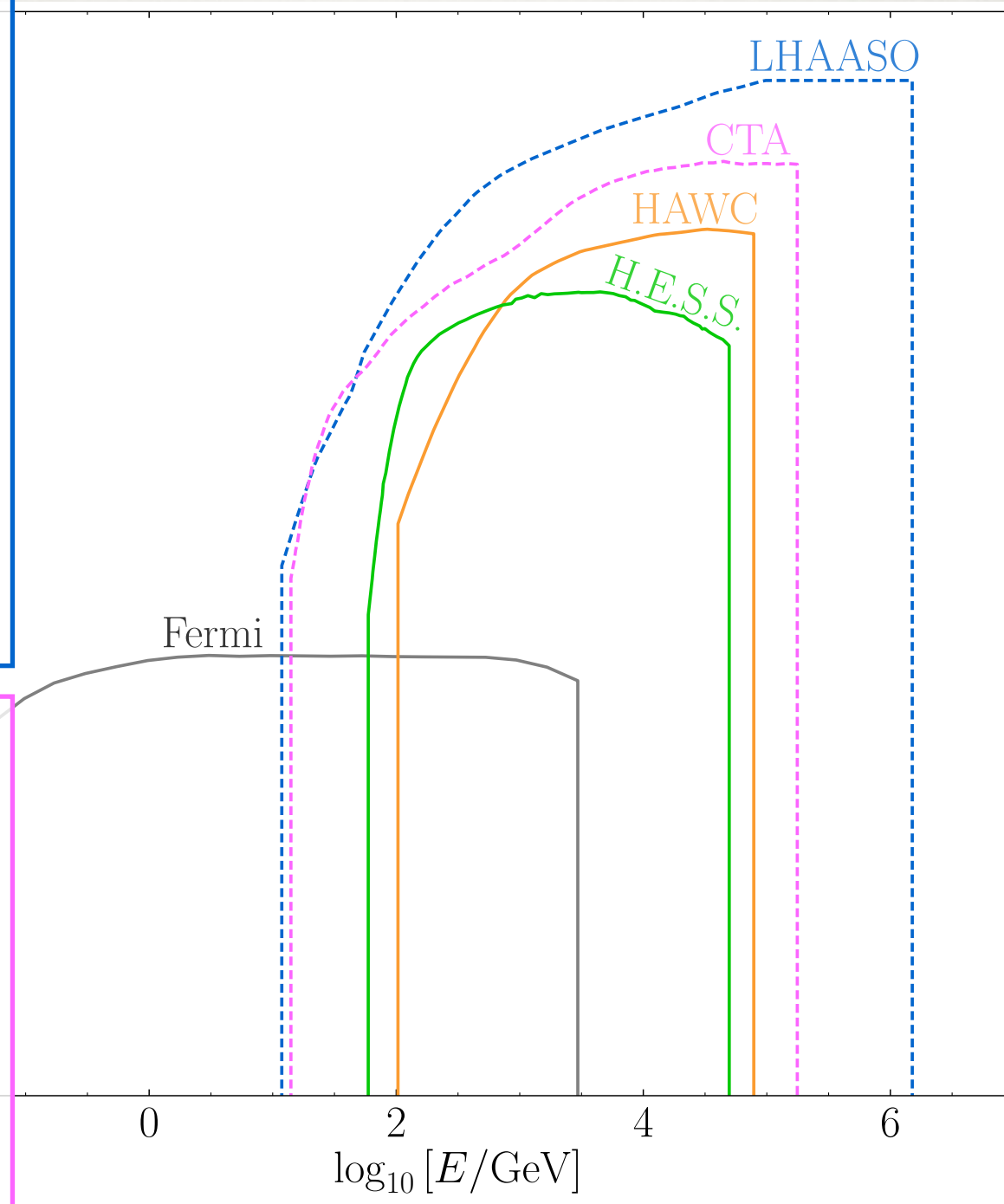
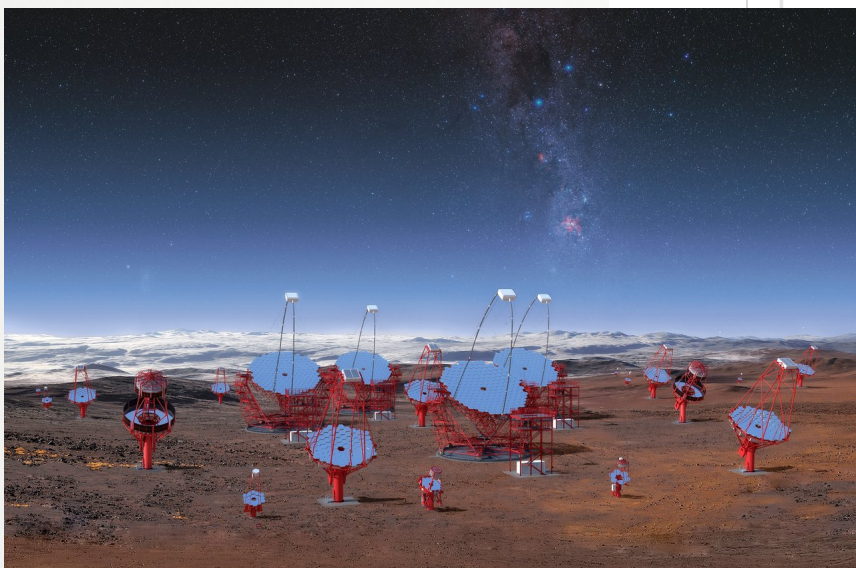


CTA

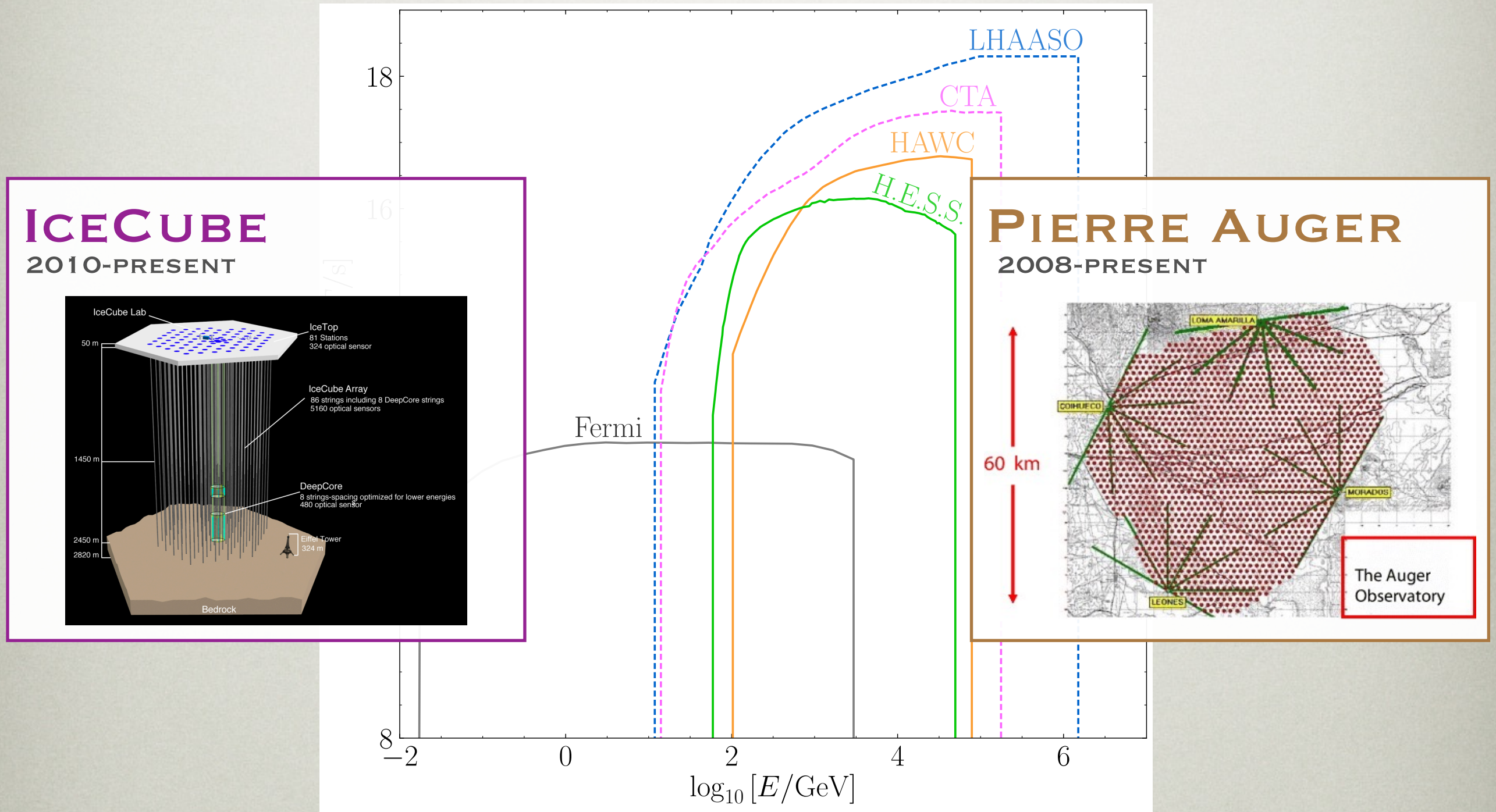
~2025

[1709.07997]

log₁₀
12



ID ABOVE 10 TEV: EXPERIMENTS



ID ABOVE 10 TEV: CUTTING EDGE



- How are the collaborations getting their spectra?
 - IceCube: **Pythia** up to EeV [1804.03848]
 - HAWC: **Pythia** up to 10 PeV [NLR+ 1710.10288]
 - HESS: **Pythia** up to 20 TeV [1012.5602]
 - LHAASO: **PPPC4DMID** up to 100 TeV [1910.05017]
 - CTA: **PPPC4DMID** up to 10 TeV [1408.4131]
 - Pierre Auger:* **DGLAP** to GUT scale, e.g. [Kalashev, Kuznetsov 1606.07354]



PPPC 4 DM ID: A Poor Particle Physicist Cookbook for Dark Matter Indirect Detection

Marco Cirelli^{a,b}, Gennaro Corcella^{c,d,e}, Andi Hektor^f,
Gert Hütsi^g, Mario Kadastik^f, Paolo Panci^{a,h,i,j},
Martti Raidal^f, Filippo Sala^{d,e}, Alessandro Strumia^{a,e,f,k}

[1012.4515]



*See also: [Barbot, Drees 2002, 2003] for a full treatment in the MSSM



ID ABOVE 10 TEV: CUTTING EDGE

- How are the collaborations getting their spectra?
 - IceCube: **Pythia** up to EeV [1804.03848]
 - HAWC: **Pythia** up to 10 PeV [NLR+ 1710.10288]

NO IDEAL OPTION

- **Pythia**: no electroweak TGC, difficult to run above \sim PeV
- **PPPC4DMID**: Stops at 100 TeV, only LO electroweak
- **DGLAP**: can evolve arbitrarily high, but just done for b



PPPC 4 DM ID:
**A Poor Particle Physicist Cookbook
 for Dark Matter Indirect Detection**

Marco Cirelli^{a,b}, Gennaro Corcella^{c,d,e}, Andi Hektor^f,
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OUR APPROACH

- Basic question: probability for an initial particle j to produce a particle i carrying momentum fraction x
- General description given by a **fragmentation function**

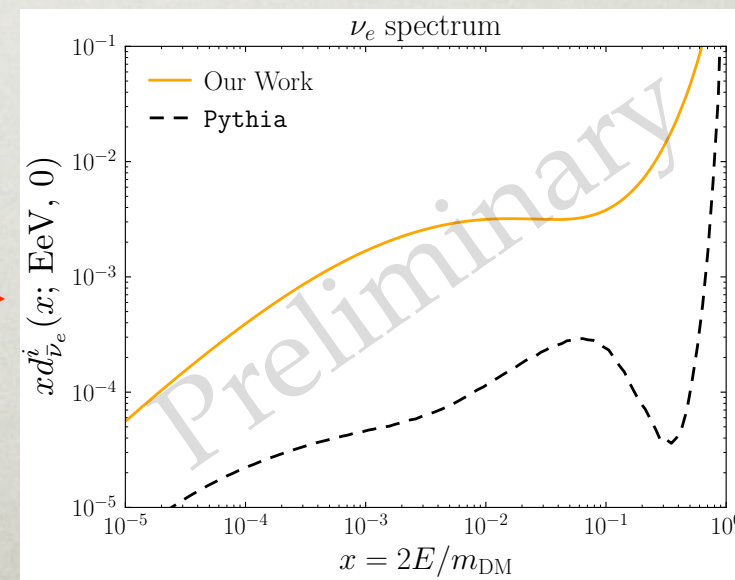
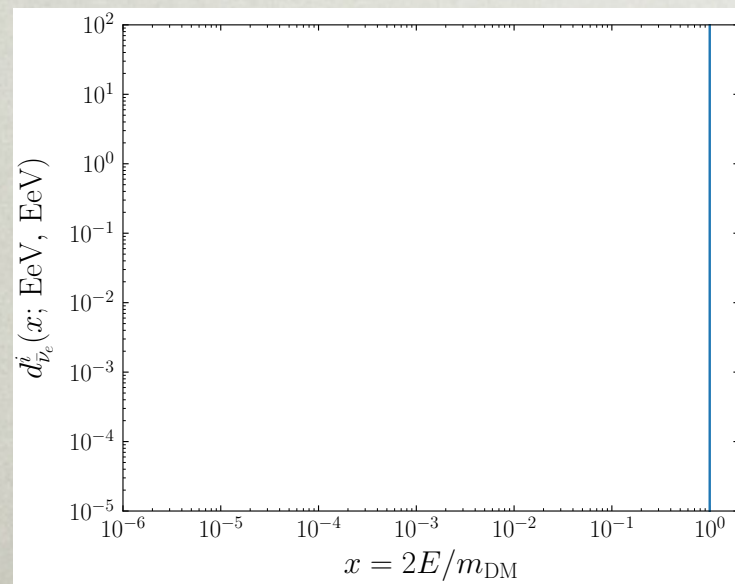
$$d_j^i(x; Q, 0)$$

OUR APPROACH

- Basic question: probability for an initial particle j to produce a particle i carrying momentum fraction x
- General description given by a **fragmentation function**

$$d_j^i(x; Q, 0)$$

- Calculate in three steps:
 1. $Q \rightarrow q_W^+$: DGLAP evolution using the full unbroken SM
 2. $q_W^+ \rightarrow q_W^-$: integrate out weak states ($q_W^\pm = q_W(1 \pm \epsilon)$)
 3. $q_W^- \rightarrow 0$: low energy showering and hadronisation in Pythia

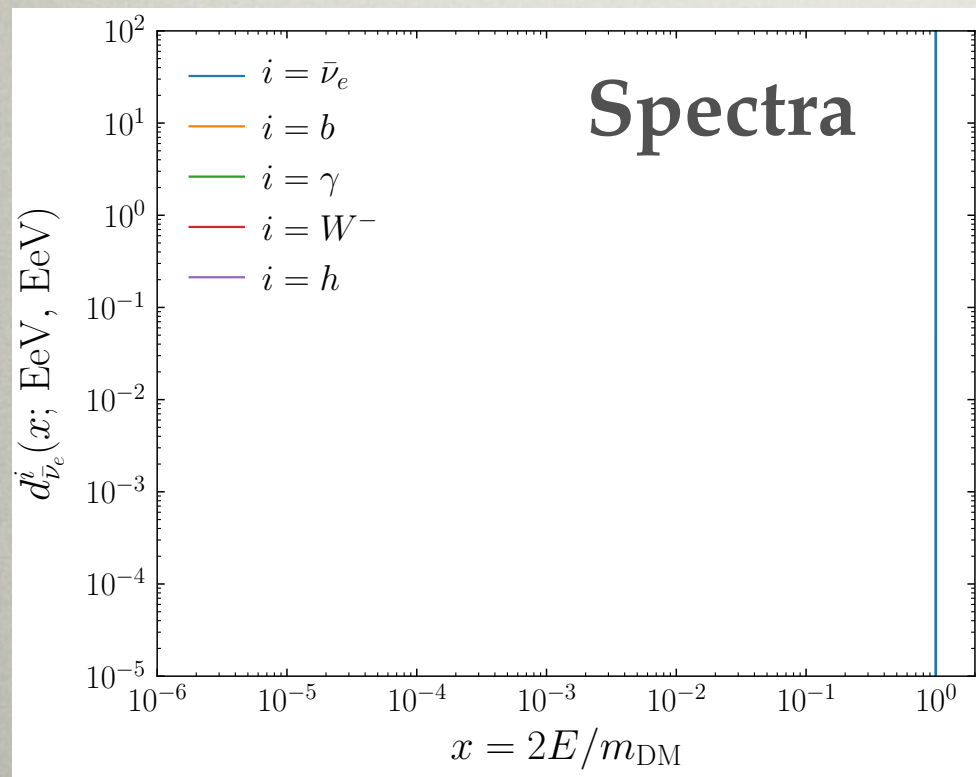




STEP 1: DGLAP

Initial Condition

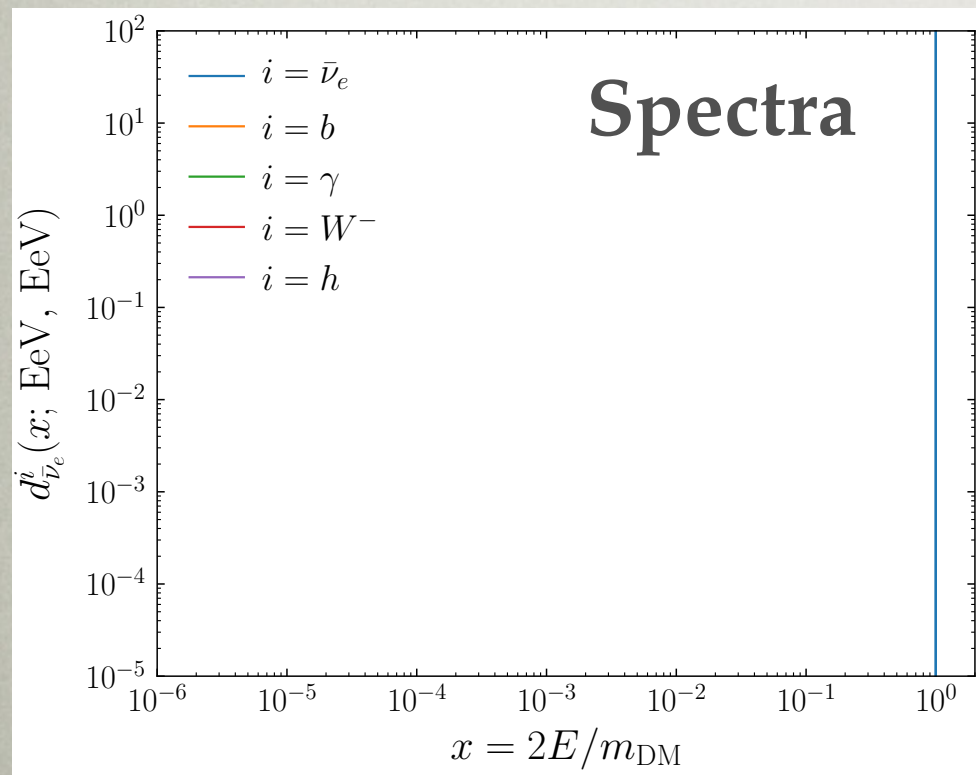
$$d(x; Q, Q) = \delta(1 - x)$$



STEP 1: DGLAP

Evolve with DGLAP,* using all interactions in the unbroken SM

$$\mu \frac{\partial}{\partial \mu} d(x; \mu) = \frac{\alpha}{\pi} \int_x^1 dz \hat{P}(z) d(x/z; \mu) - \frac{\alpha}{\pi} d(x; \mu) \int_0^1 dz \hat{P}(z)$$



***Suppressing particle indices.** In truth 1740 coupled FFs
In truth should have d_i^j , where j can take values:

$$\{u_{L/R}, d_{L/R}, e_{L/R}, \nu_L\} \times 3 \times 2 = 42$$

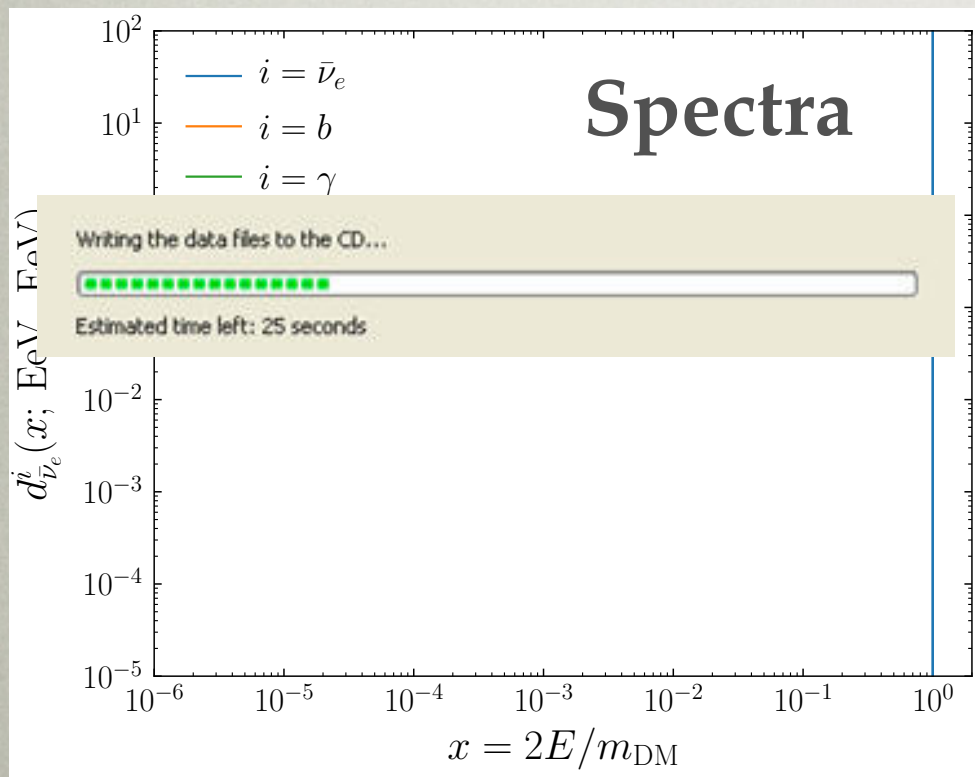
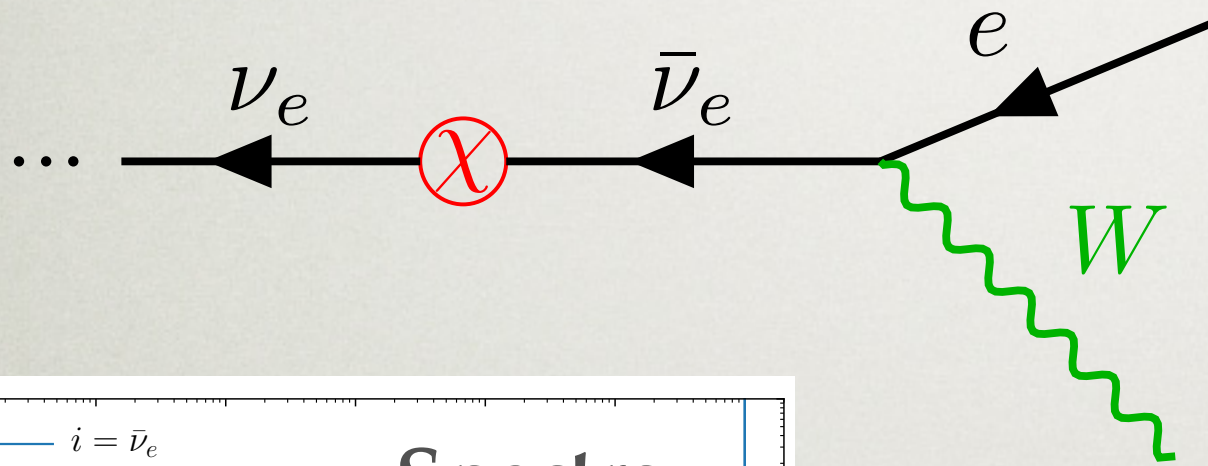
Lots of interesting physics, including electroweak double logs and polarisation generation

For detail see [Bauer, Provasoli, Webber 1806.10157],
[Bauer, Webber 1808.08831]



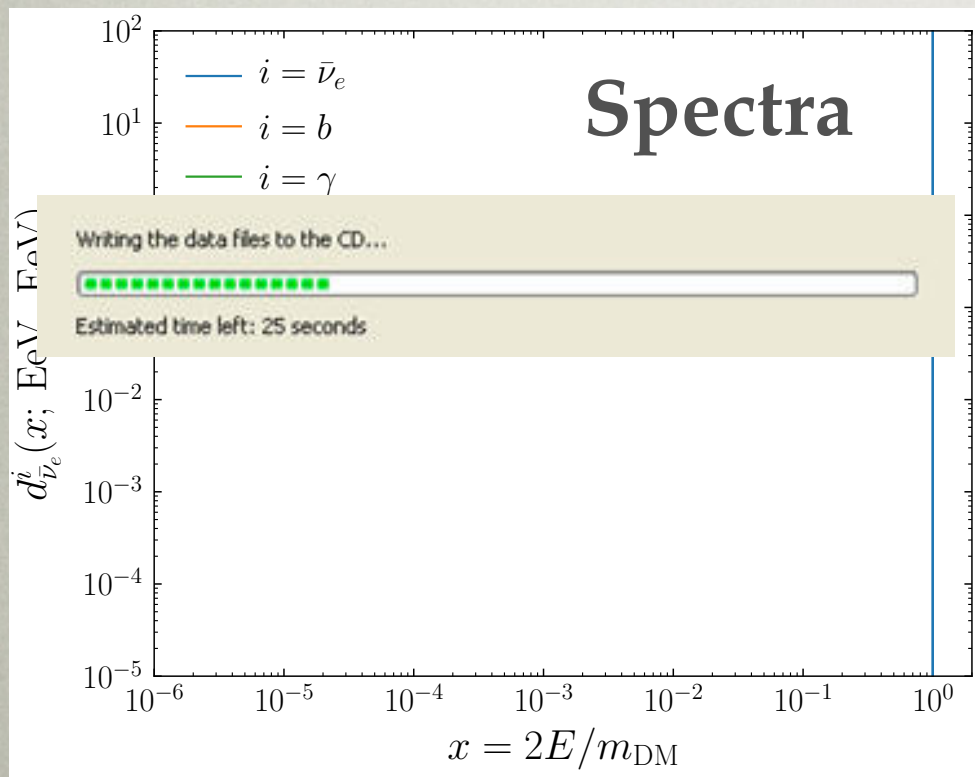
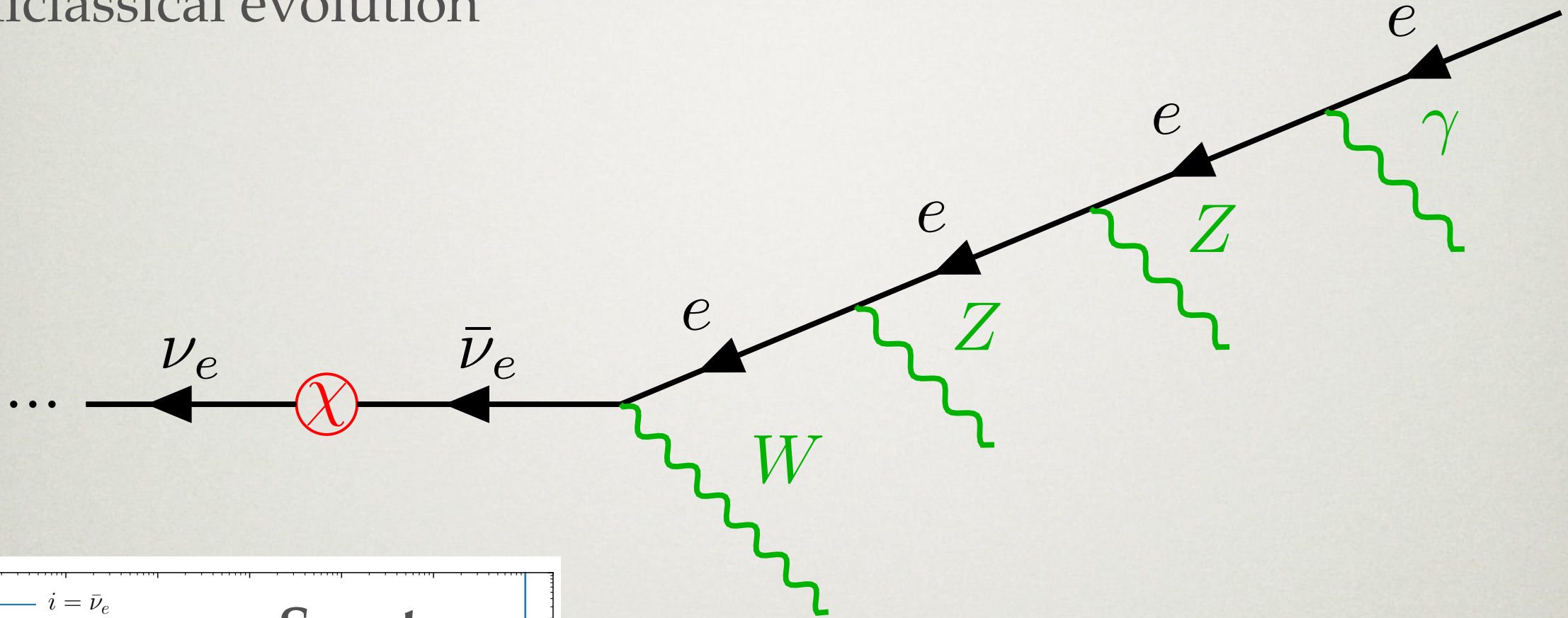
STEP 1: DGLAP

Semiclassical evolution



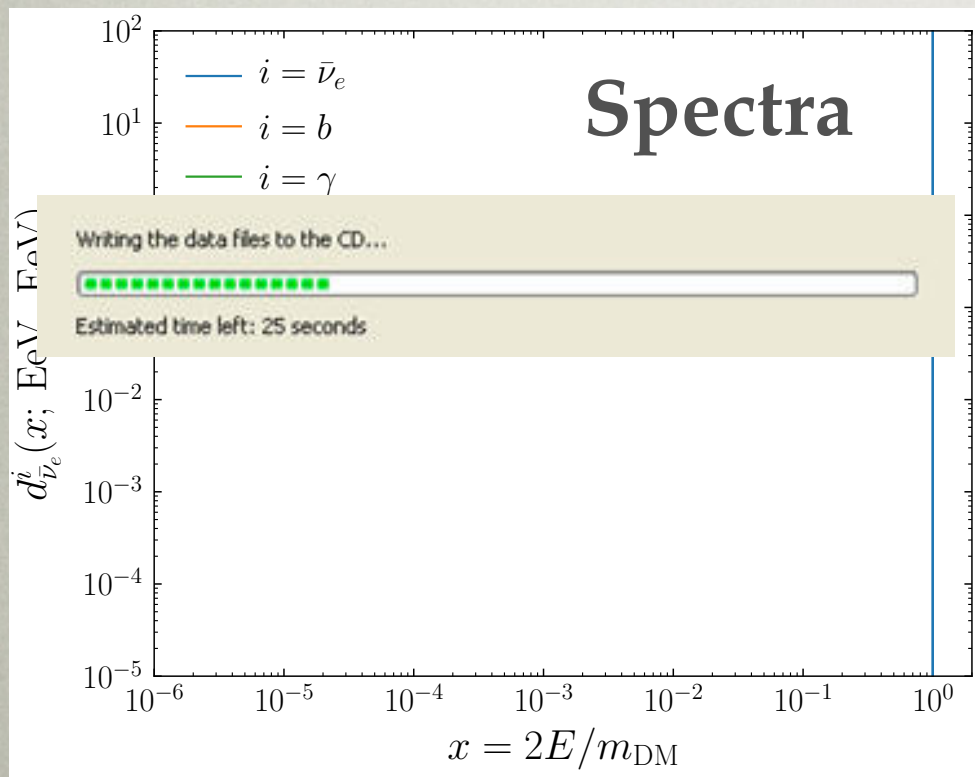
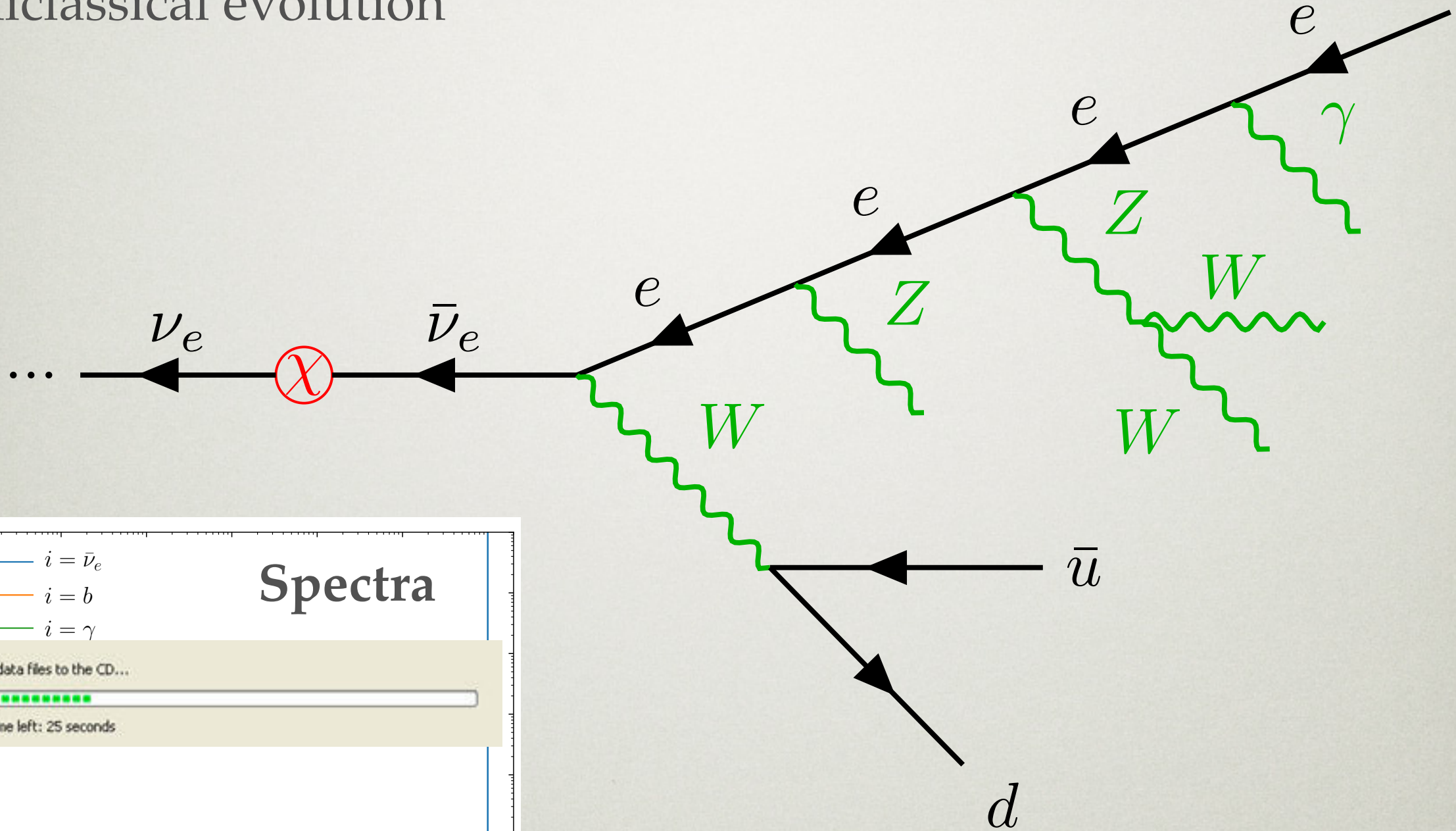
STEP 1: DGLAP

Semiclassical evolution



STEP 1: DGLAP

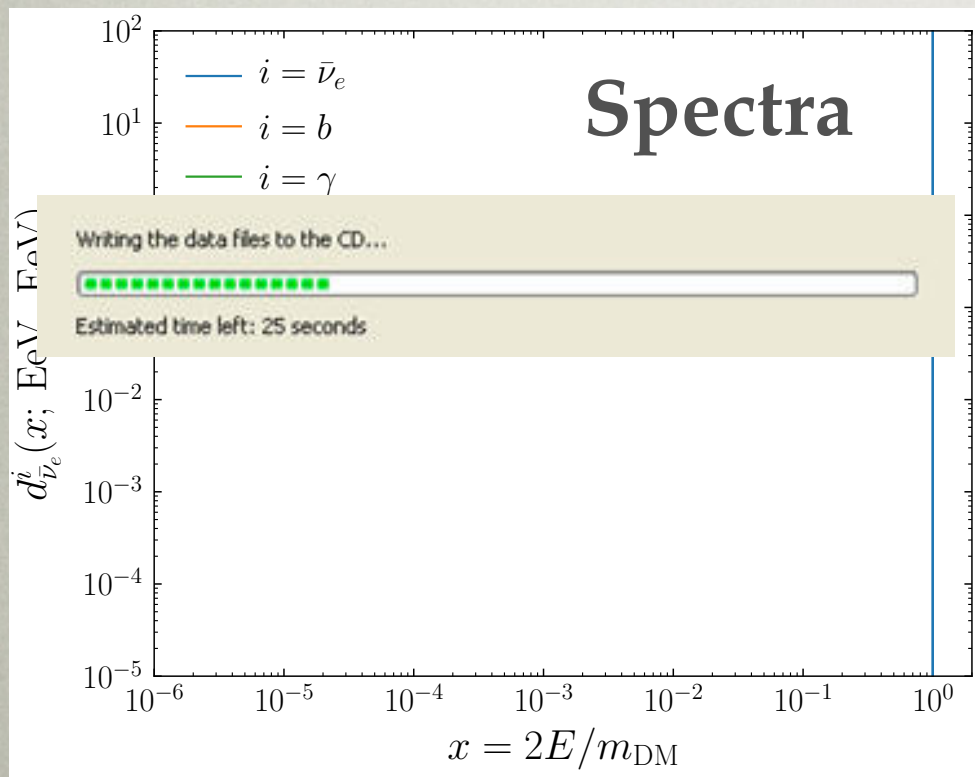
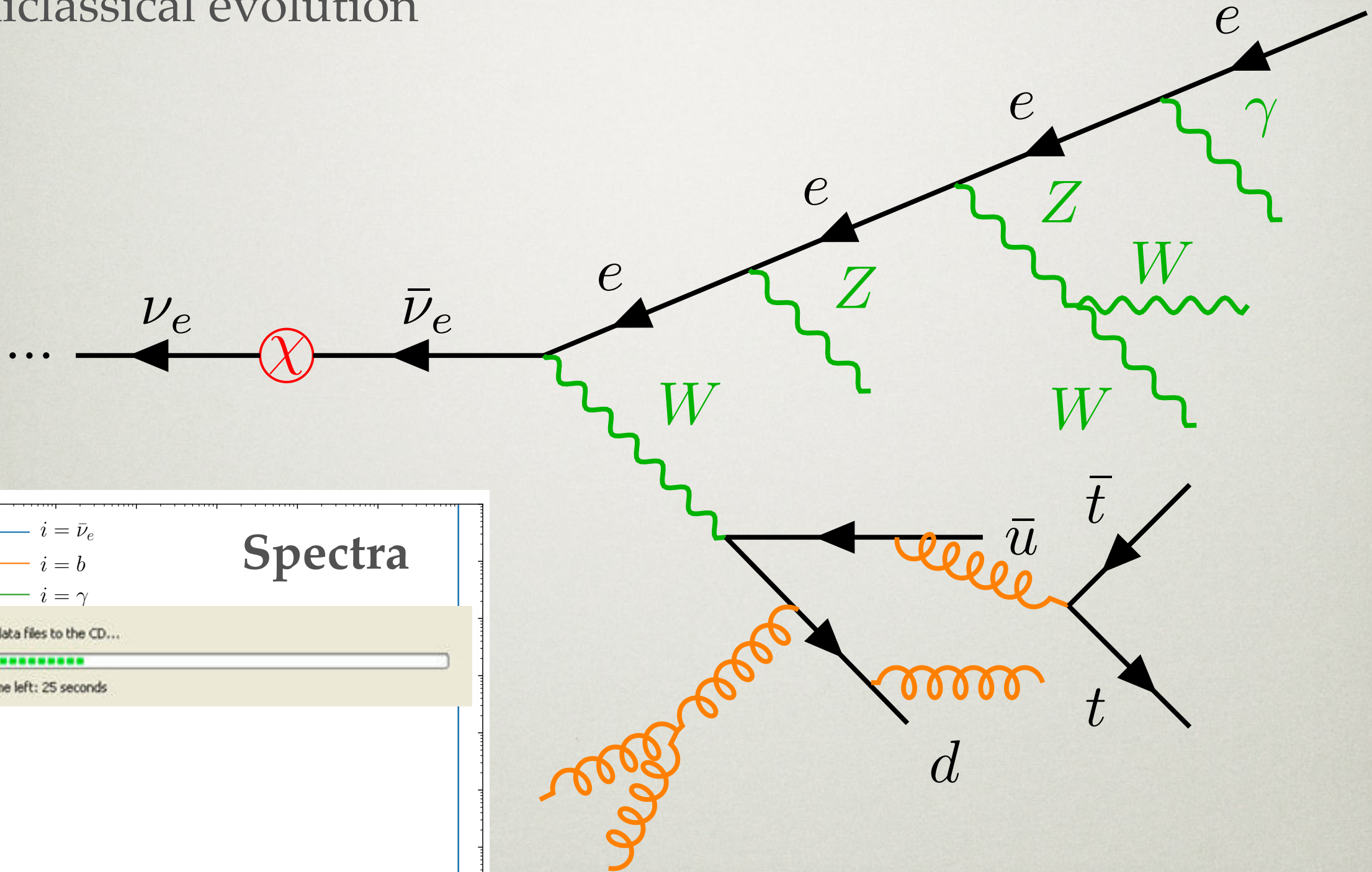
Semiclassical evolution





STEP 1: DGLAP

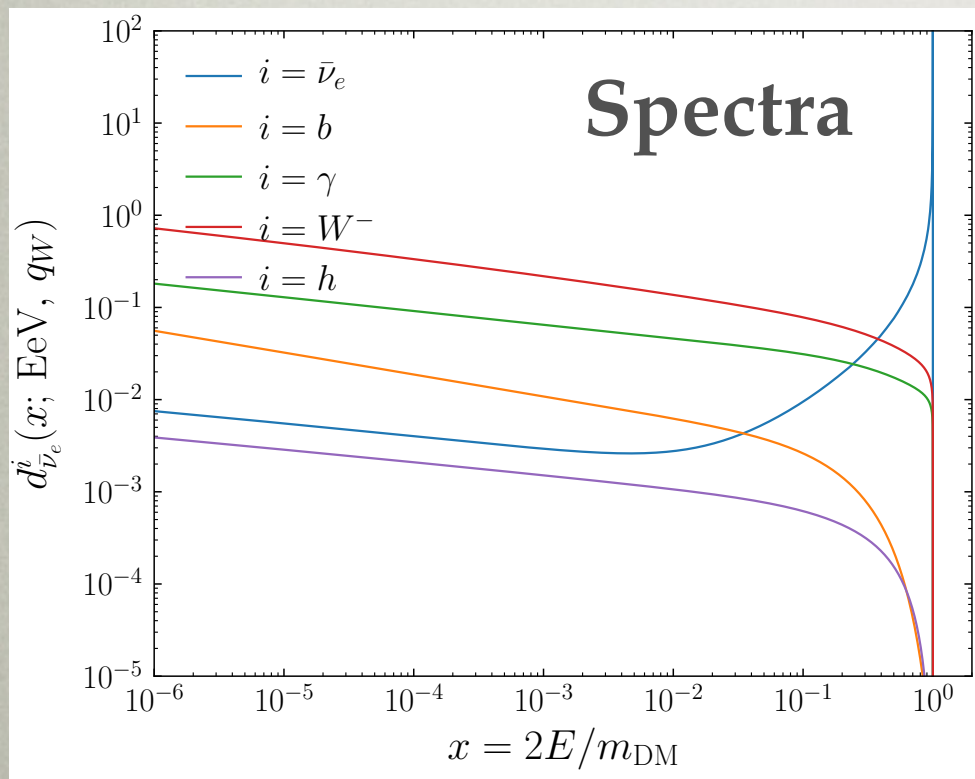
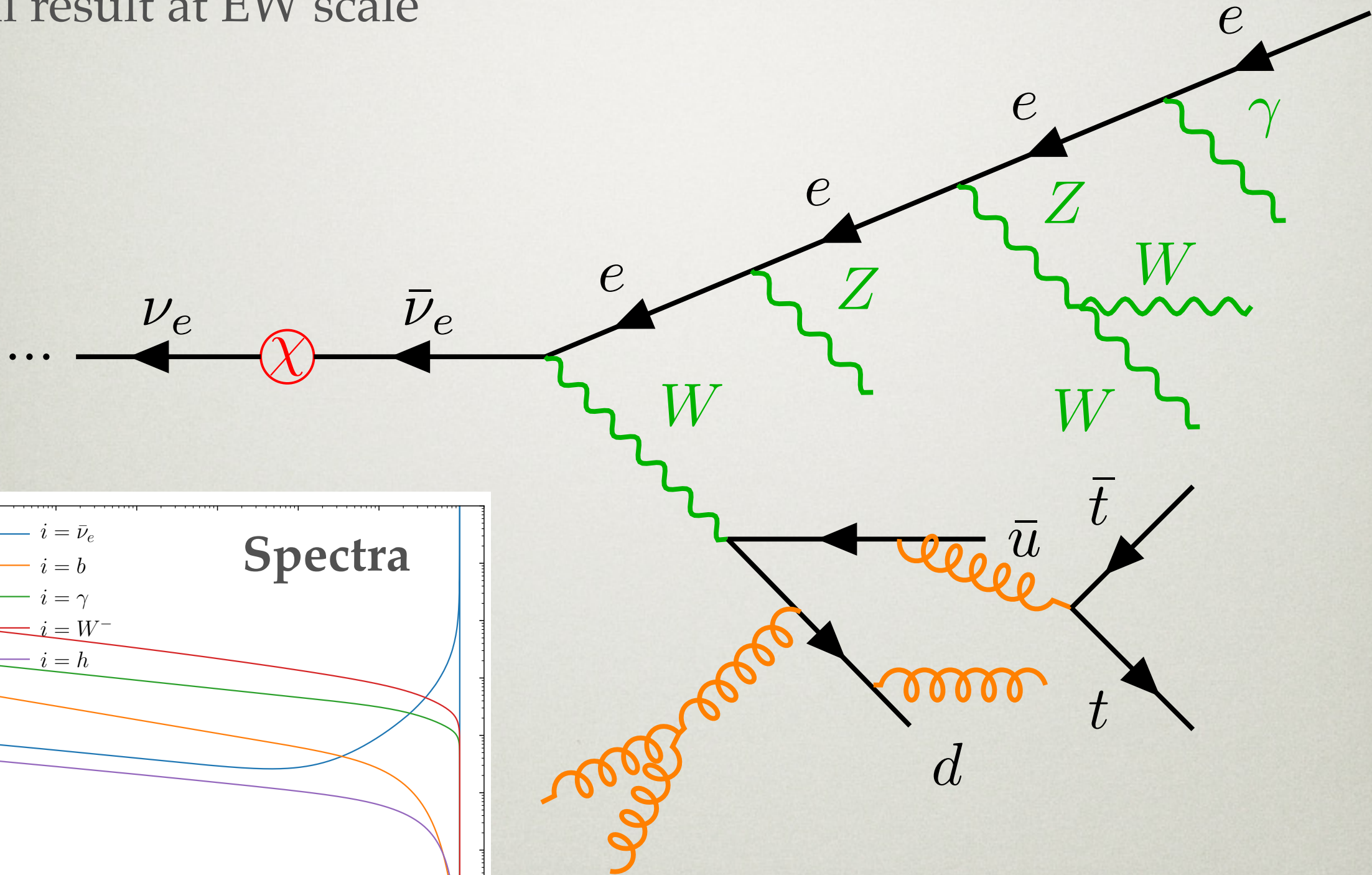
Semiclassical evolution





STEP 1: DGLAP

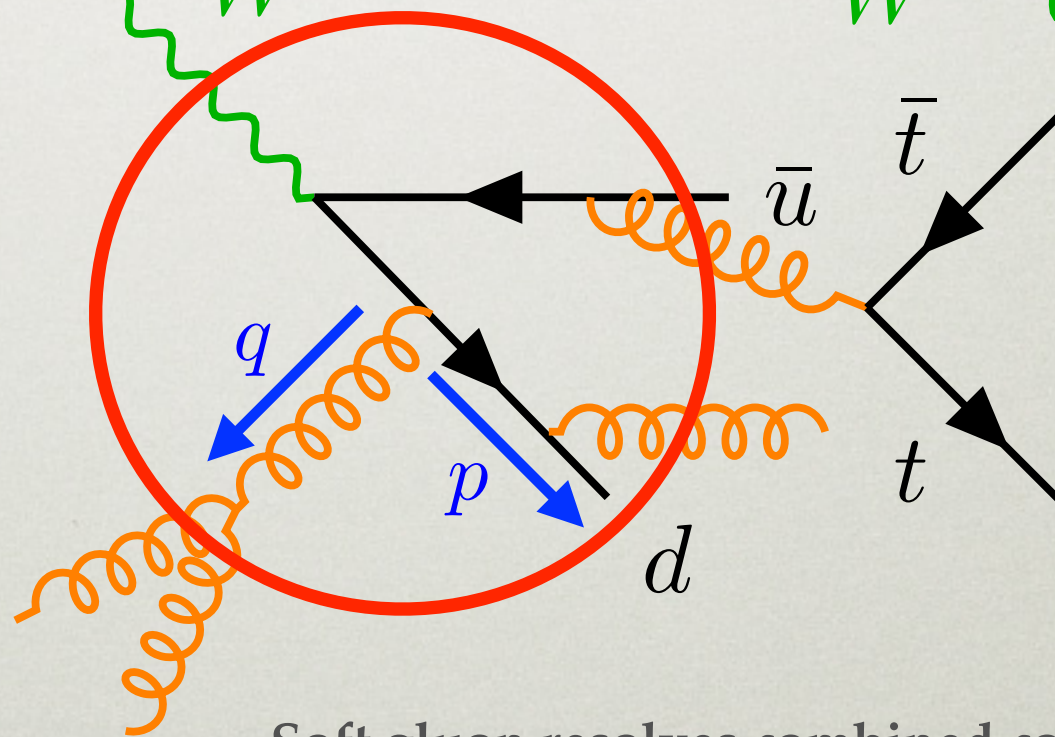
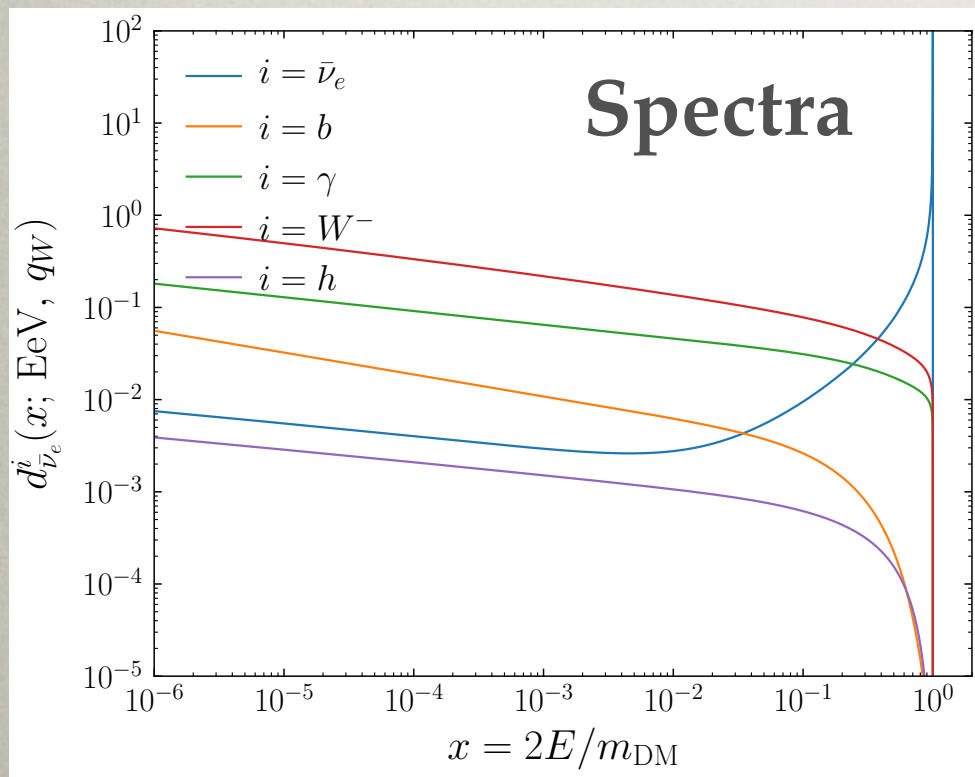
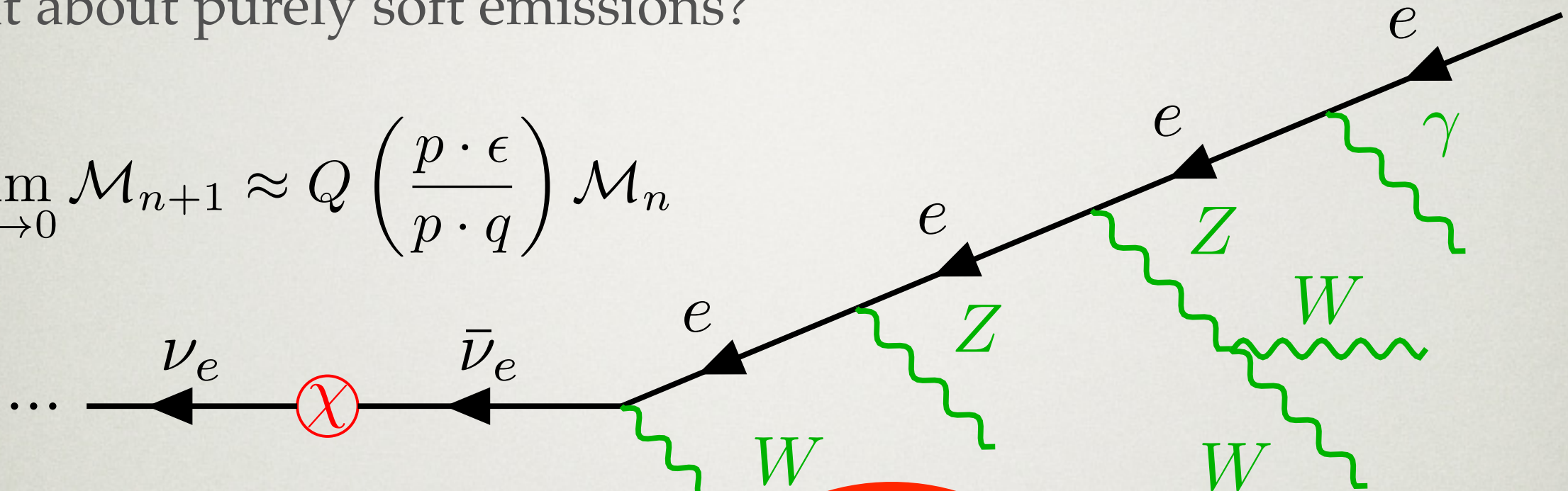
Final result at EW scale



STEP 1: DGLAP

What about purely soft emissions?

$$\lim_{q \rightarrow 0} \mathcal{M}_{n+1} \approx Q \left(\frac{p \cdot \epsilon}{p \cdot q} \right) \mathcal{M}_n$$

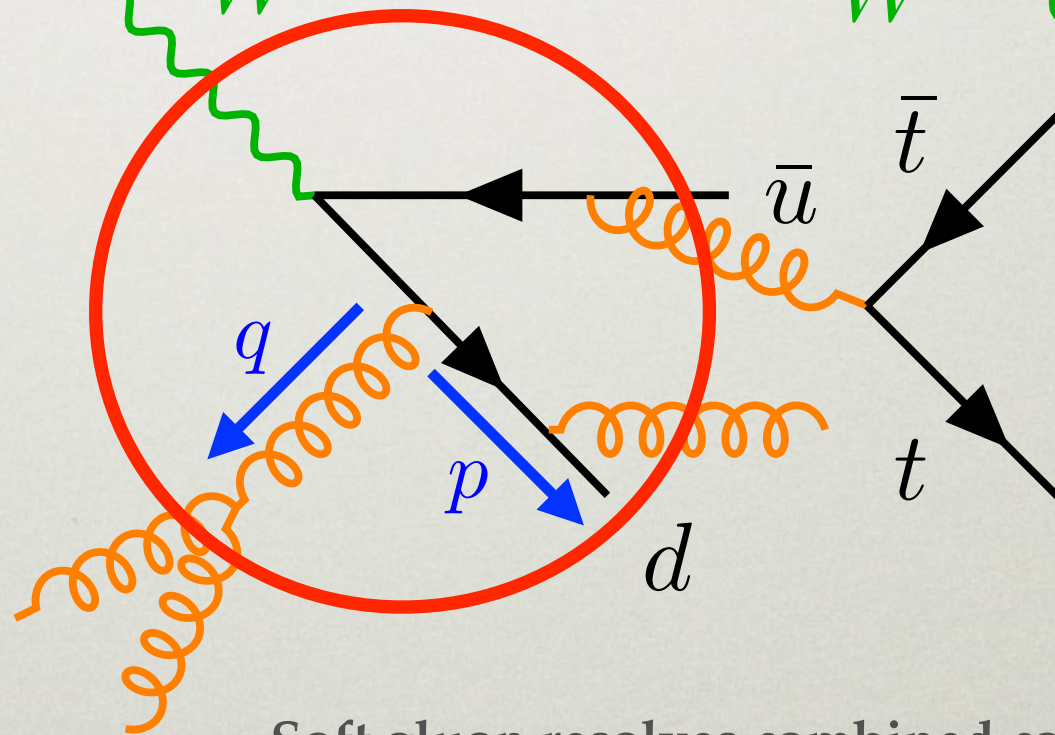
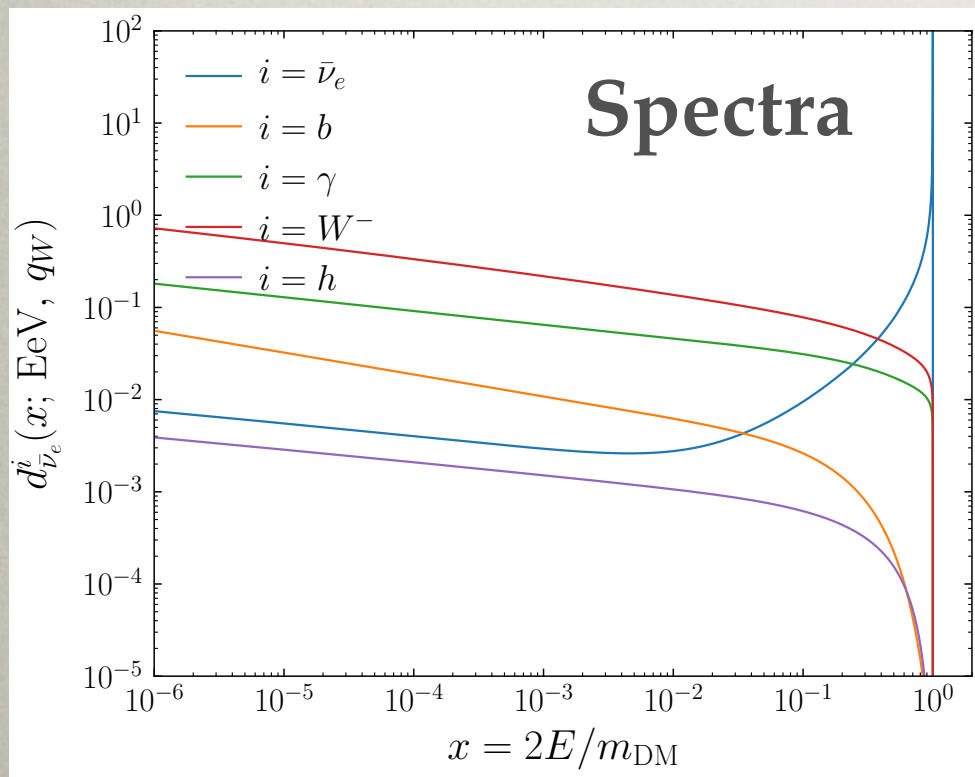
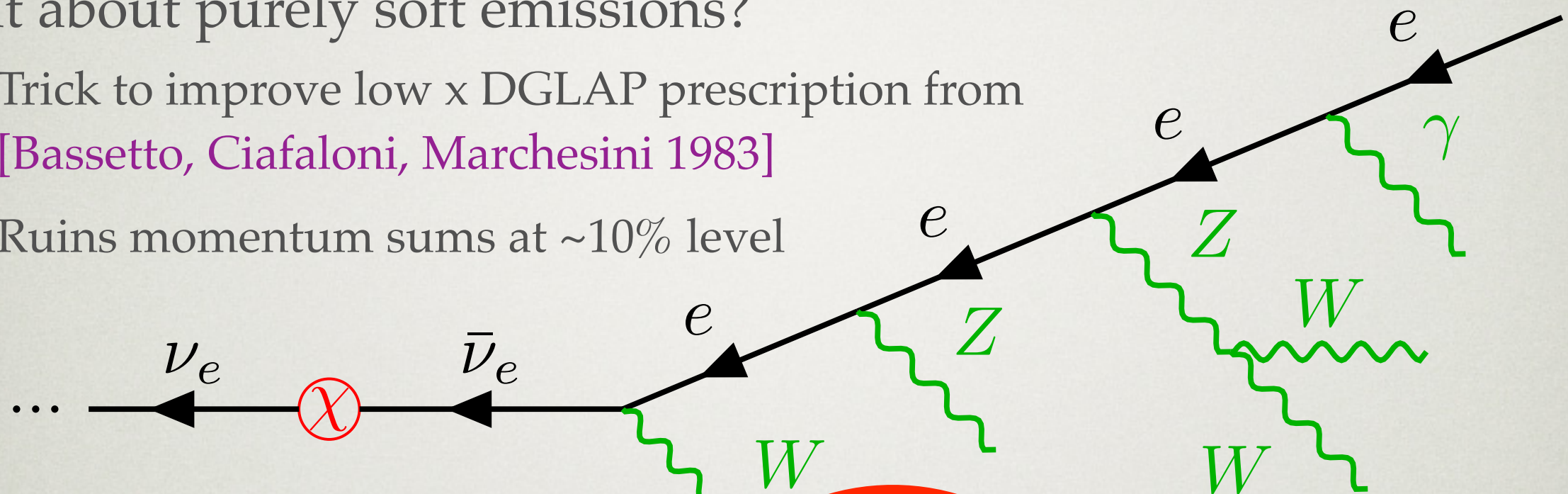


Soft gluon resolves combined color = 0, cannot emit!
 Generically destructive interference; "color coherence"

STEP 1: DGLAP

What about purely soft emissions?

- Trick to improve low x DGLAP prescription from [Bassetto, Ciafaloni, Marchesini 1983]
- Ruins momentum sums at $\sim 10\%$ level

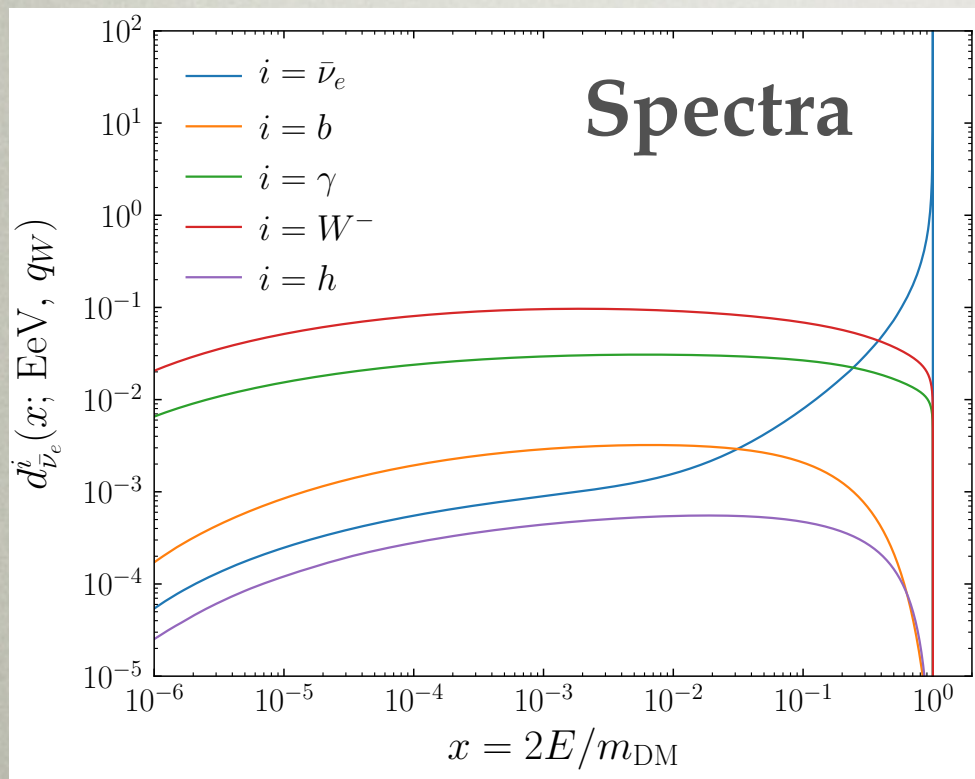
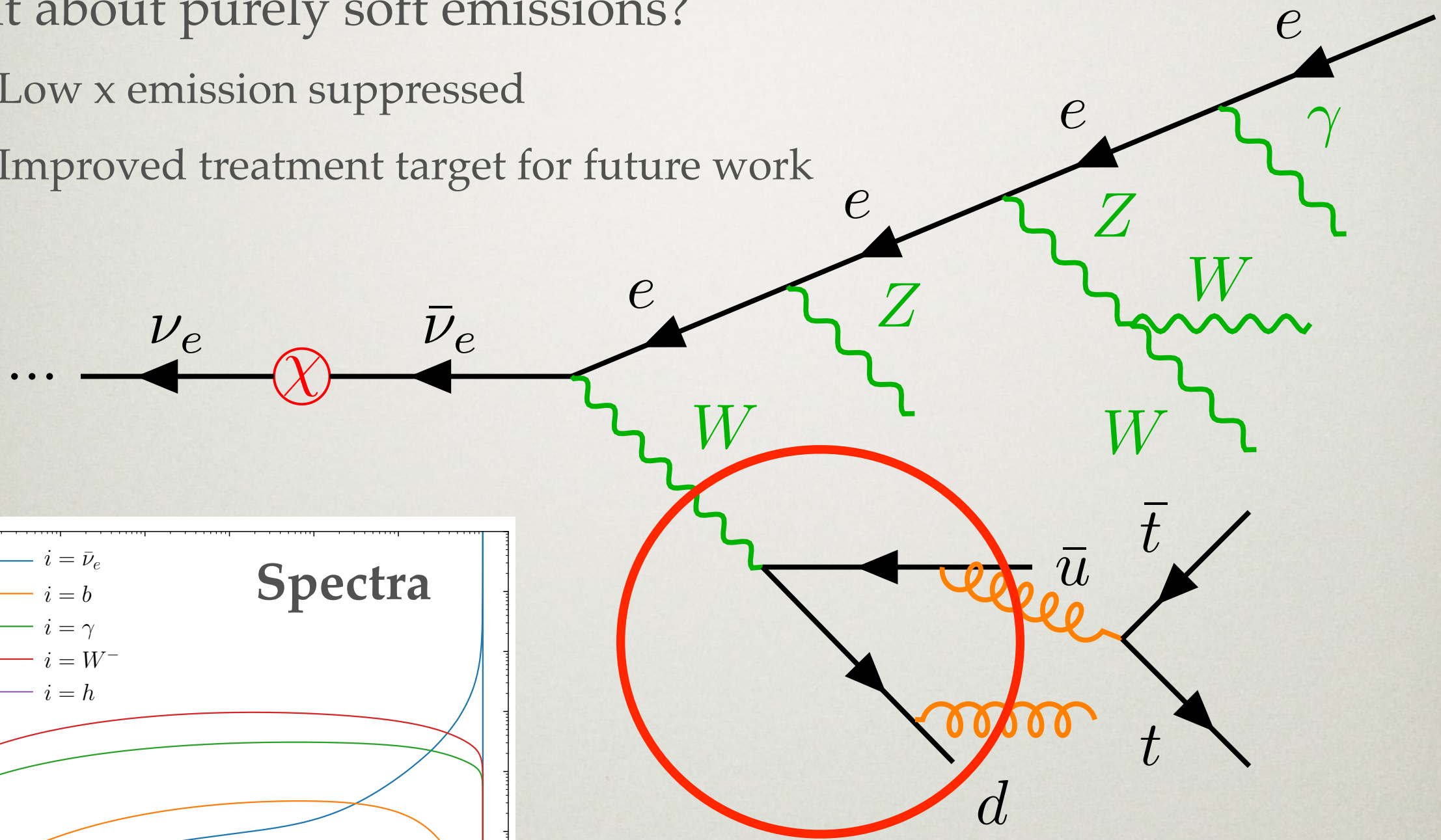


Soft gluon resolves combined color = 0, cannot emit!
 Generically destructive interference; "color coherence"

STEP 1: DGLAP

What about purely soft emissions?

- Low x emission suppressed
- Improved treatment target for future work



Soft gauge boson radiation suppressed



STEP 2: WEAK MATCHING

- Exploit FF convolution property

$$\begin{aligned}d(x; Q, q_{\bar{W}}^-) &= \int_x^1 \frac{dz}{z} d(z; Q, q_W^+) d(x/z; q_W^+, q_{\bar{W}}^-) \\ &= d(x; Q, q_W^+) \otimes d(x; q_W^+, q_{\bar{W}}^-)\end{aligned}$$

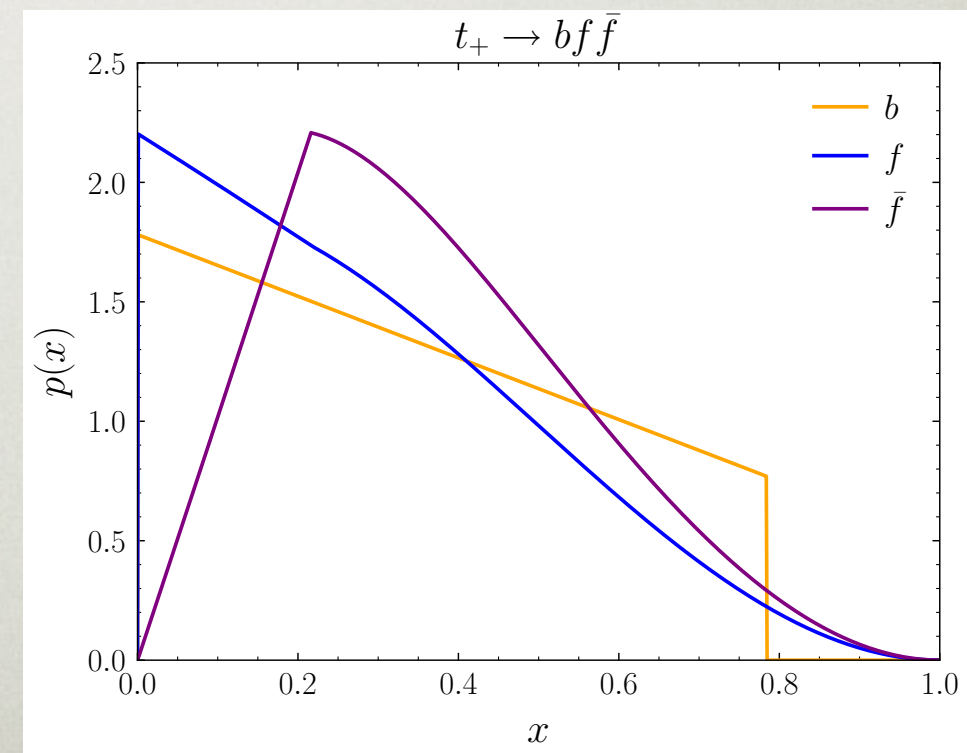
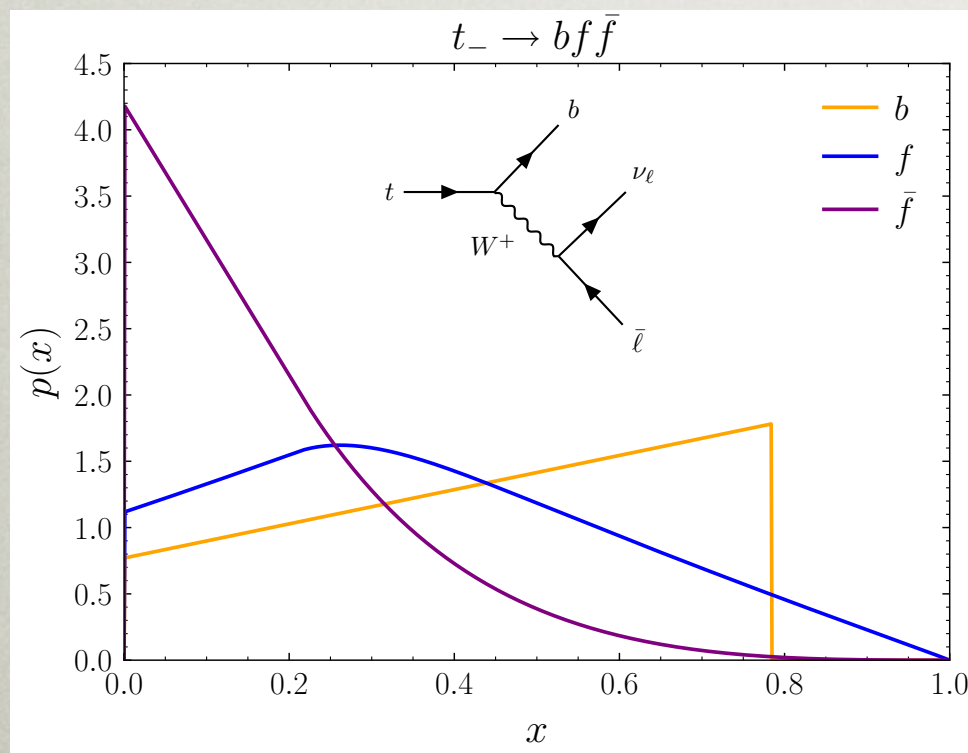
STEP 2: WEAK MATCHING

- Exploit FF convolution property

$$d(x; Q, q_{\bar{W}}^-) = \int_x^1 \frac{dz}{z} d(z; Q, q_W^+) d(x/z; q_W^+, q_{\bar{W}}^-)$$

$$= d(x; Q, q_W^+) \otimes d(x; q_W^+, q_{\bar{W}}^-)$$

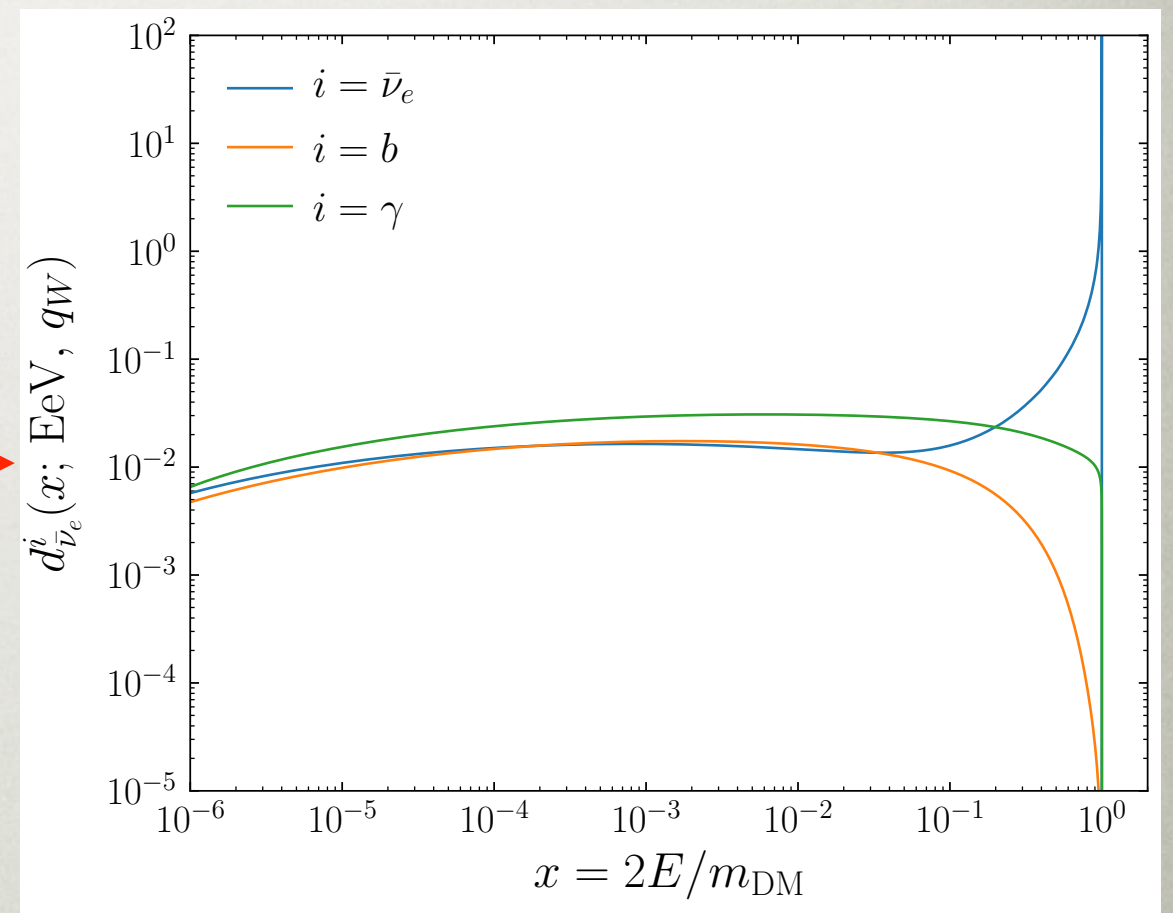
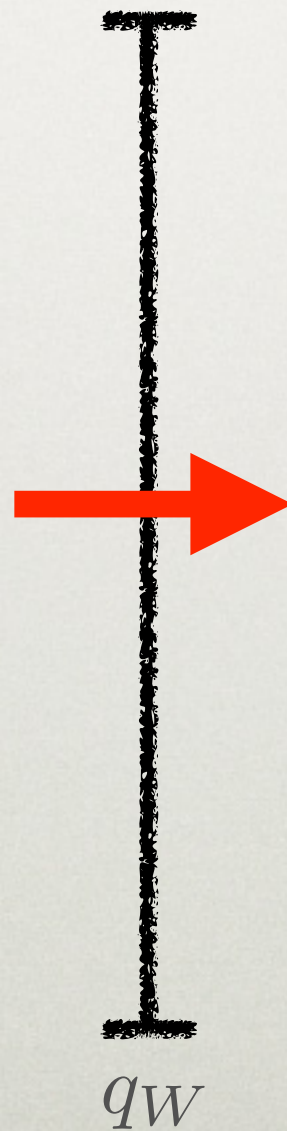
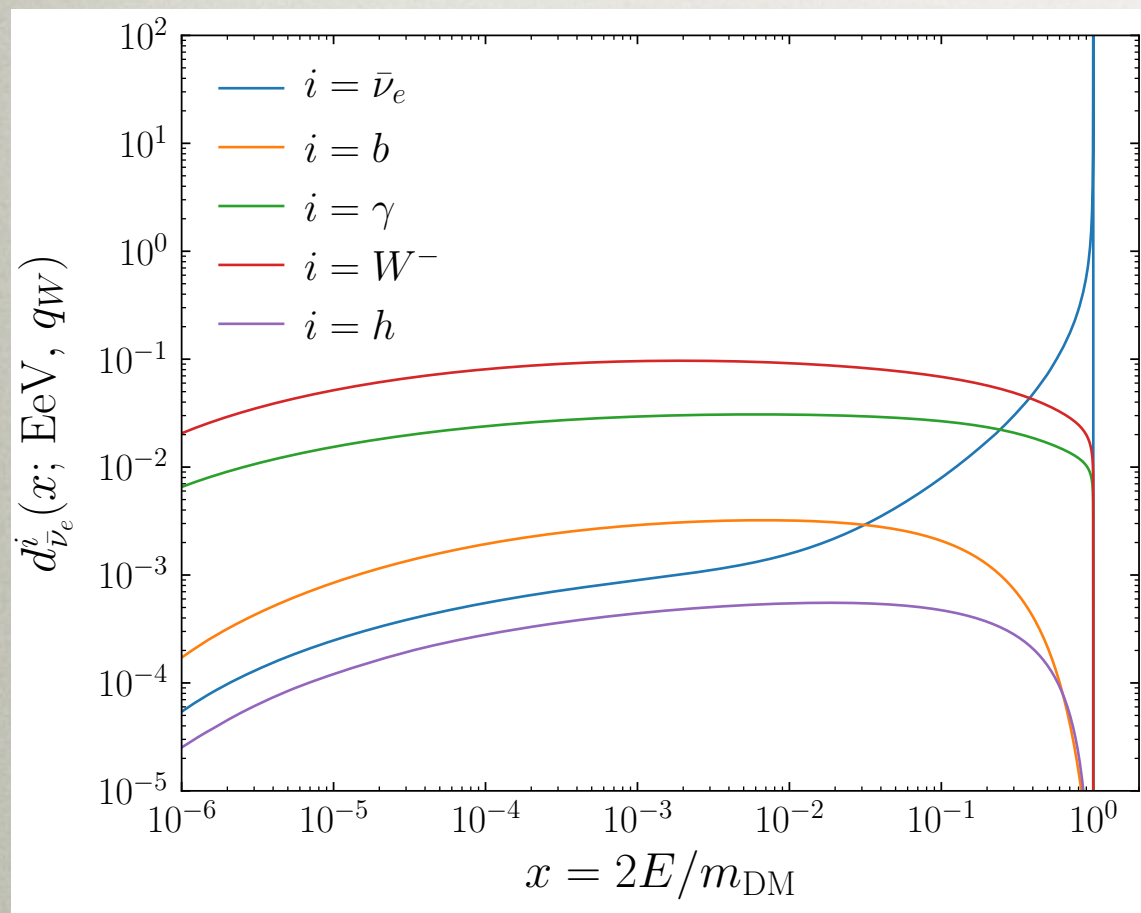
- Most states do not evolve, EW scale states (Z, W, h, t) decay
- Must account for polarised decays, all analytically



STEP 2: WEAK MATCHING

Impact on the spectra

$$d(x; Q, q_{\bar{W}}^-) = d(x; Q, q_W^+) \otimes d(x; q_W^+, q_{\bar{W}}^-)$$



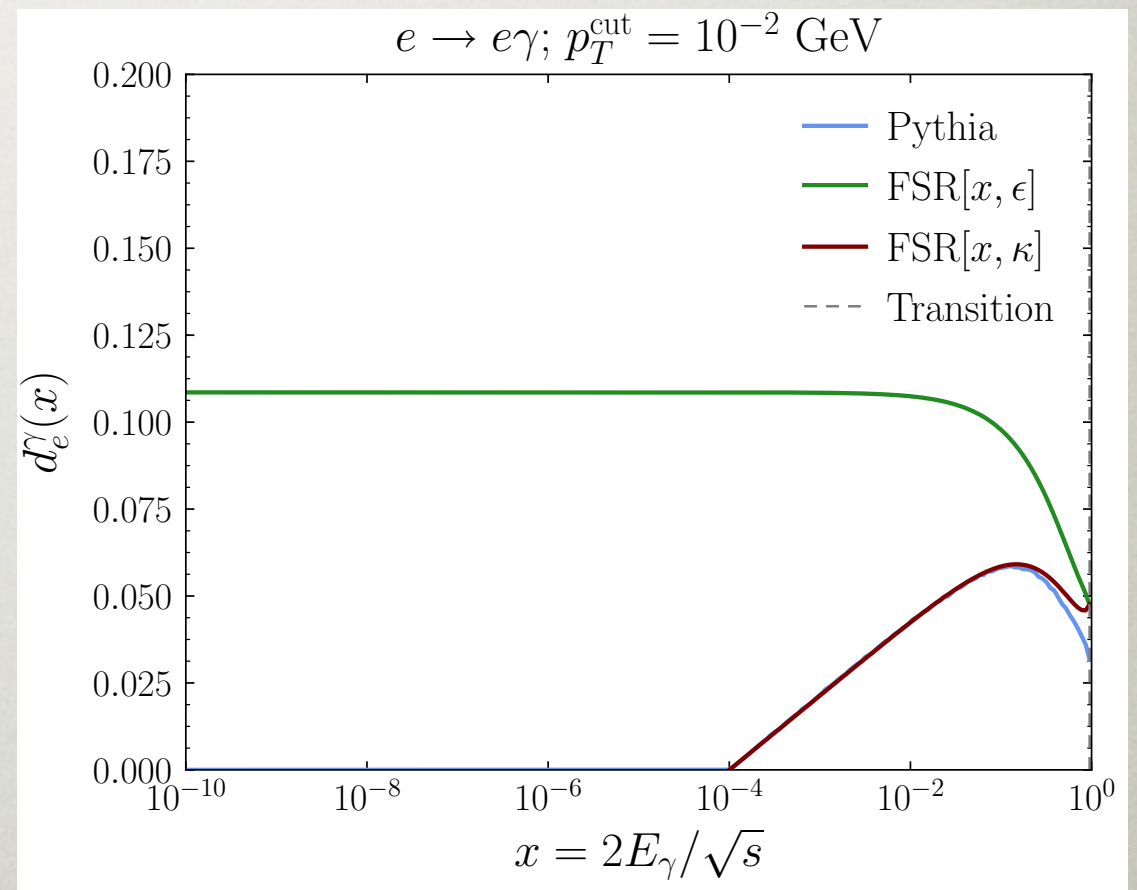
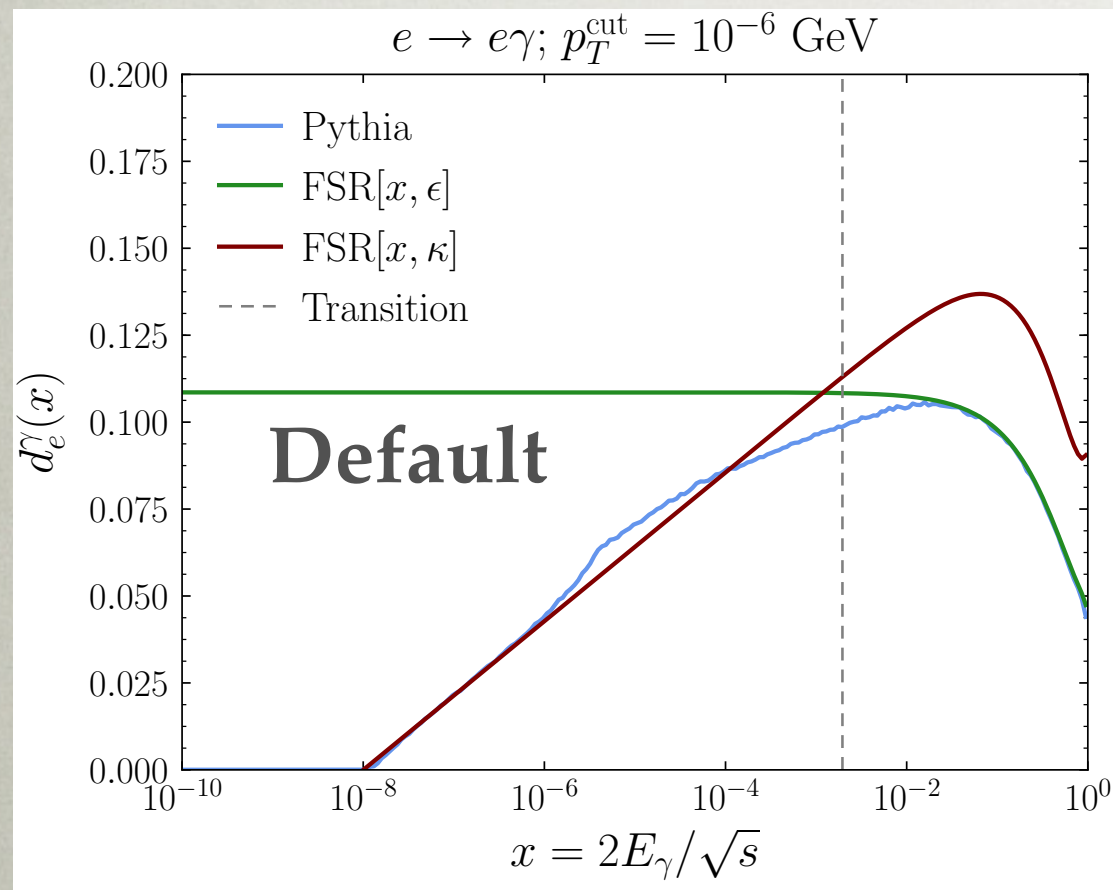


STEP 3: CONVOLVE WITH PYTHIA

- Use Pythia for showering and hadronization below EW scale

$$d(x; Q, 0) = d(x; Q, q_{W^-}) \otimes d(x; q_{W^-}, 0)$$

- **Caveat:** Pythia's treatment of photon FSR
- Implement emission (& momentum corrections) analytically



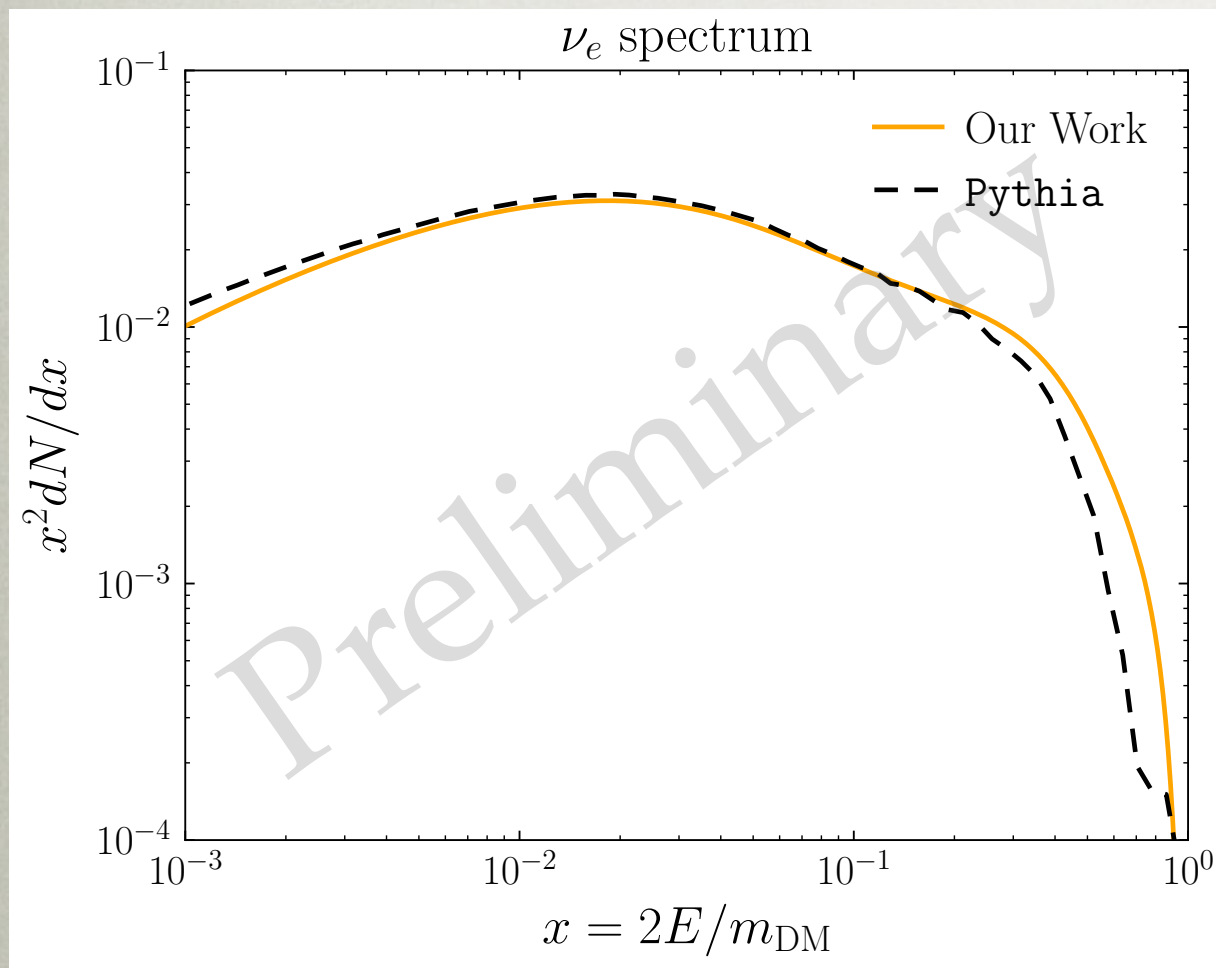


RESULTS

- All together obtain spectra for all stable final state particles
- Impact for hadronic channels less pronounced

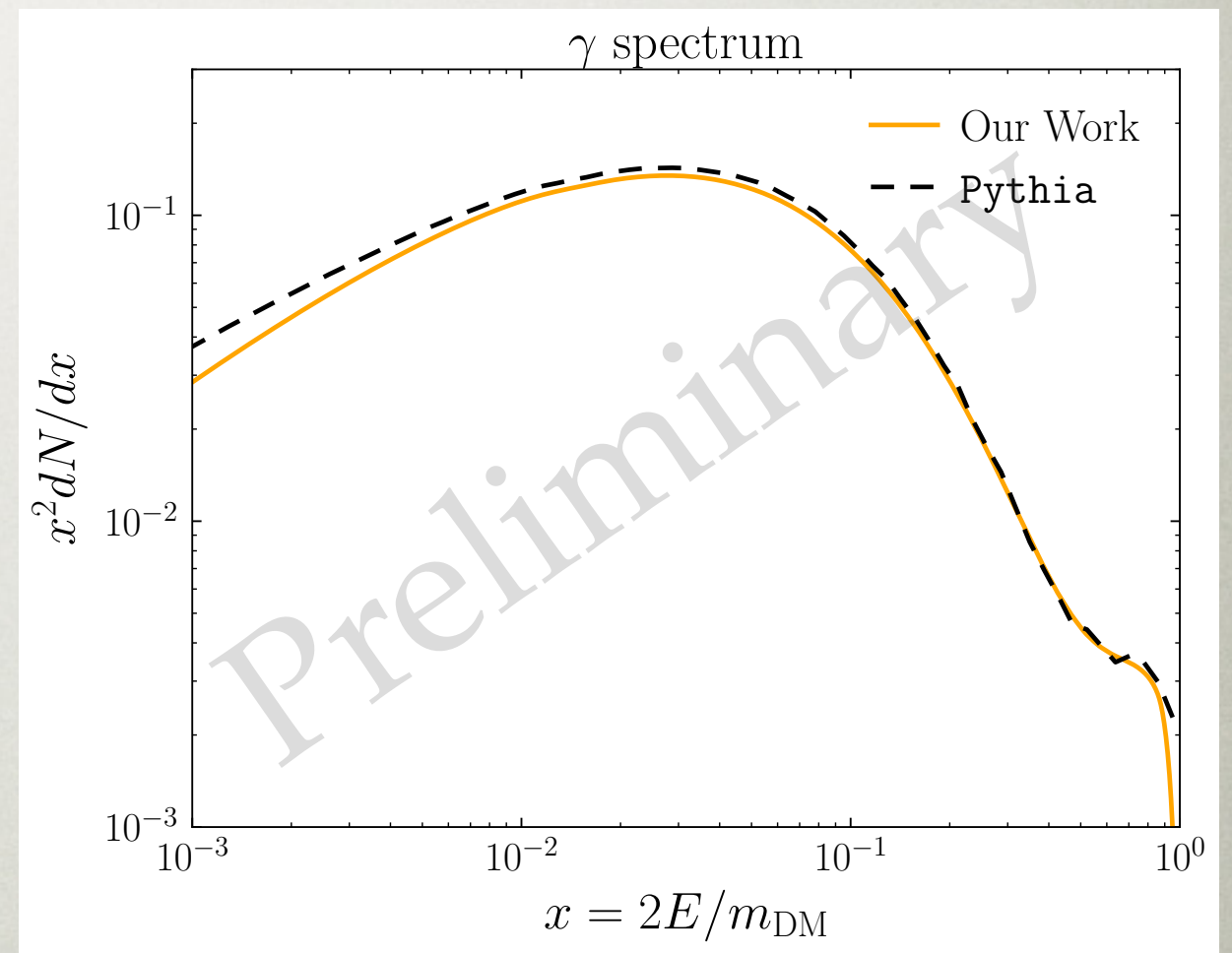
$$\chi \rightarrow b\bar{b}$$

ν_e spectrum



$$m_\chi = 2 \text{ PeV}$$

γ spectrum

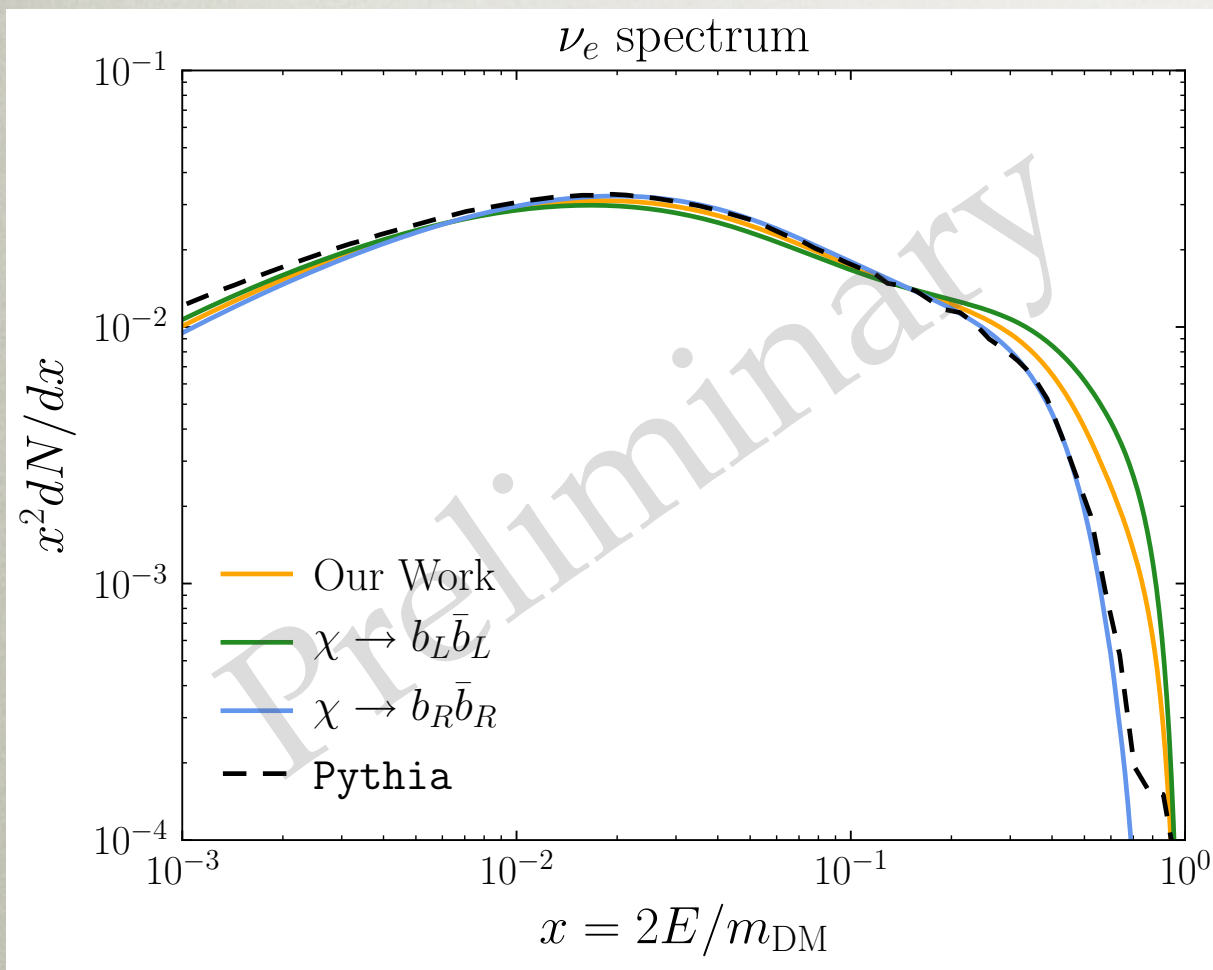


RESULTS

- We have results for arbitrary unbroken initial states
- Can also go to higher energies than are practical in Pythia

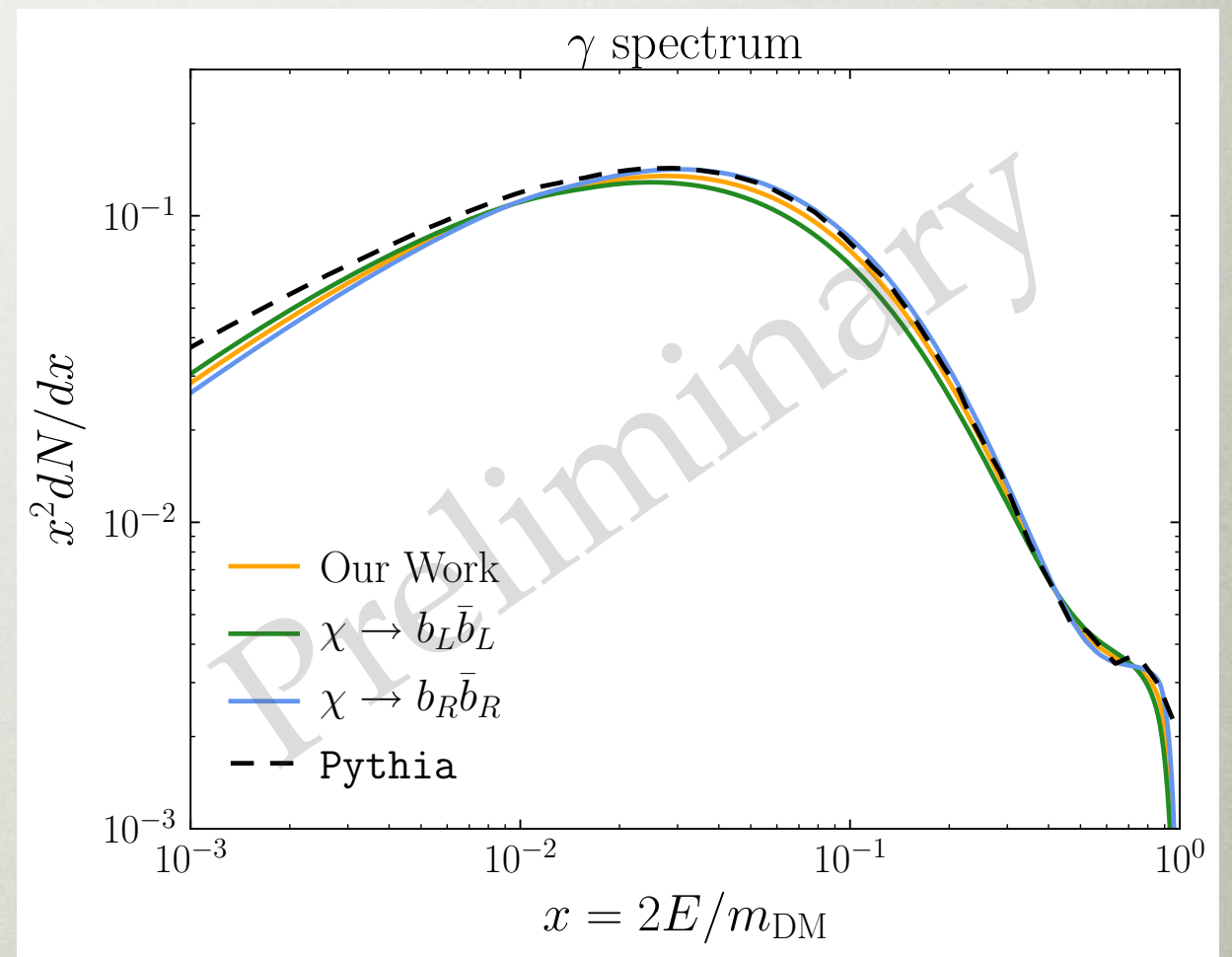
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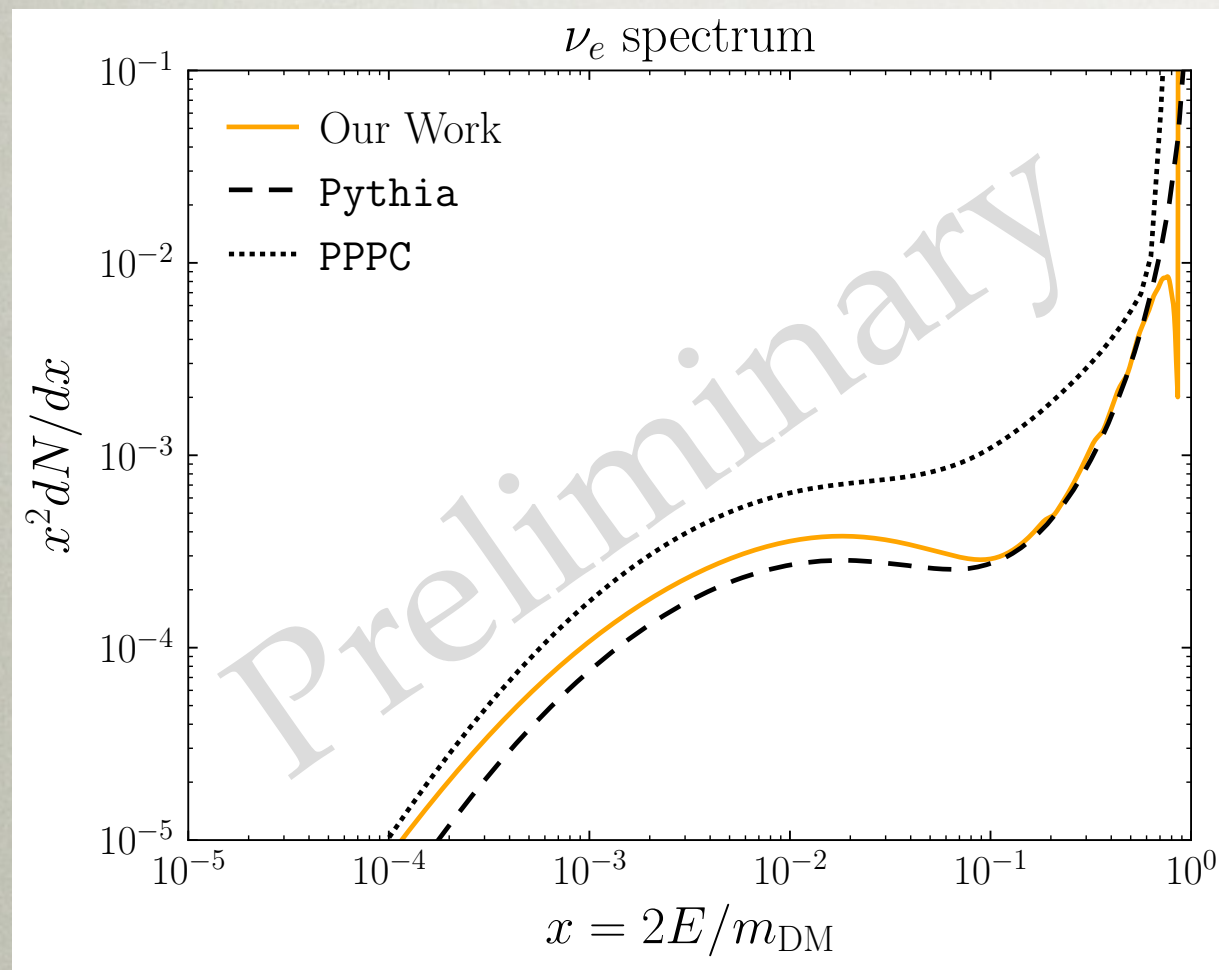
γ spectrum



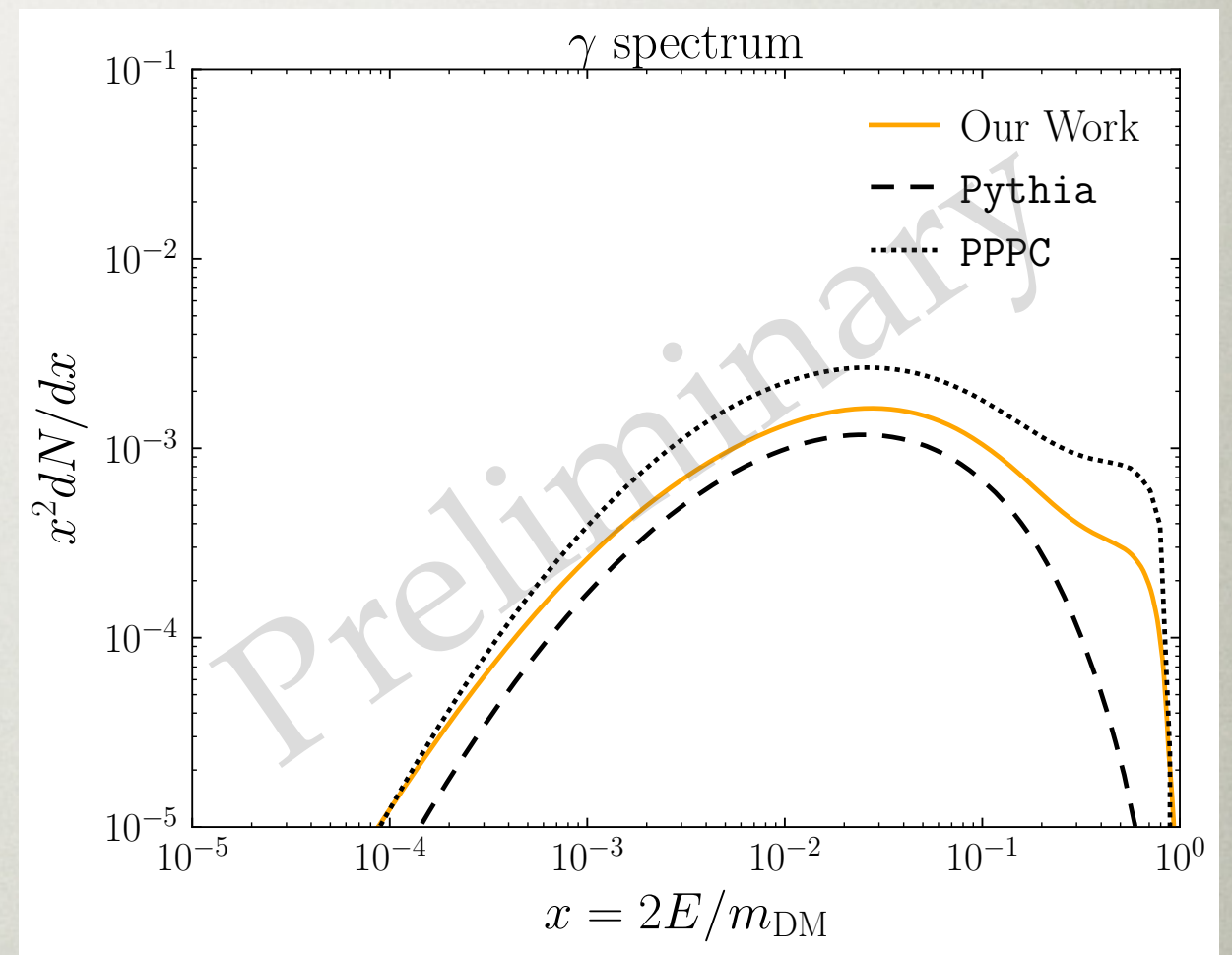
RESULTS

- Larger impact for electroweak dominated final states
- Differ from existing results even at lower masses

$$\chi \rightarrow \nu_e \bar{\nu}_e$$



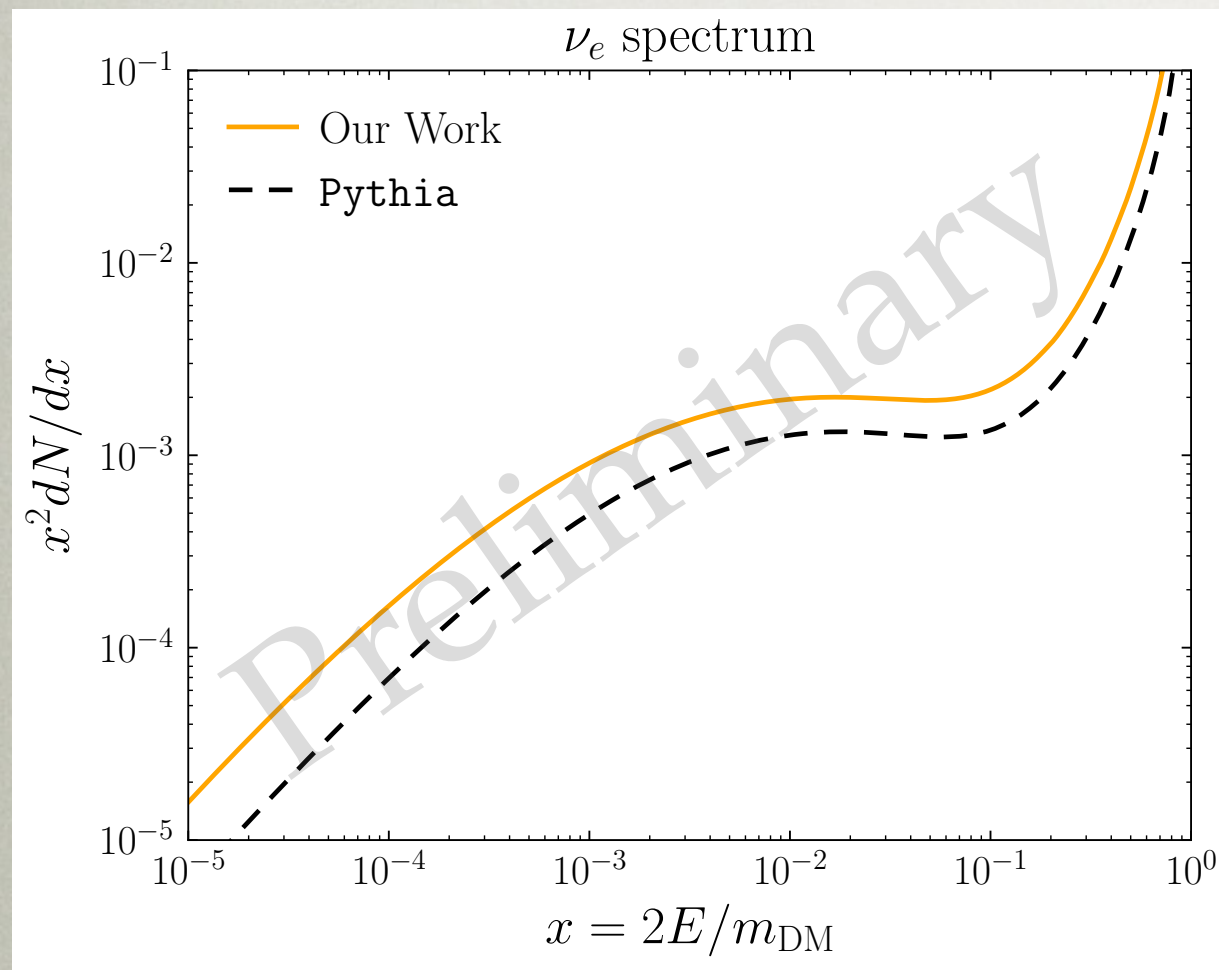
$$m_\chi = 2 \text{ TeV}$$



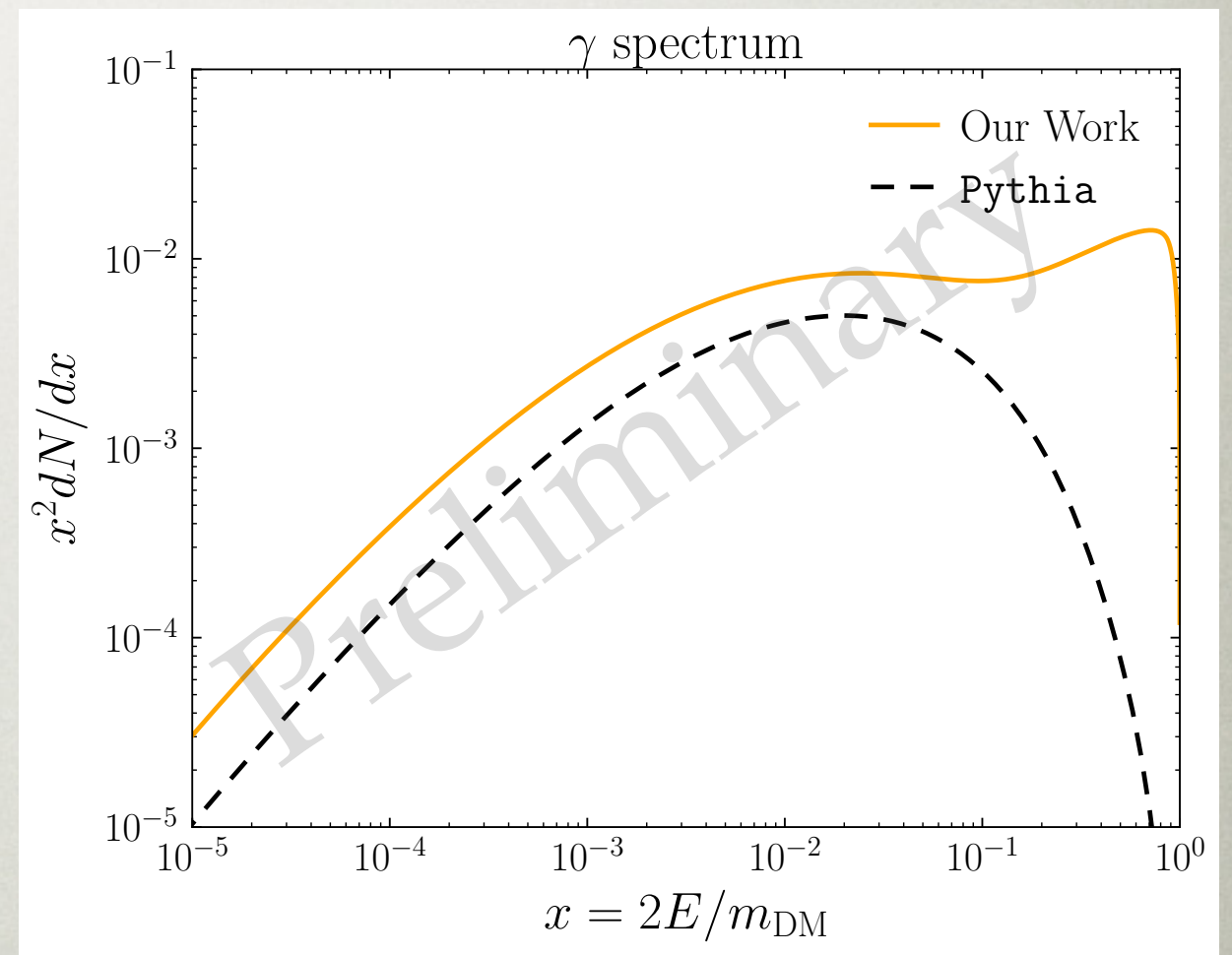
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$$m_\chi = 2 \text{ PeV}$$

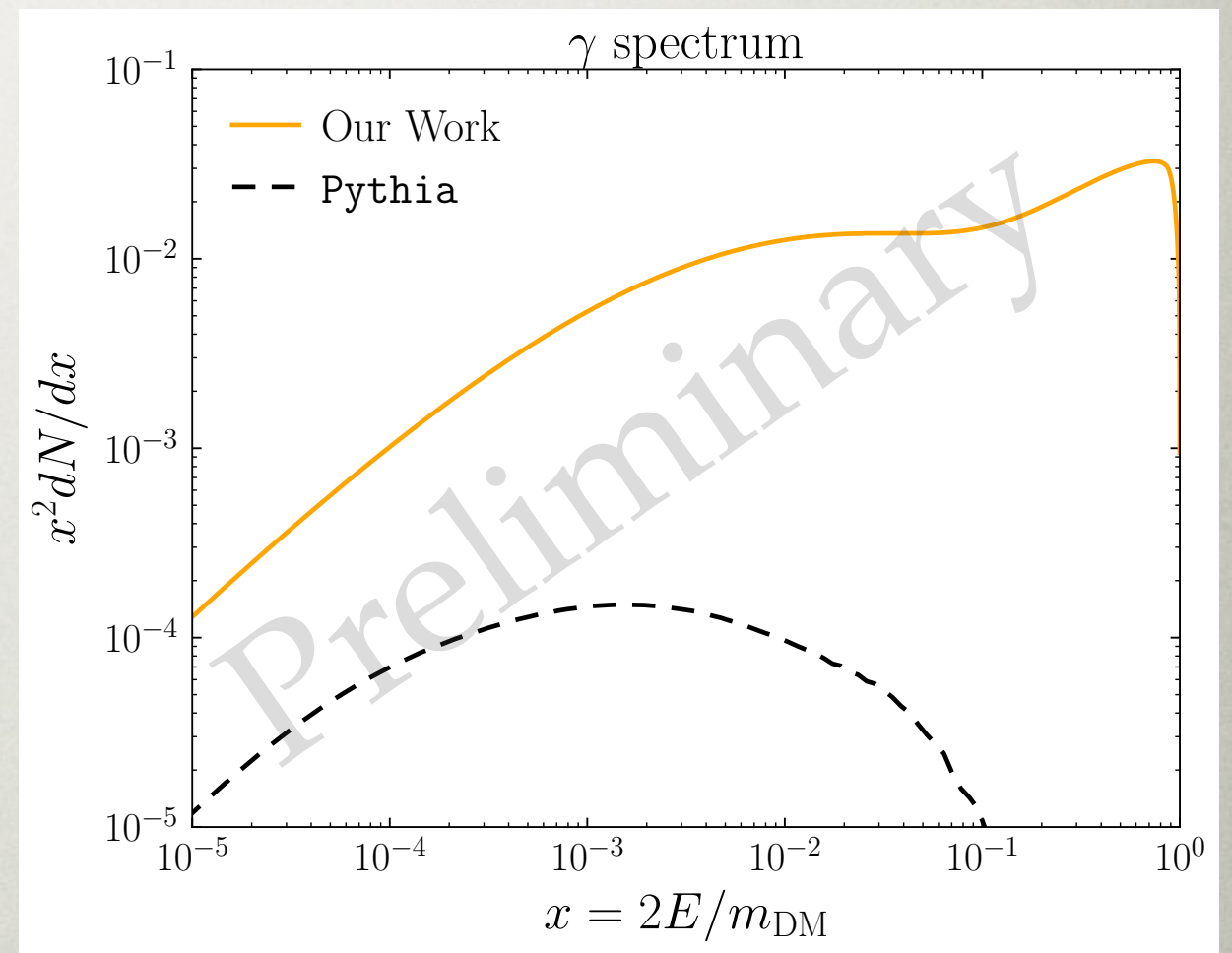
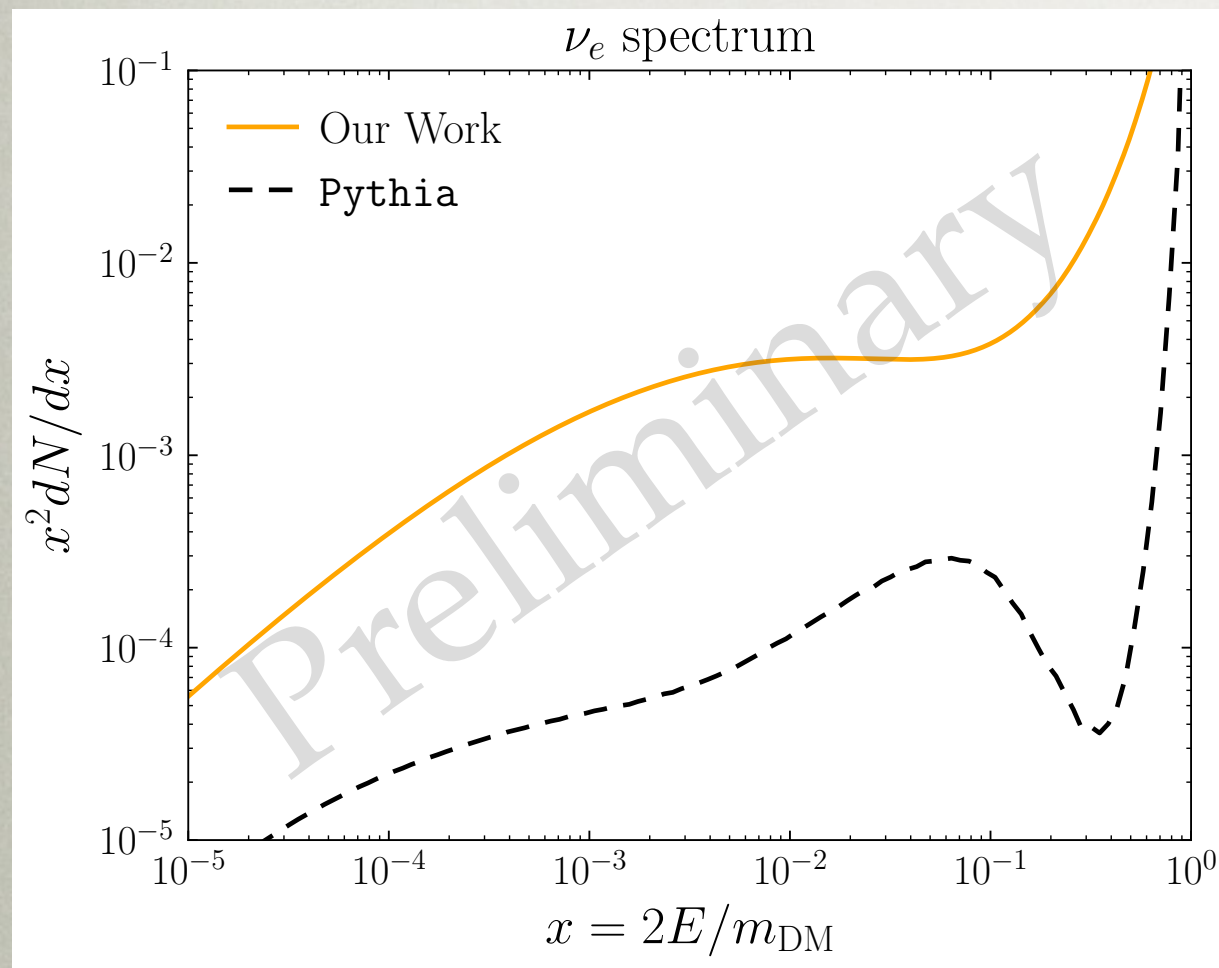


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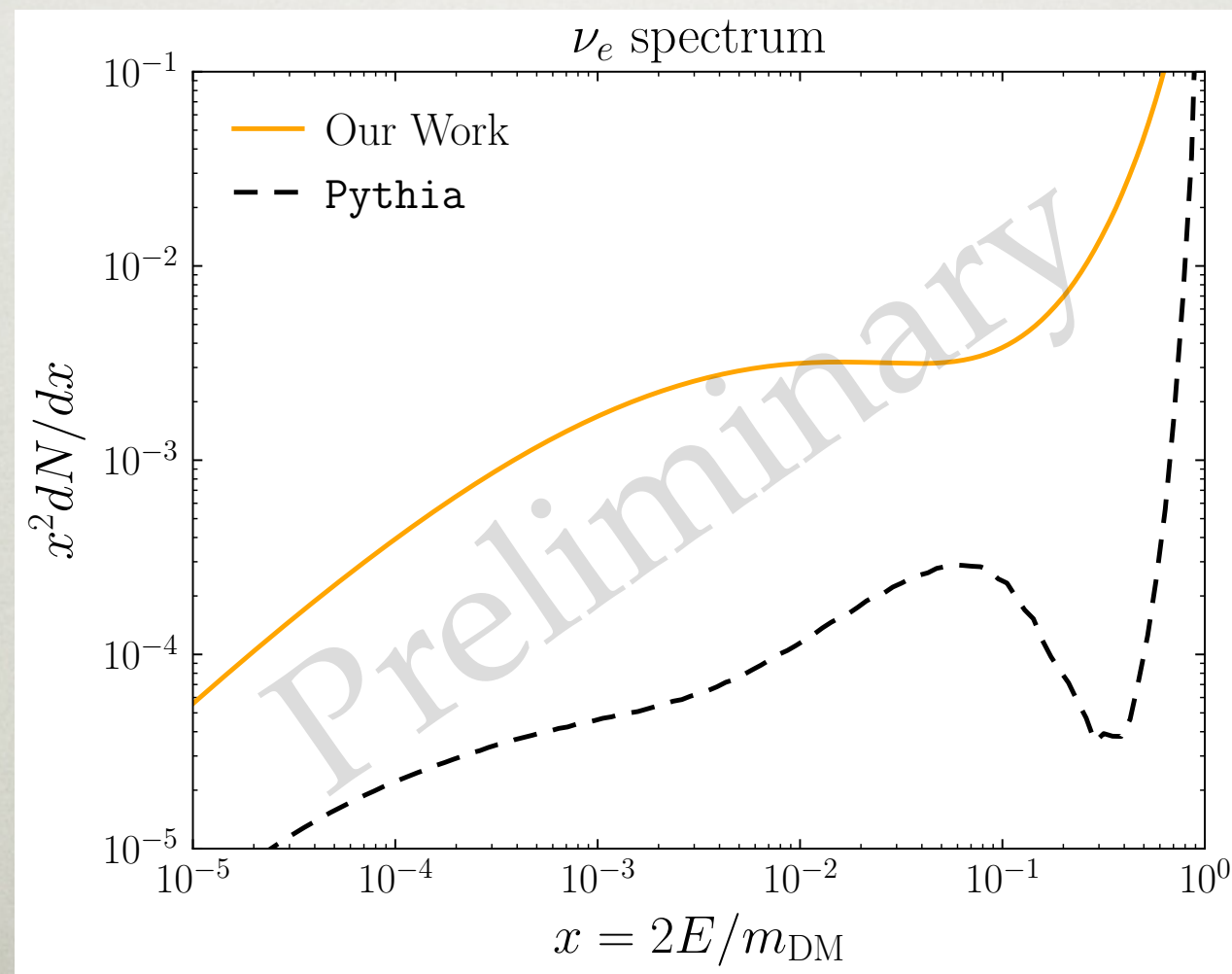
$$\chi \rightarrow \nu_e \bar{\nu}_e$$

$$m_\chi = 2 \text{ EeV}$$



CONCLUSION

- There is a robust experimental program sensitive to HDM decay
- Spectra currently used are known to be wrong
- **Our results represent a significant improvement**
- Spectra will be made publicly available, à la *PPPC4HDMID*
- Future goal: obtain NLL accuracy with soft coherence





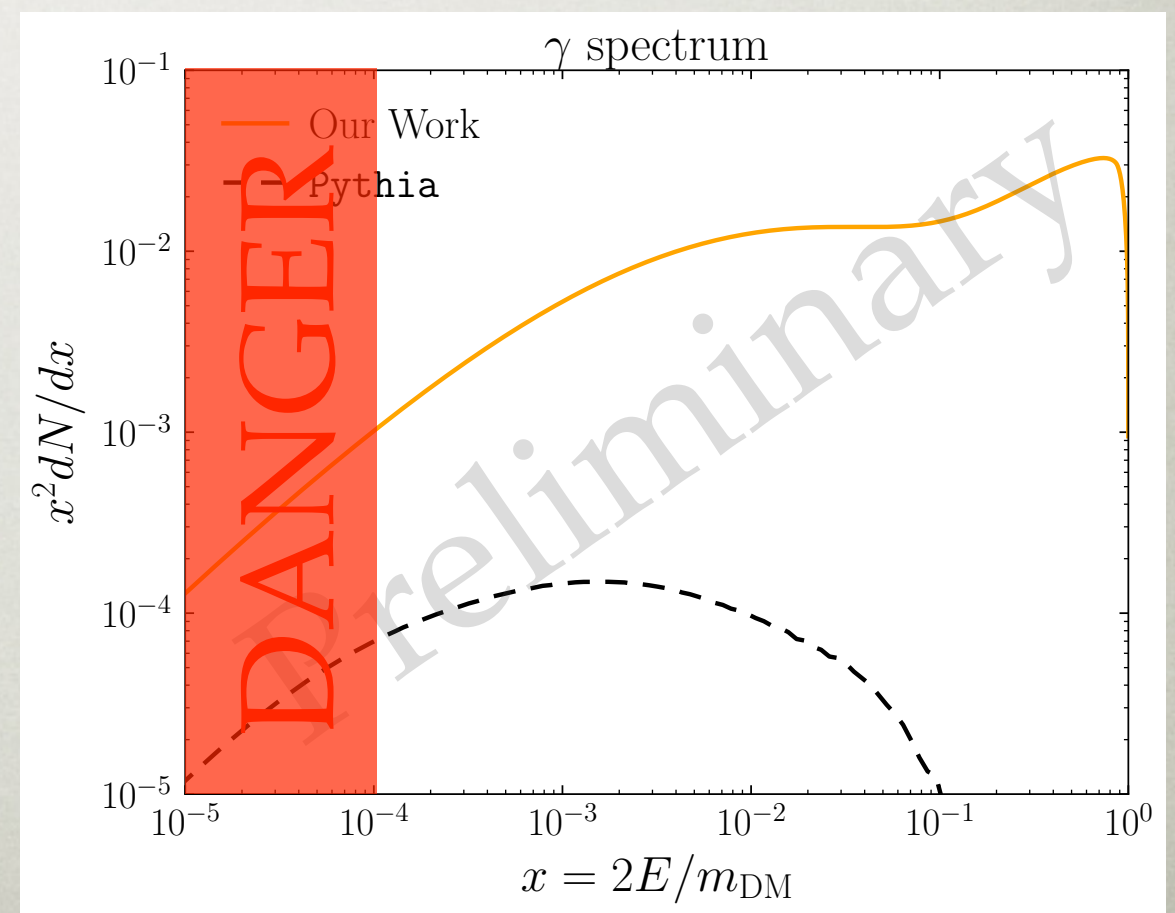
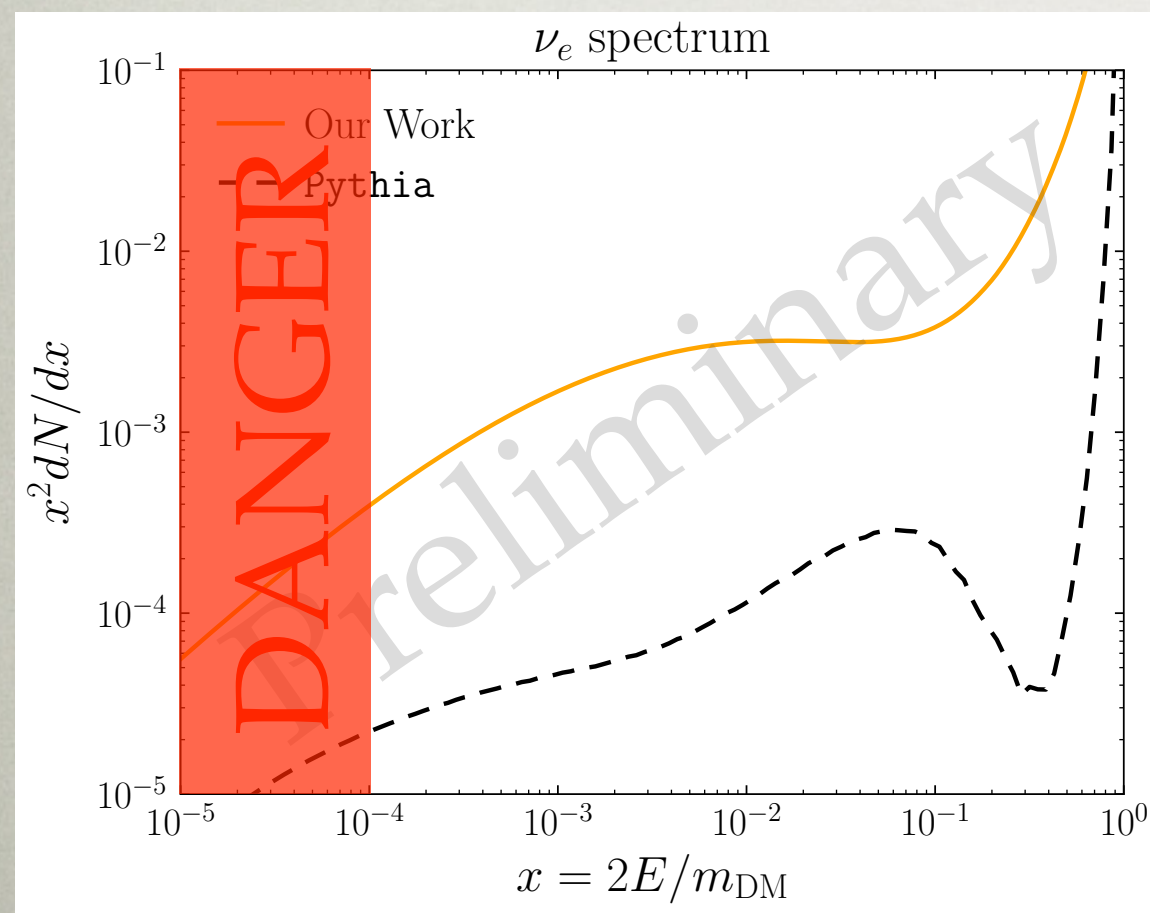
BACKUP SLIDES

BREAKDOWN OF OUR RESULTS

- We are missing a complete treatment of (soft) QCD single logs
- Our result breaks down for $\alpha_s L \sim 1$
- i.e. $x \approx e^{-1/\alpha_s} \approx 10^{-4}$
- Simple argument, only an estimate of where results break

$$\chi \rightarrow \nu_e \bar{\nu}_e$$

$$m_\chi = 2 \text{ EeV}$$

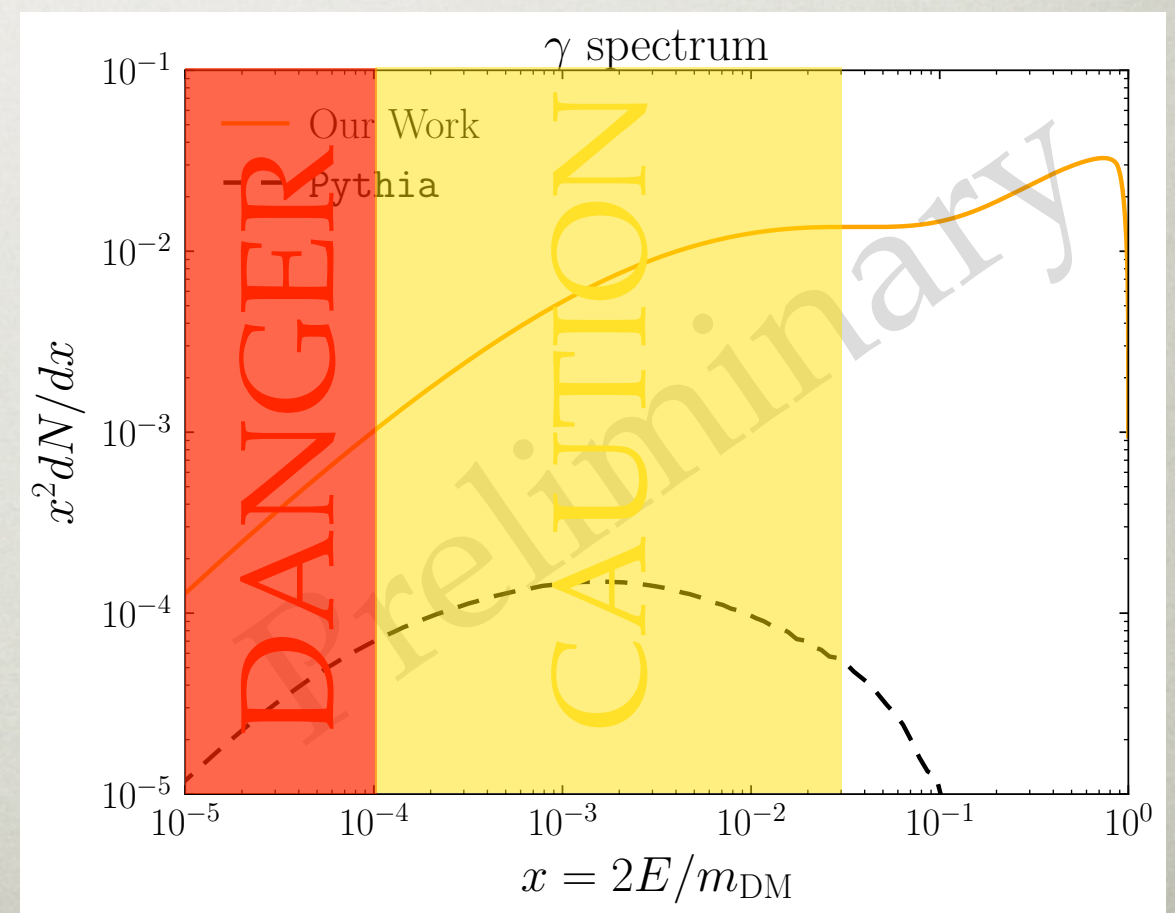
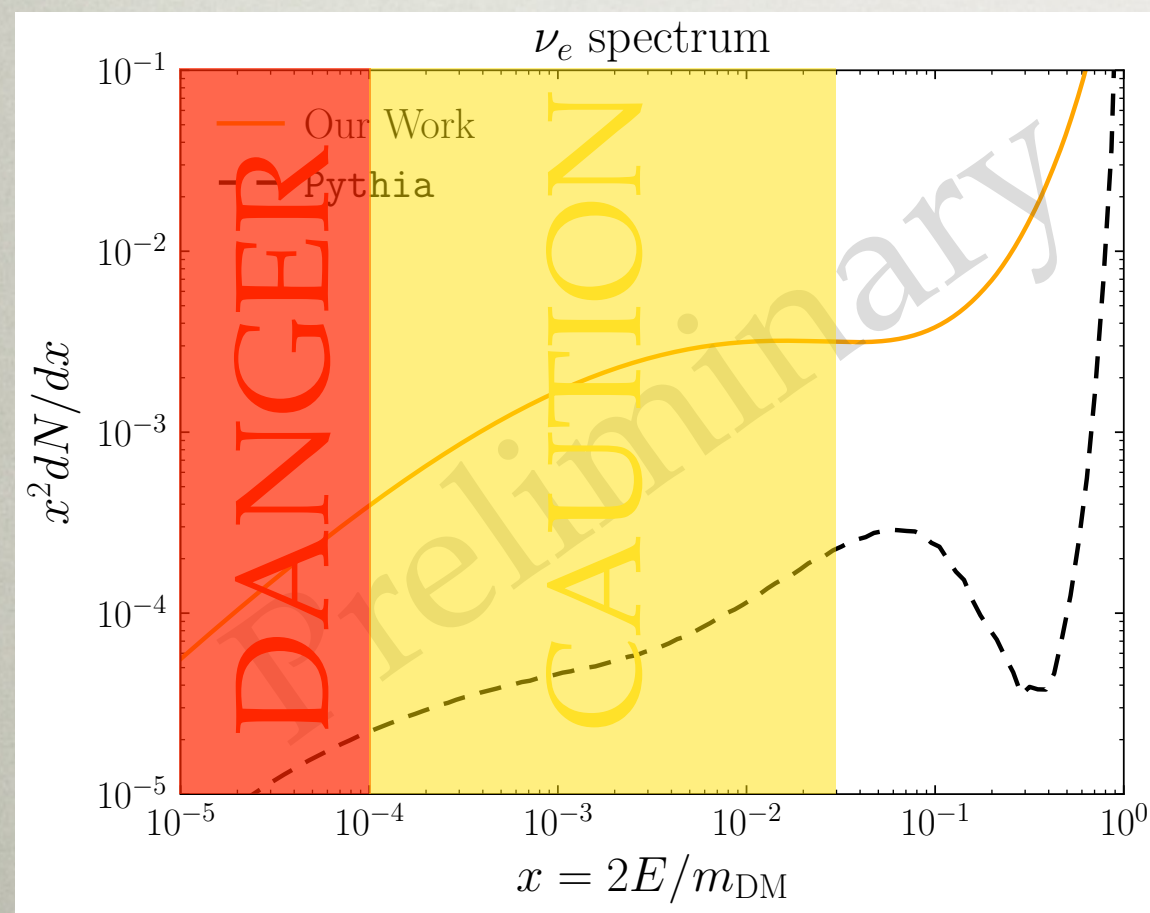


BREAKDOWN OF OUR RESULTS

- Know the expected distribution due to color coherence
[Fong, Webber 1989,1990]
- Important for $3 \ln(1/x) \sim \ln^{3/4}(Q/\Lambda)$
- At EeV energies $x \approx 10^{-1.5}$

$$\chi \rightarrow \nu_e \bar{\nu}_e$$

$$m_\chi = 2 \text{ EeV}$$



PYTHIA SPECTRA VS ENERGY

