



# Sub-GeV Dark Matter and $U(1)_{T3R}$

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2002.01137, 20xx.xxxxx



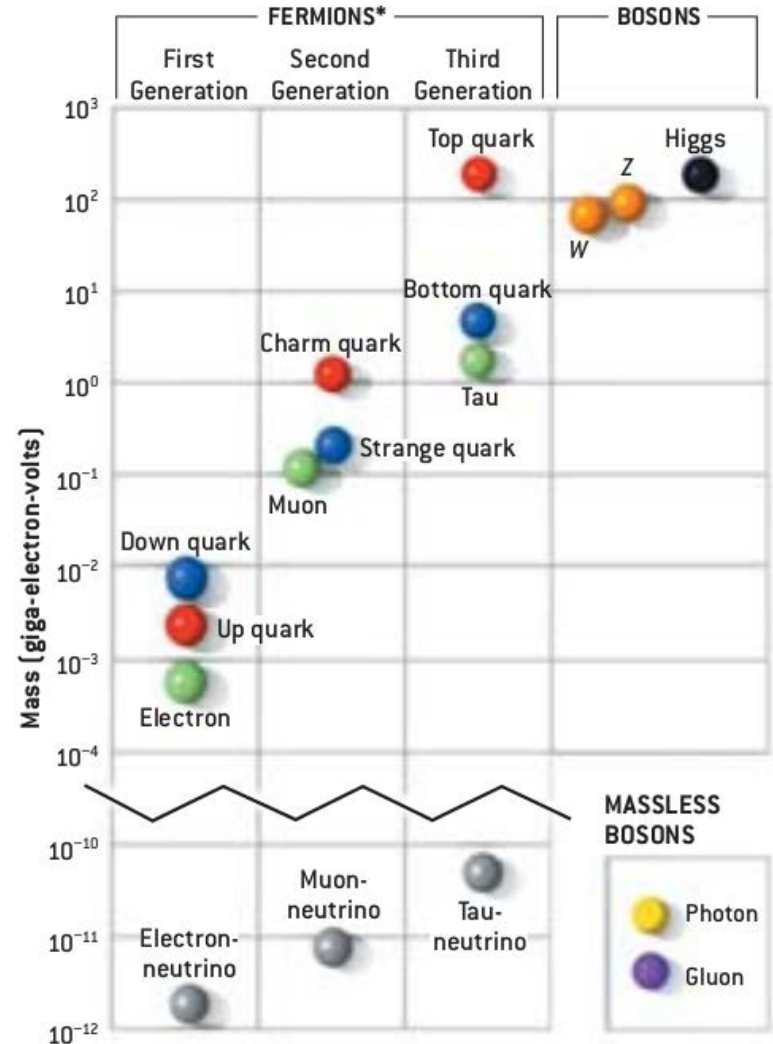
# low-mass dark matter

- there has been recent interest in **sub-GeV dark matter**
  - **evades** tight constraints from current **direct detection** experiments
  - can get the right **relic density** through a variety of mechanisms in which DM is in **thermal contact with SM** (SIMPs, ELDERs)
  - most of all, can be explored by **new**, relatively **inexpensive** experiments
- **but is there any reason for a particle at the MeV scale?**
- analogous to the **WIMP miracle**
  - mechanism for getting TeV-scale particle to have the correct relic density...
  - ... and a reason why a new TeV scale particle should arise (new physics associated with EWSB)
- **reason for having a particle at the MeV scale?...**
- ... **light flavor physics**



# why MeV?

- electroweak scale is a notch on the ruler, and W, Z, h, (WIMPs?) are all around there
- there is another notch on the ruler at MeV scale....
- 1<sup>st</sup> and 2<sup>nd</sup> generation charged particle mass parameters all lie in MeV-GeV range
- can the light flavor sector feed into dark sector?
- our idea → connect DM to light flavor physics through a dark photon/Higgs interaction for right-handed SM fermions



Gordy Kane, Scientific American, May 2003



# new U(1) gauge group

- many scenarios of new physics involve **new U(1) gauge symmetry** under which SM fermions are charged
- since SM is **chiral**, need to make sure U(1) anomalies are **cancelled**
- examples studied recently
  - B-L
  - $L_i - L_j$
  - **secluded U(1)**  $\rightarrow$  SM charges induced at one-loop through **kinetic mixing**
- but we want chiral SM charges, so we use **U(1)<sub>T3R</sub>** (Pati,Salam 74; Mohapatra,Pati 75)
  - couples to **RH fermions**, with up-type and down-type having **opposite** charge
  - originally considered in **left-right models**, where RH fermions are charged under  $SU(2)_R$ , and  $U(1)_{T3R}$  is subgroup generated by diagonal generator ( $\sigma_3$ )
  - descends from  $SU(2)_R$ , so **manifestly anomaly free**
    - anomalies proportional to  $\text{Tr} [\sigma_3]$  and  $\text{Tr}[(\sigma_3)^3] \rightarrow$  vanish
    - won't embed in  $SU(2)_R$



# $U(1)_{T3R}$ and dark matter

- strategy  $\rightarrow$  charge a generation of right-handed SM fermions under  $U(1)_{T3R}$
- in EFT below electroweak scale, the  $U(1)_{T3R}$  protects fermion masses
  - $U(1)_{T3R}$  spontaneously broken down to parity by dark Higgs
  - fermion masses now scale with symmetry-breaking parameter  $V$
- if DM is a fermion also charged under  $U(1)_{T3R}$ , and odd under surviving parity
  - stabilized by parity (it's the only odd particle)
  - gets Majorana mass proportional to symmetry-breaking parameter ( $V$ )
- upshot  $\rightarrow$  two dark sector Majorana fermions with mass scale proportional to  $V$ , just as with SM light fermions
  - lightest is stable (DM), heavier particle may still be around
  - if  $V$  is small, SM fermion scale explained, and DM naturally sub-GeV



# game plan

- this is a **general framework**, but we'll develop an **explicit example**
  - **1<sup>st</sup>/2<sup>nd</sup> generation RH fermions** charged under  $U(1)_{T3R}$
  - dark matter couples to Standard Model particles through a **dark photon** ( $A'$ ) and a **dark Higgs** ( $\phi'$ )
- lots of **constraints**, but **open parameter space available**
- interesting **phenomenological features**...
- ... **spin-independent, velocity-independent** DM-nucleon scattering
  - **elastic scattering** mediated by dark Higgs
  - **inelastic isospin-violating scattering** mediated by dark photon
- get **correct relic density** through (co-)annihilation via intermediate  $A'$  or  $\phi'$
- **consistent with Planck bounds** (invisible, p-wave, or co-annihilation)



# model

- $q_R^u, q_R^d, \ell_R,$  and  $\nu_R \rightarrow Q_{T3R} = \pm 2$ 
  - need not be in same generation
  - anomalies cancel
  - Yukawa terms need  $\phi$  insertion
- $\langle \phi \rangle = V = (-\mu_\phi^2 / 2\lambda_\phi)^{1/2}$ 
  - SM fermion masses  $\propto V$
  - breaks  $U(1)_{T3R}$  to a  $Z_2$  parity
  - SM particles even under parity
  - dark sector fermion  $\eta$  is odd
- new particles
  - $A'$  (dark photon),  $\phi'$  (dark Higgs)
  - $\nu_S$  (mostly  $\nu_R$ )
  - $\eta_{1,2}$  (Majorana fermion DM)

charges of left-handed component of Weyl spinor

field	$q_R^u$	$q_R^d$	$\ell_R$	$\nu_R$	$\eta_L$	$\eta_R$	$\phi$
$q_{T3R}$	-2	+2	+2	-2	1	-1	-2

$$\begin{aligned}
 L_\phi = & -\frac{\lambda_u}{\Lambda} \tilde{H} \phi^* \bar{Q}_L q_R^u - \frac{\lambda_d}{\Lambda} H \phi \bar{Q}_L q_R^d \\
 & -\frac{\lambda_\nu}{\Lambda} \tilde{H} \phi^* \bar{L}_L \nu_R - \frac{\lambda_\ell}{\Lambda} H \phi \bar{L}_L \ell_R \\
 & -m_D \bar{\eta}_R \eta_L - \frac{1}{2} \lambda_L \phi \bar{\eta}_L^c \eta_L - \frac{1}{2} \lambda_R \phi^* \bar{\eta}_R^c \eta_R \\
 & -\mu_\phi^2 \phi^* \phi - \lambda_\phi (\phi^* \phi)^2 + \text{h.c.}
 \end{aligned}$$

$$\tilde{H} \equiv i\sigma_2 H^*, \text{ and we take } \lambda_L = \lambda_R \equiv \lambda_M$$



# masses and couplings

- EFT below EWSB scale....
  - $\phi'ff \rightarrow$  coupling  $\propto m_f / V$
  - $A'ff \rightarrow$  coupling  $\propto Q_f m_{A'} / V$
- $\eta$  has **Maj.** and **Dirac** mass terms
  - take  $m_D \ll \lambda_M V$
  - $m_{1,2} \propto V$ , with small splitting
  - SM and DM masses scale with  $V$
  - if  $V \sim 1-10$  GeV, naturally get sub-GeV SM and DM fermions, as well as sub-GeV  $A'$ ,  $\phi'$
- $A'$  coupling to  $\eta_{1,2}$  is **off-diagonal**
  - **inelastic scattering, co-annih.**
- $A'$  kinetically mixes with  $\gamma$ ,  $Z$

$$M_\eta = \begin{bmatrix} \lambda_M V & m_D \\ m_D & \lambda_M V \end{bmatrix}$$

$$M_\nu = \begin{bmatrix} 0 & \lambda_\nu V \\ \lambda_\nu V & M \end{bmatrix}$$

$$m_{A'} = \sqrt{2} g_{T3R} V, \quad m_{\phi'} = 2\lambda_\phi^{1/2} V$$

$$j_{T3R}^\mu = \frac{i}{2} (\bar{\eta}_1 \gamma^\mu \eta_2 - \bar{\eta}_2 \gamma^\mu \eta_1) + j_{T3R, \phi'}^\mu + j_{T3R, SM(R)}^\mu$$





# setup

- we can choose  $q_R^d, \ell_R$  to be **mass eigenstates**, since this is **technically natural** (extra  $U(1)^2$  flavor symmetry)
  - see Batell, Freitas, Ismail, McKeen (1712.10022)
- no symmetry reason to assume  $q_R^u$  a mass eigenstate, but we'll assume for simplicity that dominant coupling is to **one mass eigenstate**
- so we take the SM fermions charged under  $U(1)_{T3R}$  to be  $\mu_R, u_R$ , and  $d_R$ 
  - flavor diagonal
  - other choices possible, but we'll pick this for simplicity and phenomenology
- assuming perturbativity, we get  $m_{SM}, m_{1,2}, m_{A'}, m_{\phi'} < V$
- smaller  $V \rightarrow$  **lighter DM**, with **stronger coupling to SM**
- taking  $V \sim \text{GeV}$  would give us  $\mathcal{O}(1)$  couplings, but in **tension with data**
- **need some modest hierarchies**



# constraints

Batell, Freitas, Ismail, McKeen (1712.10022); Bauer, Foldenauer, Jaeckel (1803.05466)

- but lots of **constraints** on  $A'$ ,  $\phi'$  coupling to SM fermions
- main **differences** between our scenario and others
  - no coupling to  $\bar{\nu}_L \nu_L$  ( $\nu_R/\nu_A$  mixing taken small), affects **v exp./cooling constraints**
  - direct coupling to  $\mu$ , not  $e$  (1-loop)  $\rightarrow$  affects some  **$e^+e^-$  collider constraints**
  - chiral coupling of  $A'$  to fermions  $\rightarrow$  even at **weak coupling, Goldstone couples**
- **g-2 corrections** from  $\phi'$  (positive) and  $A'$  (negative) running in loop
  - corrections can be **fine-tuned** against each other or **heavy new physics**
- main constraints
  - **$e^+e^- \rightarrow 4\mu$  (BaBar)**
  - **fixed target/beam dump exps.:**  $A', \phi' \rightarrow \gamma\gamma, e^+e^-$  (E137, E774, E141, Orsay, etc.)
  - **solar/SN/Glob. Cluster cooling constraints** (production of  $A', \phi'$ )
  - **fifth force constraints/ $N_{\text{eff}}$**
- we'll take  **$V = 10$  GeV**, and will find restrictions on  $m_{\phi'}$  and  $m_{A'}$ 
  - not much dependence on dark matter mass (take  $m_{\eta}, m_{\nu_S} > 10$  MeV)



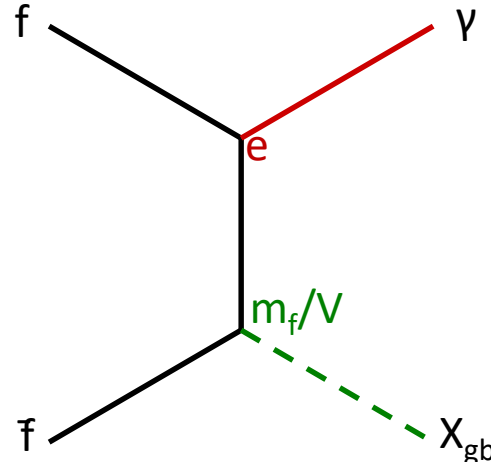
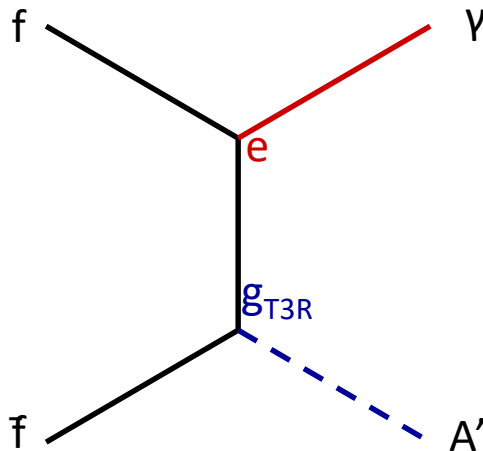
# $N_{\text{eff}}$ and $U(1)_{T3R}$ (2002.00137)

- generally **two ways** to avoid light  $A'$  or  $\phi'$  contributing too much to  $N_{\text{eff}}$
- if  $A'$  and  $\phi'$  are **heavy enough** ( $> 10$  MeV), they are gone before neutrino decoupling and don't affect  $N_{\text{eff}}$
- if coupling is **weak enough**, then  $A'$  and  $\phi'$  are never in equilibrium with SM  $\rightarrow$  never produced, so also don't affect  $N_{\text{eff}}$
- for our case,  $U(1)_{T3R}$  coupled to second-generation
  - for  $\phi'$ , coupling  $m_f/V \sim 0.01$ , so **never weakly coupled** enough
  - for  $A'$ , coupling  $m_{A'}/V$ , so can make weakly coupled just by making it **light**
- but  **$U(1)_{T3R}$  case is very different** from B-L,  $L_i$ - $L_j$ , kinetic mixing, etc.
  - no matter how weak the coupling, **always produced in the early Universe** unless  **$V > \mathcal{O}(10^7)$  GeV**
  - result of coupling to **chiral fermions**



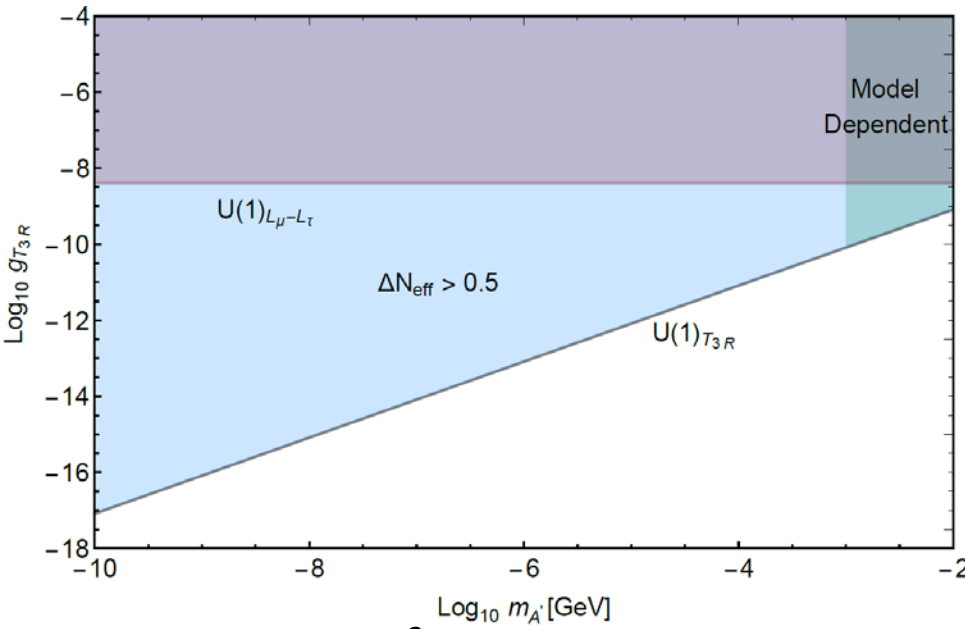
# $N_{\text{eff}}$ and chiral fermions

- weak coupling, so dominant  $A'$  production mode is **inverse decay** process
  - $ff \rightarrow \gamma A'$
- longitudinal modes get an enhancement,  $E/m_{A'}$ , so  **$A'$  thermalizes regardless of how small the mass/coupling is**
  - **enhancement killed if there is only a vector coupling**, due to Ward identity
- another way to see it... as  $m_{A'}/V \rightarrow 0$ ,  $U(1)_{T3R}$  becomes a **global symmetry**
  - massless **Goldstone mode couples as  $m_f/V$** , always thermalizes
  - for B-L,  $L_i-L_j$ , etc., ... no need for Goldstone to couple of charged SM fermions
- we'll consider case where 2<sup>nd</sup> generation couples to  $U(1)_{T3R}$





# comparing $U(1)_{T_{3R}}$ to $L_{\mu} - L_{\tau}$



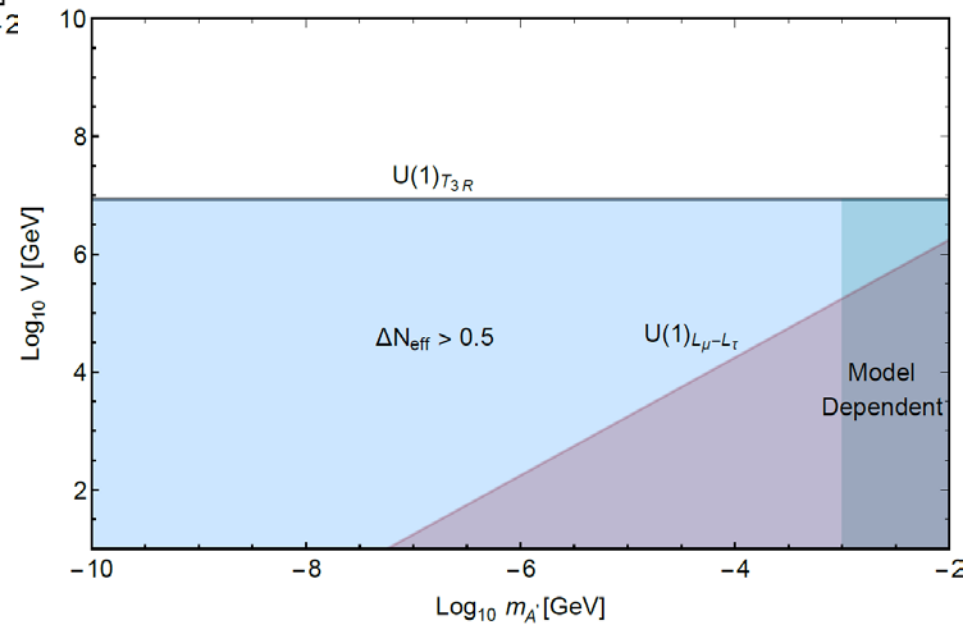
$$\sigma v \propto \frac{\alpha_{\text{em}} m_f^2}{sV^2}$$

$$\langle \sigma v \rangle_{T=m_f} \eta > H$$

just outside excluded region, can get small  $\Delta N_{\text{eff}}$ , which might help with  $H_0$  tension (Bernal, Verde, Riess 1604.0617; Escudero, Hooper, Krnjaic, Pierre, 1901.02010, for example)

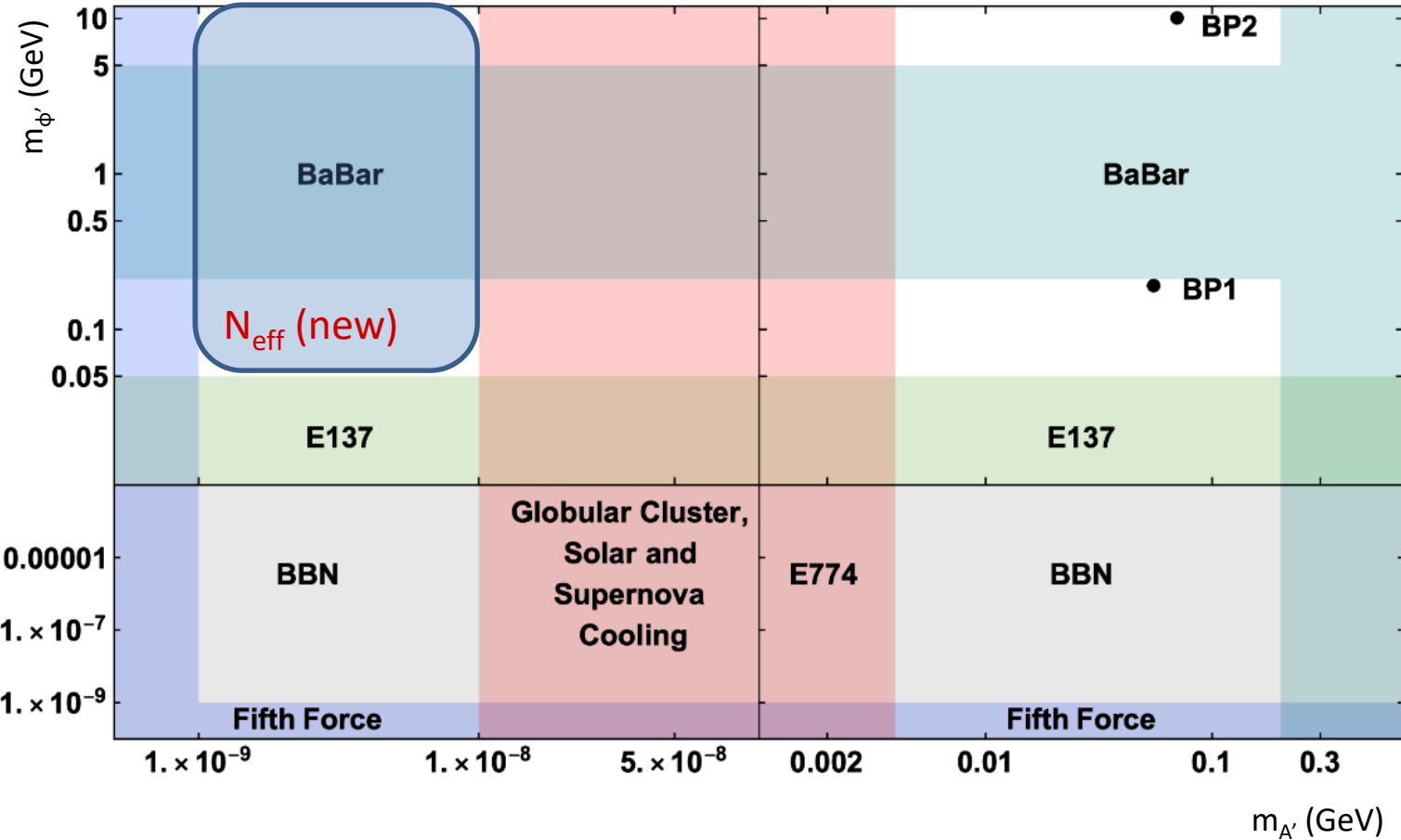
$L_{\mu} - L_{\tau}$  is from Escudero, Hooper, Krnjaic, Pierre (1901.020210, purple)

between 1-10 MeV,  $A'$  can decay to  $\nu_A \nu_A$  or  $e^+ e^-$  at one-loop, so effect on  $\Delta N_{\text{eff}}$  depends on neutrino mass matrix





# constraints



two benchmark points which we will see get the relic density right also

open parameter space limited, can be closed with future experiments!

but parameter space can also be opened up by increasing  $V_{\text{eff}}$ ...



# direct detection

- $\phi'$  mediated  $\rightarrow$  SI, velocity-independent, elastic, isospin-invariant
- $A'$  mediated  $\rightarrow$  SI, velocity-independent, **inelastic**, **isospin-violating** (IVDM)
  - opposite coupling to u and d (thus to p and n)
- mediator mass is of the same order as momentum transfer
  - **not a contact interaction**,  $d\sigma/dE_R$  suppressed by  $[1+(2m_A E_R/m_{\phi',A'})^2]^{-2}$
- current constraints (contact interaction, isospin-invariant)
  - **CRESST III**  $\rightarrow \sigma_{SI} \sim 10^{-35} \text{ cm}^2$  at  $m_\eta = 200 \text{ MeV}$
  - **CDEX-1B**  $\rightarrow \sigma_{SI} \sim 10^{-32-34} \text{ cm}^2$  at  $m_\eta = 50-180 \text{ MeV}$
  - **XENON1T**  $\rightarrow \sigma_{SI} \sim 10^{-29-30} \text{ cm}^2$  over full mass range,  $\sigma_{SI} \sim 10^{-34} \text{ cm}^2$  at  $m_\eta = 100 \text{ MeV}$  (Migdal effect, 1907.12771)
    - **cosmic rays scatter off DM**, boosted DM scatters in XENON1T (Bringmann, Pospelov -1810.10543; Dent, Dutta, Newstead, Shoemaker – 1907.03782)
- **benchmark models satisfy all bounds** (including electron scattering)



# thermal relic density

- main relevant annihilation channels are **s-channel** through  $\phi'$  or  $A'$
- a thermal relic cross section would naively violate **Planck bounds**
- **three ways out** which we can use
  - **p-wave**: factor 10 suppression at freeze-out, but much more at recombination
    - kills Planck bounds for  $\phi'$ -mediated case
  - **co-annihilation**: heavier state around at freeze-out, but decayed before recombination
    - can rescue  $A'$ -mediated co-annihilation case, if DM splitting is set right
  - **invisible final states**: if annihilation products predominantly produce  $\nu_{s,A}$
- $\phi'$  coupling suppressed by mass of incoming/outgoing particles
  - need to be **near resonance** to get correct relic density for  $\phi'$  mediator
- $A'$  coupling not suppressed if  $A'$  is not light..., need not be on resonance
  - final state is  $\nu_{s,A} \nu_{s,A}$ , or demand  $\eta_2$  **decay before recombination**





# two benchmark models

$\ell=\mu$	$m_{A'}$ (MeV)	$m_{\phi'}$ (MeV)	$m_{\eta}$ (MeV)	$m_{\nu_s}$ (MeV)	$m_{\nu_D}$ (MeV)	$\langle\sigma v\rangle$ (cm <sup>3</sup> /s)	$\sigma_{\text{SI}}^S$ (pb)	$\sigma_{\text{SI}}^V$ (pb)
BP1	55	200	100	10	10 <sup>-3</sup>	3 × 10 <sup>-26</sup>	0.51	6.50
BP2	70	10 <sup>4</sup>	50	10	10 <sup>-3</sup>	3 × 10 <sup>-26</sup>	5.7×10 <sup>-9</sup>	1.80

- first benchmark get relic density via  $\phi'$  resonance
  - $a_\mu$  corrections ( $A'/\phi'$ ) need to be **fine-tuned** against each other to **1%**
  - need a contribution from new physics for cancellation
- for second benchmark, get relic density from **co-annihilation via  $A'$** 
  - $\phi'$  corrections to  $a_\mu$  small, so need to cancel  $\delta a_\mu$  correction from  $A'$  against heavy new physics to **1%**
  - $e^+e^-$  rate (one-loop) can be non-negligible, but can suppress if **heavier state gone before recombination**
    - if splitting small enough ( $< 1$  MeV), **doesn't affect BBN**



# upshot

- **sub-GeV dark matter** is a target which experiments are focusing on....
- points to either high-scale new physics with a suppressed coupling to DM, or **low-scale new physics with less suppressed couplings**
- best-case scenario is a **GeV scale dynamically-generated parameter** from new physics coupled to DM and SM
  - natural SM coupling is the **light-flavor sector**
- but the very best-case scenario is in **tension with data...**
- ... need to push the parameter scale up, and the couplings down, to avoid tight constraints → **need some fine-tuning** ( $V, g-2$ )
- but **points to a window** where we get the **correct relic density**, and have interesting **future prospects for experiments**
- **inelastic scattering is a generic feature** whenever DM is charged under a **broken continuous symmetry** (mediated by dark gauge boson)

# conclusion

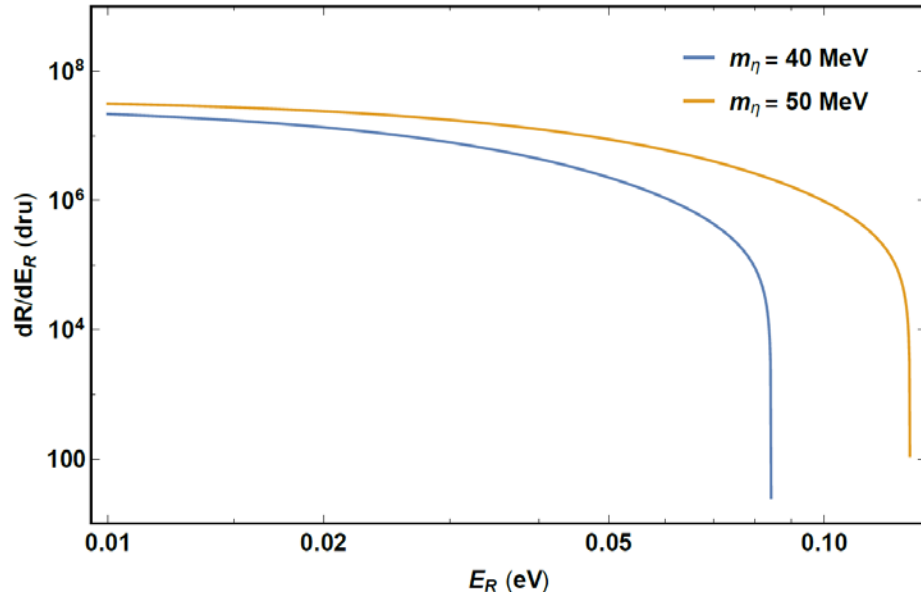
- **dark matter** experiments are set to probe the **MeV-scale**, but still need theory guidance
  - MeV scale naturally arises in models which **connect dark sector to the light flavor sector**
  - **many constraints** narrow parameter space, but **some room left**
- 
- implications for **MeV-range searches**, including inelastic scat. of IVDM
  - **future work** → prospects for future experiments to definitively test, coupling to electrons while avoiding atomic parity violation exps.



# Backup Slides

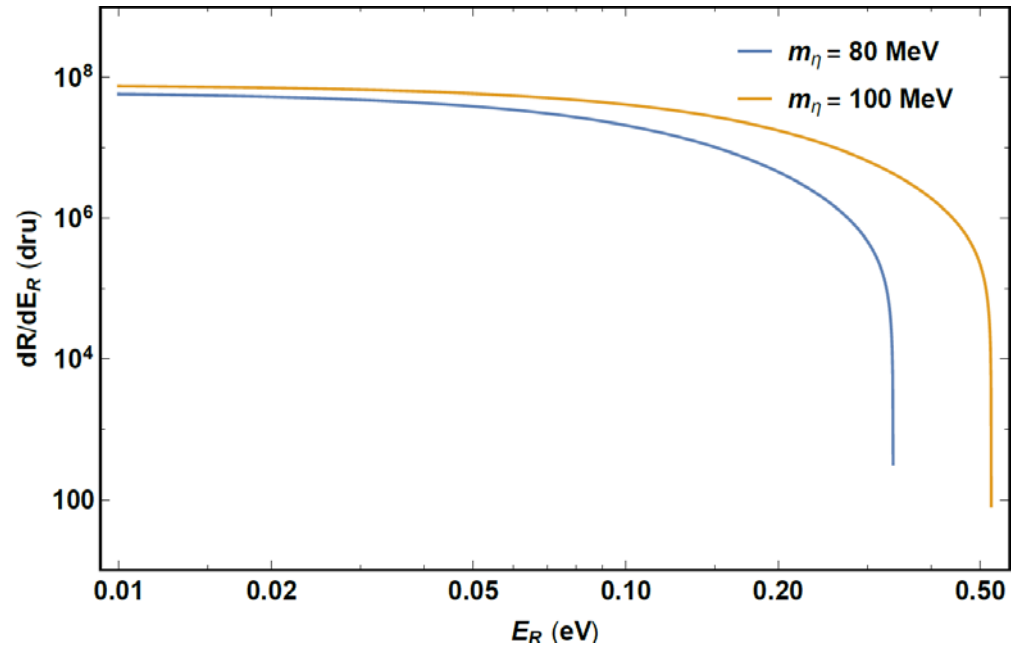


# elastic scattering rates



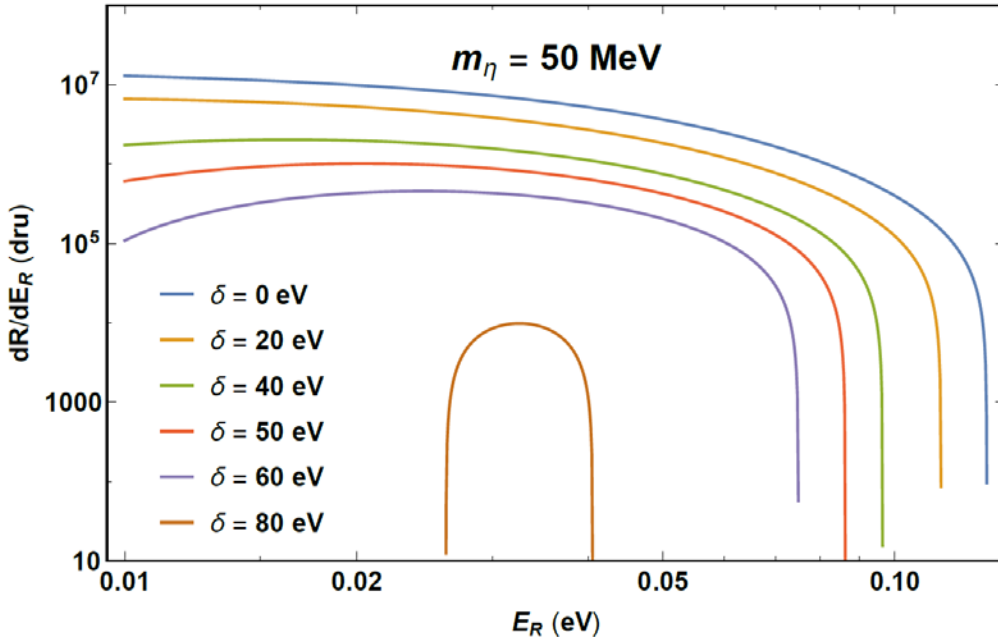
$$m_{\phi'} = 200 \text{ MeV}, V = 10 \text{ GeV}$$

differential cross section suppressed  
by  $[1 + (2m_A E_R / m_{\phi', A'}^2)]^{-2}$



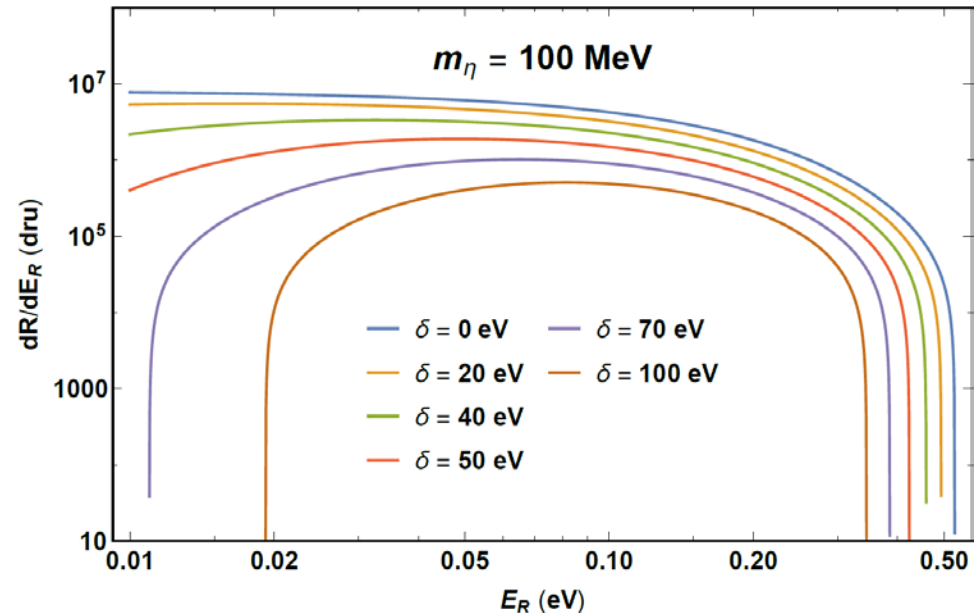


# inelastic scattering rates



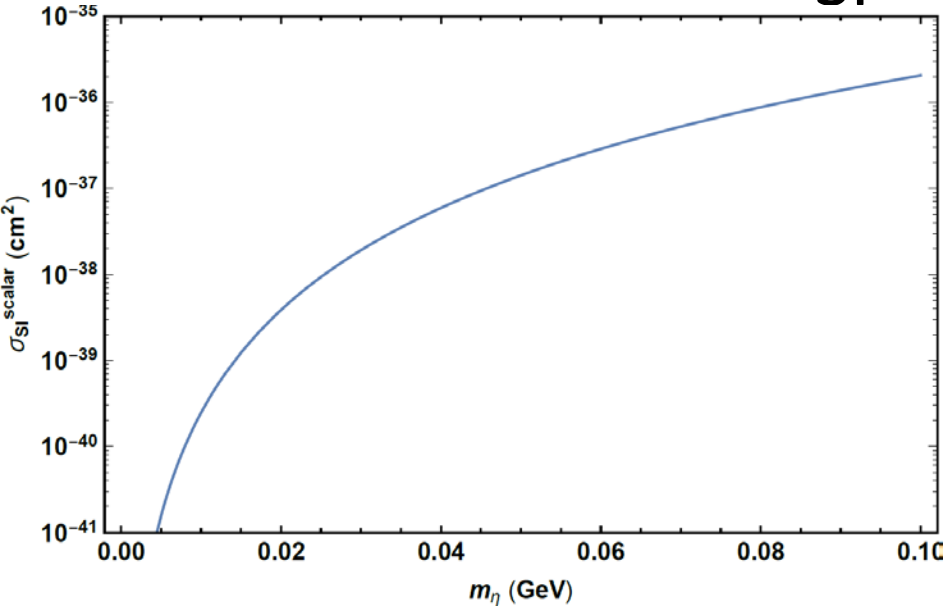
$$m_{A'} = 55 \text{ MeV}, V = 10 \text{ GeV}$$

differential cross section suppressed by  $[1 + (2m_{A'} E_R / m_{\phi', A'}^2)]^{-2}$





# DM-nucleon $\sigma_{SI}$ (zero-mom. trans.)



actual differential cross section suppressed by  $[1+(2m_A E_R/m_{\phi',A'}^2)]^{-2}$

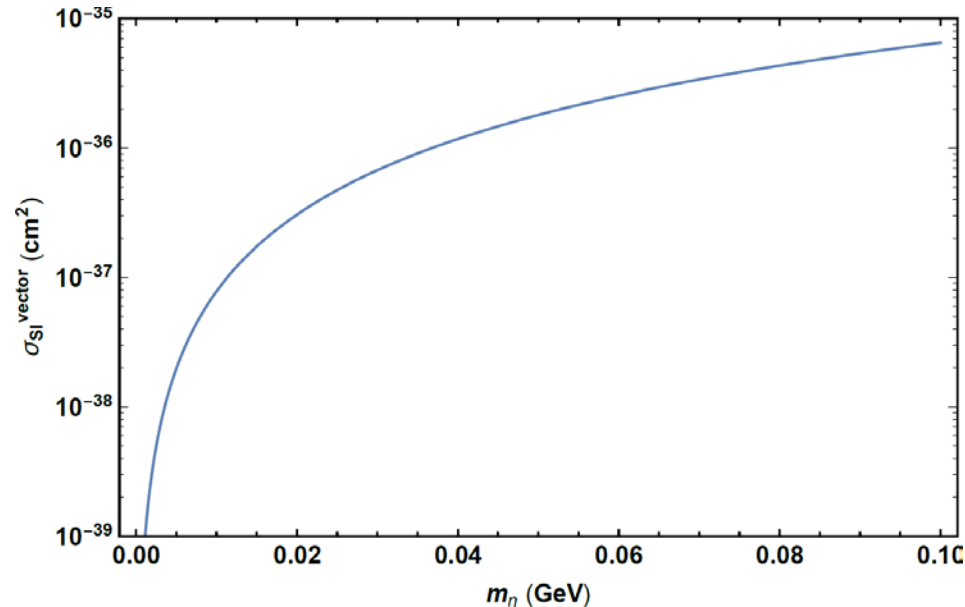
$$\sigma_{SI(0)}^{\text{vector}(p,n)} = \frac{\mu_{\eta N}^2}{16\pi V^4} \delta = 0$$

$V=10 \text{ GeV}$

$m_{\phi'} = 200 \text{ MeV}, V=10 \text{ GeV}$

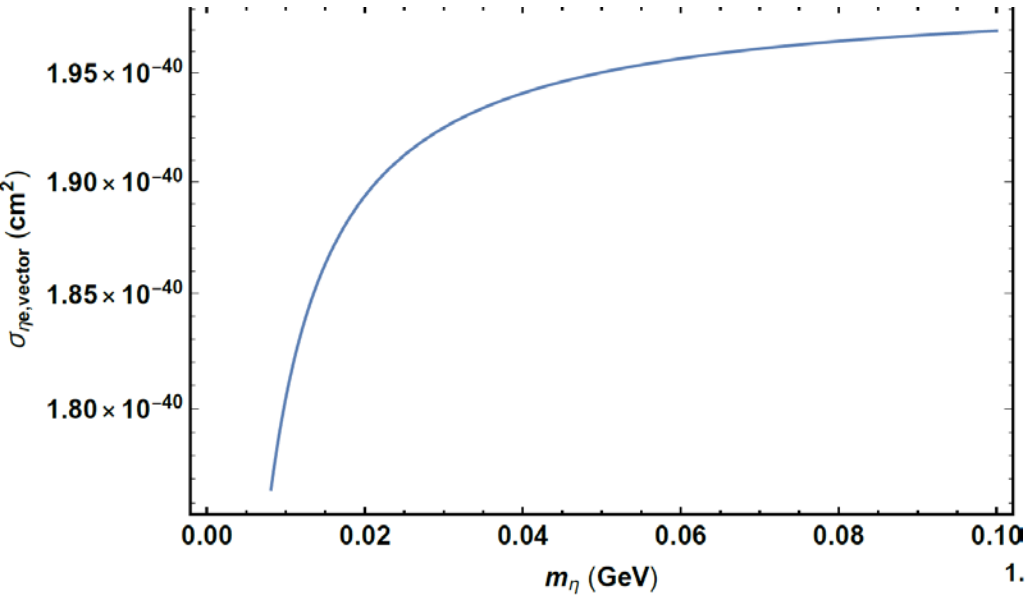
$$\sigma_{SI(0)}^{\text{scalar}(p,n)} = \frac{\mu_{\eta N}^2 m_\eta^2}{4\pi V^4 m_{\phi'}^4} f_{p,n}^2$$

$$f_{p,n} \propto m_N$$

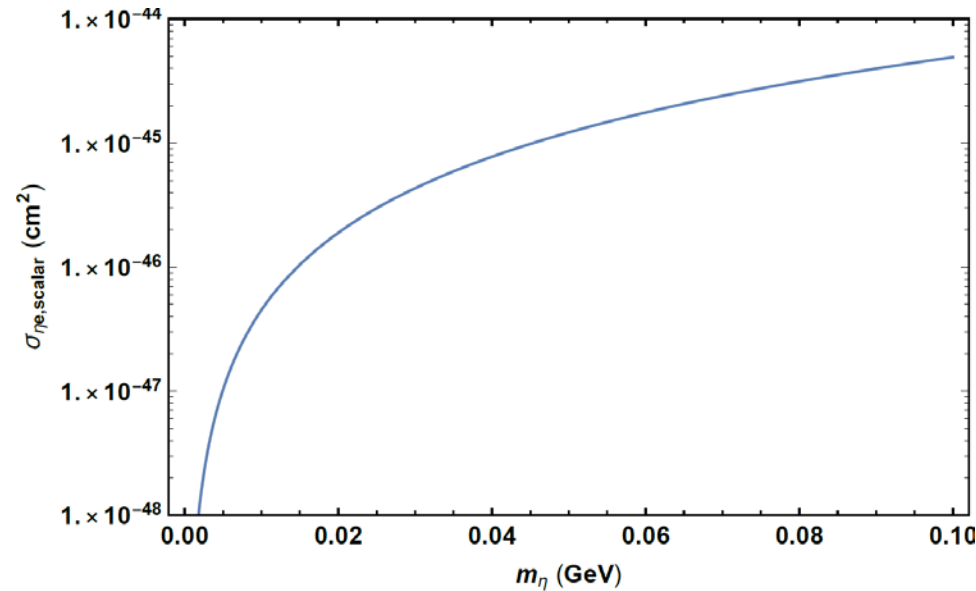




# DM-electron cross sections



$m_{\phi'} = 200 \text{ MeV}, \delta = 0,$   
 $V = 10 \text{ GeV}$

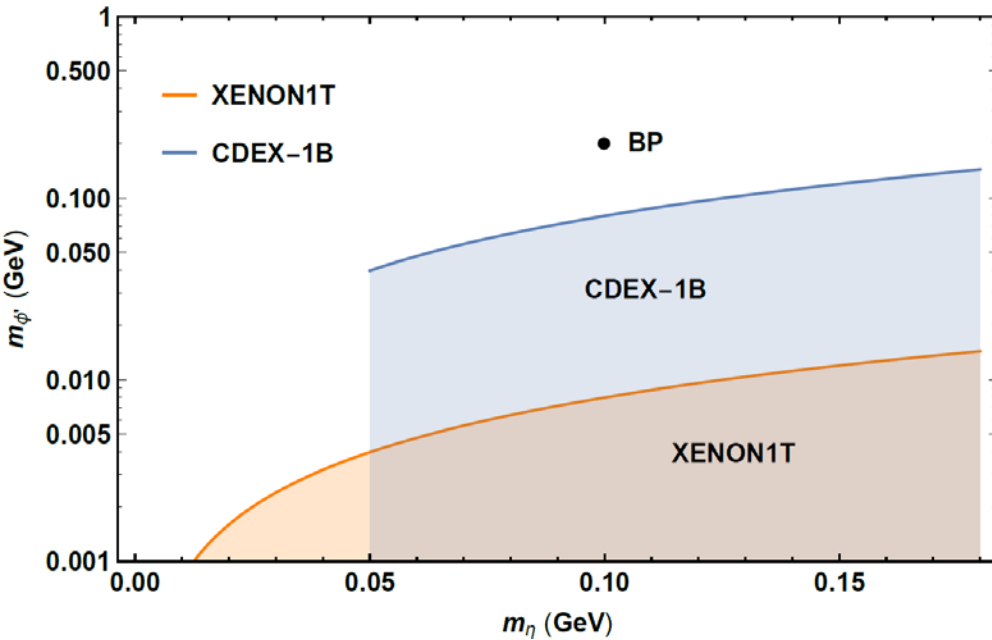


at zero mass splitting and momentum transfer,  $A'$ -mediated scattering gets close to recent XENON1T bound ( $2E-40 \text{ cm}^2$ , 1907.11485) but suppression from mass splitting can ease tension





# exclusion contours

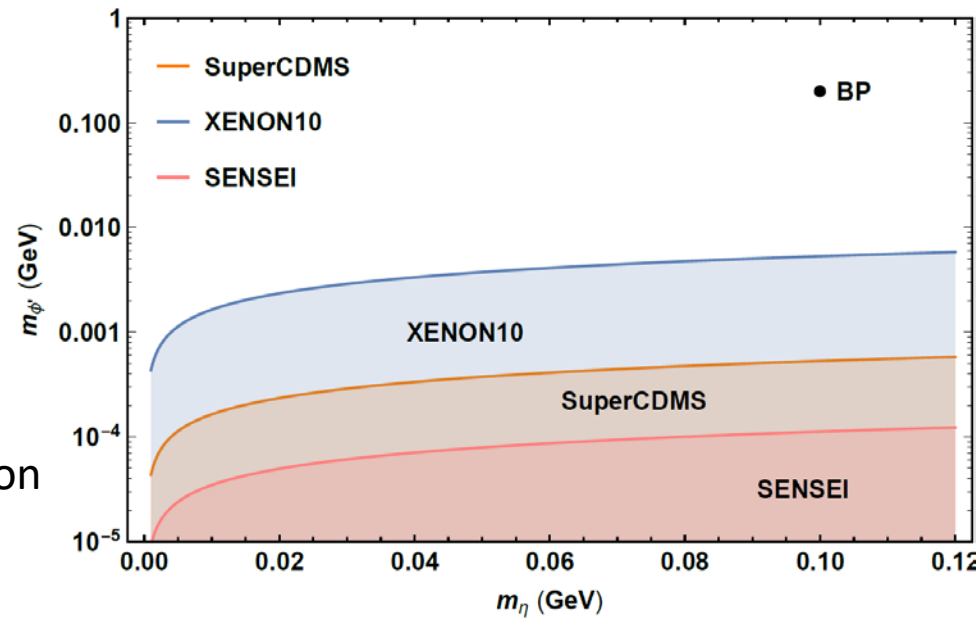


elastic scattering ( $\phi'$ -mediated)

BP:  $m_{\phi'}=200$  MeV,  $m_{\eta}=100$  MeV,  
 $V=10$  GeV

DM-nucleon

DM-electron





# g-2 correction

- correction from  $\phi'$  is positive, but correction from  $A'$  is negative
  - vector + axial
- as  $m_{A'} \rightarrow 0$ , coupling goes to zero and transverse polarizations decouple, but longitudinal polarization does not
  - becomes massless Goldstone mode of a global symmetry
  - g-2 correction becomes that of pseudoscalar with Goldstone's coupling
- all corrections go away as  $m_\ell \ll m_{A'}$

$$\delta a_\ell = \frac{m_\ell^4}{16\pi^2 V^2} \int_0^1 \frac{(1-x)^2(1+x)}{(1-x)^2 m_\ell^2 + x m_{\phi'}^2} + \frac{m_\ell^2}{32\pi^2 V^2} \int_0^1 \frac{2x(1-x)(x-2)m_{A'}^2 - 2x^3 m_\ell^2}{x^2 m_\ell^2 + (1-x)m_{A'}^2}$$



# constraint considerations

- $\ell=\mu$ 
  - $\phi'^{(*)} \rightarrow \gamma\gamma$  through a  $\mu$  loop is always open, dominates if  $v_s$  heavy enough
    - kills cooling bounds through off-shell  $\phi'$ , gives beam-dump bounds
  - $A' \rightarrow \gamma\gamma$  forbidden by Landau-Yang theorem
    - cooling through  $A'$  has to be killed by heavy  $A'$ , weak coupling, or suppressed by heavy  $v_s$  (coupling to  $v_A$  is one-loop)
    - $A' \rightarrow e^+e^-$  proceeds through one loop kinetic mixing, but subdominant to  $v_A v_A$ 
      - $A' \rightarrow v_A v_A$  allowed because of  $\gamma^5$  coupling
      - gives beam dump bounds
- $\ell=e$ 
  - $\phi'^{(*)} \rightarrow e^+e^-$  tree-level, but suppressed by small coupling, beam dump bounds
  - $\phi'^{(*)} \rightarrow \gamma\gamma$  kills cooling bounds if  $v_s$  is heavy enough to suppress invis. decay
  - $A' \rightarrow e^+e^-$  at tree-level gives beam dump bounds
  - if  $A'$  light enough, get cooling bounds from  $A' \rightarrow v_A v_A$



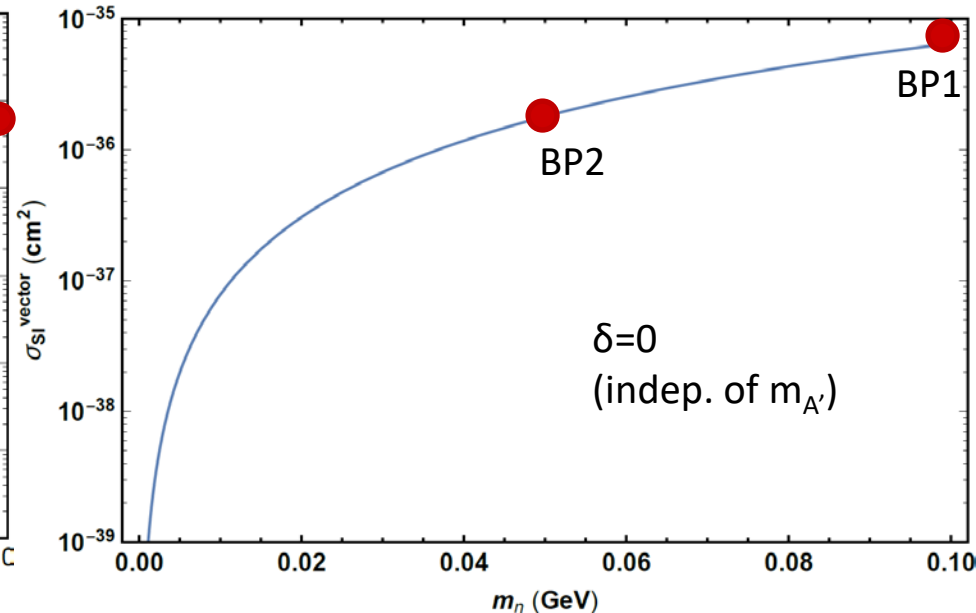
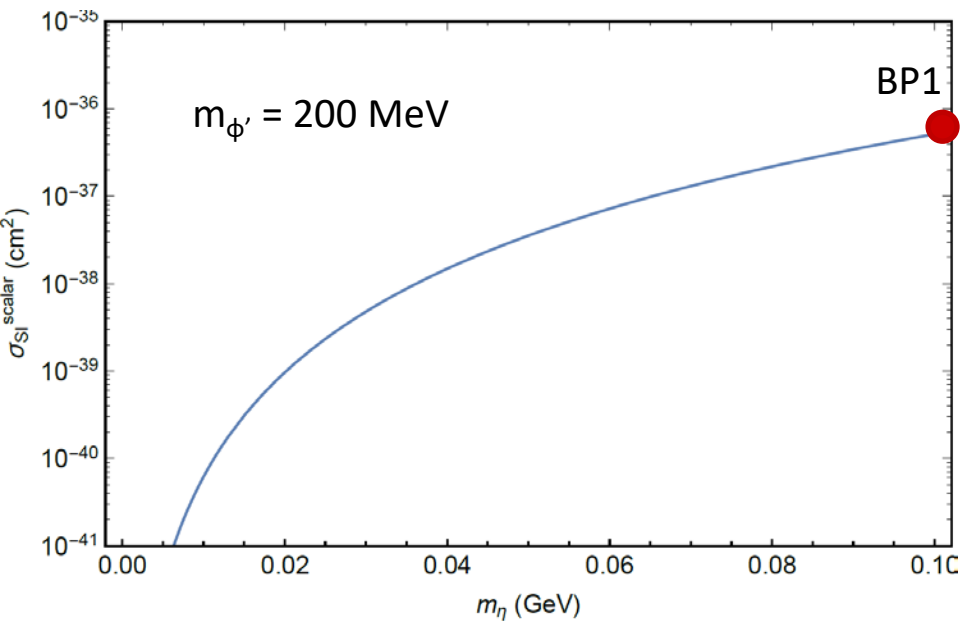
# direct detection

- $\phi'$  mediated  $\rightarrow$  SI, velocity-independent, elastic, isospin-invariant
- $A'$  mediated  $\rightarrow$  SI, velocity-independent, **inelastic**, **isospin-violating** (IVDM)
  - opposite coupling to u and d (thus to p and n)
- mediator mass can be of the same order as momentum transfer
  - **not a contact interaction**,  $d\sigma/dE_R$  suppressed by  $[1+(2m_A E_R/m_{\phi',A'}^2)]^{-2}$
- current strategies for direct detection of low-mass DM
  - **low threshold**
  - **Migdal effect** – nuclear recoil results in electrons being kicked out
  - DM upscattered by **cosmic ray** interactions
    - boosted relativistic DM well above threshold (Bringmann, Pospelov -1810.10543; Dent, Dutta, Newstead, Shoemaker – 1907.03782)
  - **DM-electron scattering** (one-loop suppressed... not constraining)
- **future experiments upcoming**



# direct detection

- current constraints (contact interaction, isospin-invariant,  $\delta=0$ )
  - CRESST III  $\rightarrow \sigma_{SI} \sim 10^{-35} \text{ cm}^2$  at  $m_\eta = 200 \text{ MeV}$
  - CDEX-1B  $\rightarrow \sigma_{SI} \sim 10^{-32-34} \text{ cm}^2$  at  $m_\eta = 50-180 \text{ MeV}$
  - XENON1T  $\rightarrow \sigma_{SI} \sim 10^{-29-30} \text{ cm}^2$  over full mass range, up-scattering
    - $\sigma_{SI} \sim 10^{-34} \text{ cm}^2$  at  $m_\eta = 100 \text{ MeV}$  (Migdal effect, 1907.12771)
- benchmark models satisfy all bounds





# constraints

Batell, Freitas, Ismail, McKeen (1712.10022); Bauer, Foldenauer, Jaeckel (1803.05466)

- but lots of **constraints** on  $A'$ ,  $\phi'$  coupling to SM fermions
  - couplings fixed in terms of masses and  $V$
- main **differences** between our scenario and others
  - **no coupling to  $\bar{\nu}_L \nu_L$**  ( $\nu_R/\nu_A$  mixing taken small),
    - suppresses **v experiment and astrophysical cooling constraints** when  $\nu_A$  involved
  - **no direct coupling to e**
    - some  **$e^+e^-$  collider constraints** suppressed at one-loop
  - **chiral coupling** of  $A'$  to SM fermions
    - even at weak coupling ( $g_{T3R} \rightarrow 0$ ), longitudinal mode (Goldstone) does not decouple
- **$g_\mu-2$  corrections** from  $\phi'$  (positive) and  $A'$  (negative) running in loop
  - corrections can be **fine-tuned** against each other or **heavy new physics**
  - even weakly coupled  $A'$  contributes to  $g-2$  via **massless Goldstone mode**



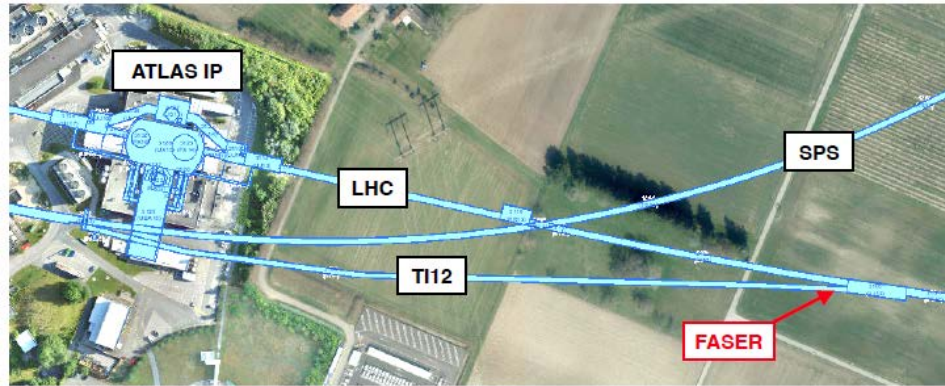
# constraints

- main constraints
  - $e^+e^- \rightarrow 4\mu$  (BaBar)
  - fixed target/beam dump exps.:  $A', \phi' \rightarrow \gamma\gamma, e^+e^-$  (E137, E774, E141, Orsay, etc.)
  - solar/SN/Glob. Cluster cooling constraints (production of  $A', \phi'$ )
  - fifth force constraints  $\rightarrow$  constrains light mediators
    - when  $A'$  light, longitudinal mode still coupled like a Goldstone boson
  - BBN/CMB  $\rightarrow$  need  $\Delta N_{\text{eff}}$  small ( $\lesssim 0.5$ )
- we'll take  $V = 10 \text{ GeV}$ , and will find restrictions on  $m_{\phi'}$  and  $m_{A'}$ 
  - not much dependence on dark matter mass
  - take neutrino mixing angle small
  - take  $m_{\nu_s} > 10 \text{ MeV}$  also (doesn't affect  $N_{\text{eff}}$ )
  - take  $m_{\eta} > 40 \text{ MeV}$  (freezeout before BBN, doesn't affect  $N_{\text{eff}}$ )



# future experiments

- **laboratory production** of  $A'$ ,  $\phi'$ 
  - FASER, SHiP, SeaQuest, MATHUSLA
  - DUNE, COHERENT, CCM
  - LDMX
  - NA64 $\mu$
  - COMPASS
  
- dark matter **direct detection**
  - LZ, CDEX
  - SuperCDMS, CRESST, NEWS-G
  - LBECA, SENSEI



FASER (1811.12522)



LZ





# laboratory production

- produce  $A'$ ,  $\phi'$  from a p beam
    - FASER, SHiP, SeaQuest, MATHUSLA
      - detect visible decay at distant detector
    - DUNE, COHERENT, CCM
      - decay to DM, which scatters in distant detector
  - NA64  $\mu \rightarrow A'$ ,  $\phi'$  produced by  $\mu$ -beam, and decay invisibly
  - LDMX  $\rightarrow$  electrons scattering off nuclei, producing MET
  - reanalysis of COMPASS data
    - muon beam strikes a thin target
- plots in progress!



# direct detection

- need to account for **DM mass splitting** for **A'-mediated** case
  - not such a big effect for CR-upscattering, since already above threshold
  - more important for **low-threshold** searches
- **new solid-state detectors** can lower thresholds further
- when detectors are within reach of models, need to address **recoil spectrum** (non-contact interaction) and **isospin violation**
  
- analysis in progress!



# electron coupling

- what if the **electron** is coupled to  $U(1)_{T3R}$ , not muon?
- basic problem is  $A'$ 
  - if **low-mass**, ruled out by constraints on  $N_{\text{eff}}$  (preliminary)
  - if **higher-mass**, **decays early**, but ruled out by **atomic parity violation** experiments
  - **right-handed coupling violates parity**
- can potentially **fine-tune this away**, either by **cancelling against new physics**, or **scaling up  $V$**
- other constraints modified by direct coupling to  $e$
- **DM-electron scattering** becomes more important
- **future work to expand on this....**